

# Introduction to Bubble Chamber: Ops Training

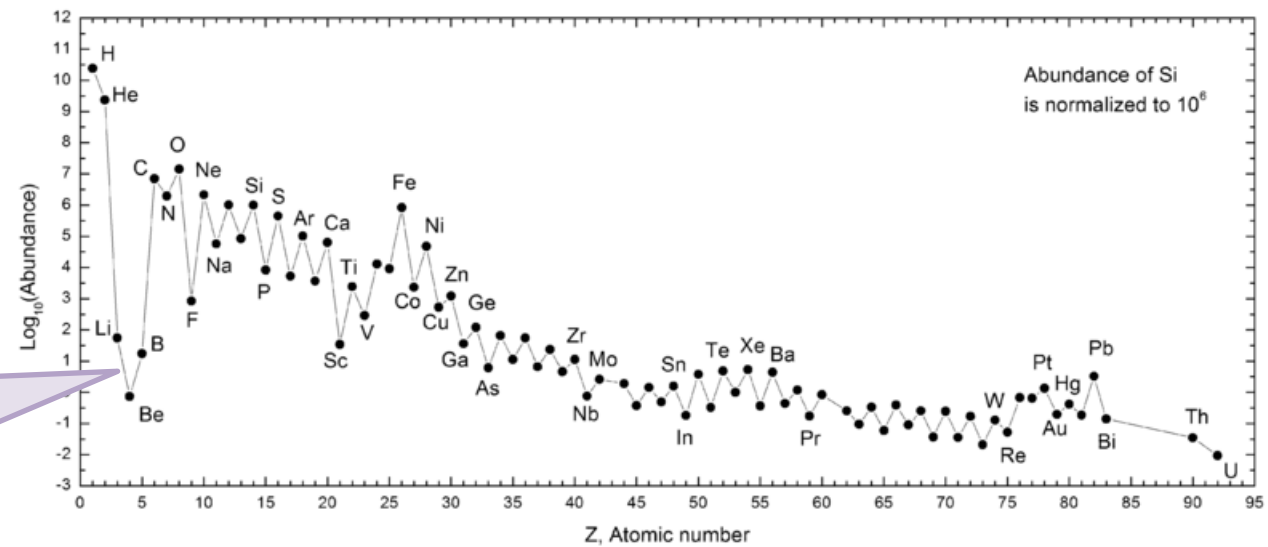
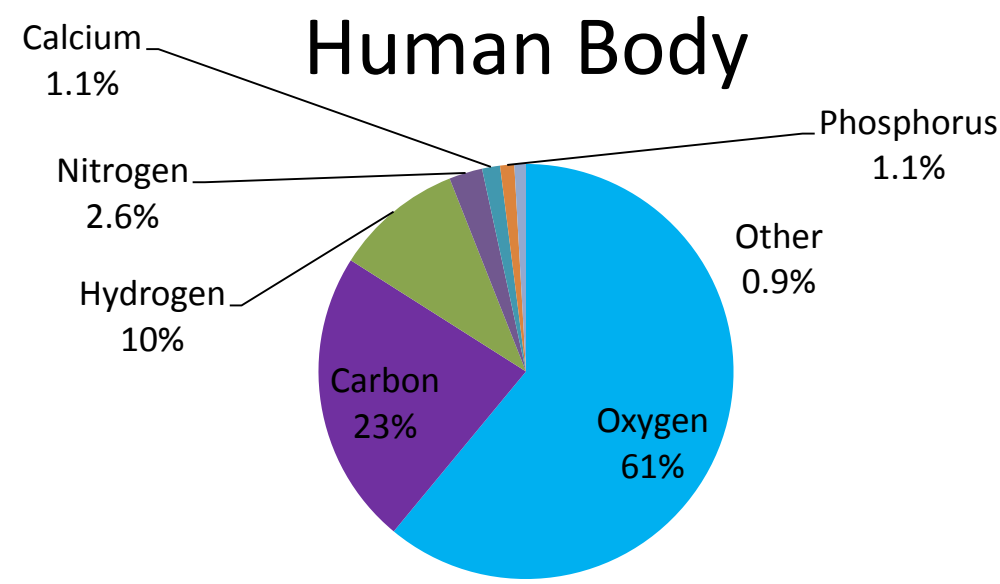
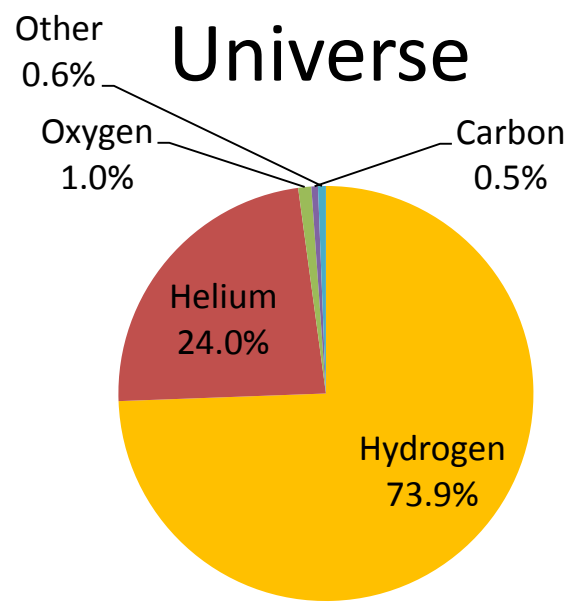
September 9, 2015

[https://wiki.jlab.org/ciswiki/index.php/Bubble Chamber](https://wiki.jlab.org/ciswiki/index.php/Bubble_Chamber)

# OUTLINE

- Nucleosynthesis and  $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$  Reaction
- Time Reversal Reaction:  $^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$
- Bubble Chamber
- Electron Beam Requirements
- Bremsstrahlung Beam
- Penfold-Leiss Cross Section Unfolding
- Bubble Chamber Test Plans
- Bubble Chamber Safety

# RELATIVE ABUNDANCE OF ELEMENTS BY WEIGHT



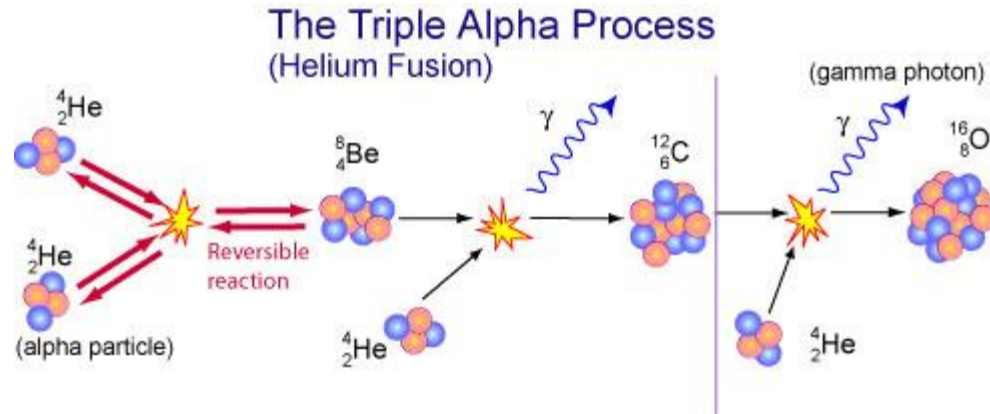
# NUCLEOSYNTHESIS

- **Big Bang Nucleosynthesis:** quark–gluon plasma  $\rightarrow$  p, n, He
- **Stellar Nucleosynthesis:** H burning, He burning, NCO cycle
- **Supernovae Nucleosynthesis:** Si burning
- **Cosmic Ray Spallation**

<b>H</b> B																	<b>He</b> B
<b>Li</b> C	<b>Be</b> C																
<b>Na</b> L	<b>Mg</b> L																
<b>K</b> L	<b>Ca</b> L	<b>Sc</b> L	<b>Ti</b> \$ L	<b>V</b> \$ L	<b>Cr</b> L	<b>Mn</b> L	<b>Fe</b> \$ L	<b>Co</b> \$	<b>Ni</b> \$	<b>Cu</b> L	<b>Zn</b> L	<b>Ga</b> \$	<b>Ge</b> \$	<b>As</b> L	<b>Se</b> \$	<b>Br</b> \$	<b>Kr</b> \$
<b>Rb</b> \$	<b>Sr</b> L	<b>Y</b> L	<b>Zr</b> L	<b>Nb</b> L	<b>Mo</b> \$ L	<b>Tc</b> L	<b>Ru</b> \$ L	<b>Rh</b> \$	<b>Pd</b> \$ L	<b>Ag</b> \$ L	<b>Cd</b> \$ L	<b>In</b> \$ L	<b>Sn</b> \$ L	<b>Sb</b> \$	<b>Te</b> \$	<b>I</b> \$	<b>Xe</b> \$
<b>Cs</b> \$	<b>Ba</b> L		<b>Hf</b> \$ L	<b>Ta</b> \$ L	<b>W</b> \$ L	<b>Re</b> \$	<b>Os</b> \$	<b>Ir</b> \$	<b>Pt</b> \$	<b>Au</b> \$	<b>Hg</b> \$ L	<b>Tl</b> \$ L	<b>Pb</b> \$	<b>Bi</b> \$	<b>Po</b> \$	<b>At</b> \$	<b>Rn</b> \$
<b>Fr</b> \$	<b>Ra</b> \$																
			<b>La</b> L	<b>Ce</b> L	<b>Pr</b> \$ L	<b>Nd</b> \$ L	<b>Pm</b> \$ L	<b>Sm</b> \$ L	<b>Eu</b> \$	<b>Gd</b> \$	<b>Tb</b> \$	<b>Dy</b> \$	<b>Ho</b> \$	<b>Er</b> \$	<b>Tm</b> \$	<b>Yb</b> \$ L	<b>Lu</b> \$
			<b>Ac</b> \$	<b>Th</b> \$	<b>Pa</b> \$	<b>U</b> \$	<b>Np</b> \$	<b>Pu</b> \$	<b>Am</b> M	<b>Cm</b> M	<b>Bk</b> M	<b>Cf</b> M	<b>Es</b> M	<b>Fm</b> M	<b>Md</b> M	<b>No</b> M	<b>Lr</b> M

# NUCLEOSYNTHESIS AND $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$

Stellar Helium  
burning



➤ The *holy grail* of nuclear astrophysics:

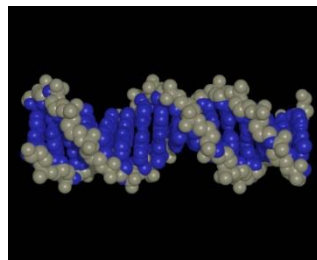
# Periodic Table of the Elements

1	H	2	He																	18	Ar	19	K	20	Ca	21	Sc	22	Ti	23	V	24	Cr	25	Mn	26	Fe	27	Co	28	Ni	29	Cu	30	Zn	31	Ga	32	Ge	33	As	34	Se	35	Br	36	Kr																																
2	Li	Be																	10	Ne	11	Na	12	Mg	13	Al	14	Si	15	P	16	S	17	Cl	18	Ar																																																					
3	Na	Mg	3	Al	4	Si	5	P	6	S	7	Cl	8	Ar	9	K	10	Ca	11	Sc	12	Ti	13	V	14	Cr	15	Mn	16	Fe	17	Co	18	Ni	19	Cu	20	Zn	21	Ga	22	Ge	23	As	24	Se	25	Br	26	Kr																																							
4	K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr	1	H	2	He	3	Li	4	Be	5	B	6	C	7	N	8	O	9	F	10	Ne																																																			
5	Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe	21	Bi	22	Pb	23	Bi	24	Po	25	At	26	Rn	27	Fr	28	Ra	29	Ac	30	Th	31	Pa	32	U	33	Np	34	Pu	35	Am	36	Cm	37	Bk	38	Cf	39	Es	40	Fm	41	Md	42	No	43	Lr																									
6	Cs	Ba	La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu	1	H	2	He	3	Li	4	Be	5	B	6	C	7	N	8	O	9	F	10	Ne	11	Na	12	Mg	13	Al	14	Si	15	P	16	S	17	Cl	18	Ar	19	K	20	Ca	21	Sc	22	Ti	23	V	24	Cr	25	Mn	26	Fe	27	Co	28	Ni	29	Cu	30	Zn	31	Ga	32	Ge	33	As	34	Se	35	Br	36	Kr
7	Fr	Ra	Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No	Lr	1	H	2	He	3	Li	4	Be	5	B	6	C	7	N	8	O	9	F	10	Ne	11	Na	12	Mg	13	Al	14	Si	15	P	16	S	17	Cl	18	Ar	19	K	20	Ca	21	Sc	22	Ti	23	V	24	Cr	25	Mn	26	Fe	27	Co	28	Ni	29	Cu	30	Zn	31	Ga	32	Ge	33	As	34	Se	35	Br	36	Kr

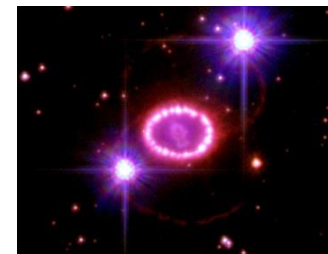
\* Lanthanide Series

+ Actinide Series

Affects synthesis of  
most of elements in  
periodic table



Sets  $N(^{12}\text{C})/N(^{16}\text{O})$   
( $\approx 0.4$ ) ratio in universe



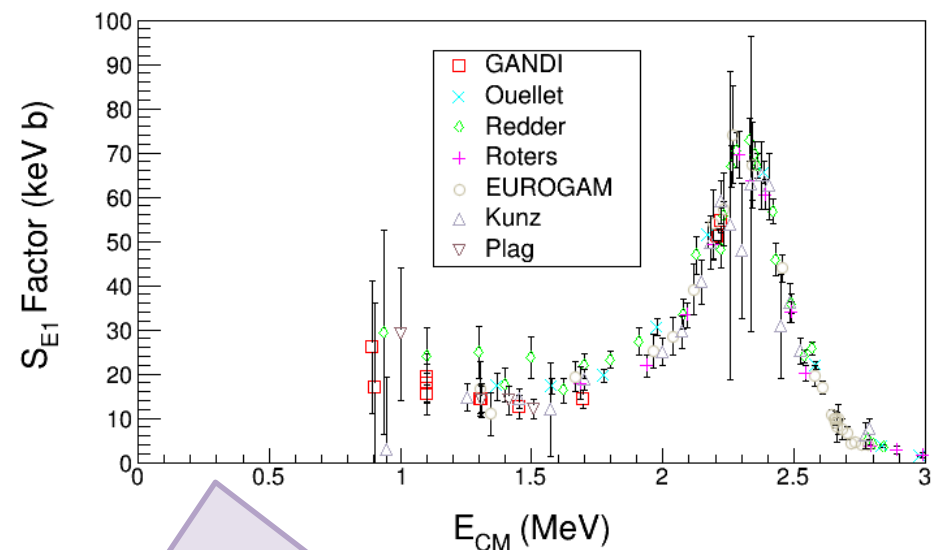
Determines minimum  
mass star requires to  
become supernova

# HEROIC EFFORTS IN SEARCH OF $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$

- Previous cross section measurements:
  - Helium ions on carbon target:  $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$
  - Carbon ions on helium gas:  $^4\text{He}(^{12}\text{C},\gamma)^{16}\text{O}$  or  $^4\text{He}(^{12}\text{C},^{16}\text{O})\gamma$
- Define *S-Factor* to remove both  $1/E$  dependence of nuclear cross sections and Coulomb barrier transmission probability:

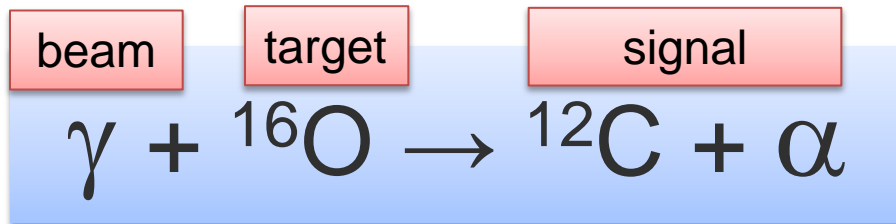
$$S \equiv E_{CM} \sigma(\alpha, \gamma) e^{2\pi\eta}$$

Author	$S_{\text{tot}}(300)$ (keV b)
Hammer (2005)	$162 \pm 39$
Kunz (2001)	$165 \pm 50$



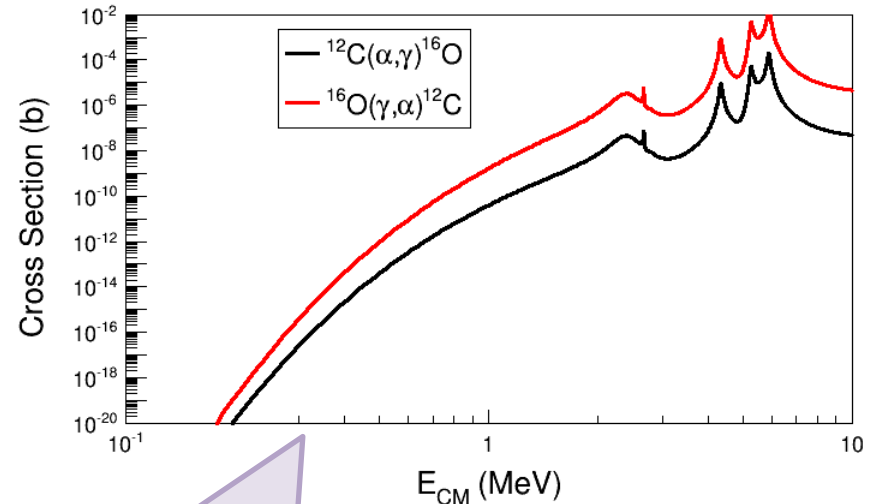
R-matrix Extrapolation to stellar  
helium burning at  $E = 300$  keV

# NEW APPROACH: REVERSAL REACTION + BUBBLE CHAMBER



$$Q = +7.162 \text{ MeV}$$

$$E_{\gamma} \cong E_{CM} + Q$$

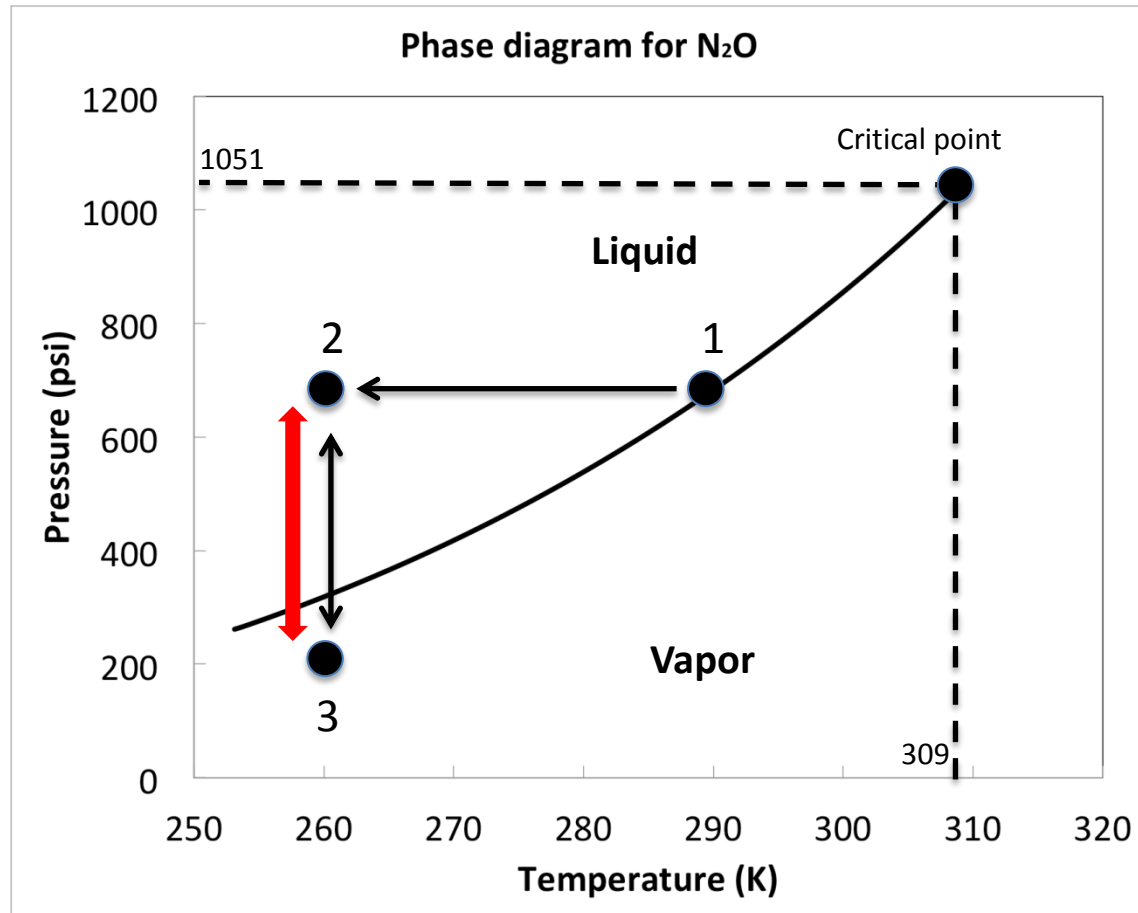


Stellar helium burning at  
 $E = 300 \text{ keV}$ ,  $T = 200 \times 10^6 \text{ K}$

- Extra gain (factor of 100) by measuring time reversal reaction
- Bremsstrahlung at JLab  $\sim 10^9 \text{ } \gamma/\text{s}$  (top 250 keV)
- Target density up to  $10^4$  higher than conventional targets. Number of  $^{16}\text{O}$  nuclei =  $3.5 \times 10^{22}/\text{cm}^2$  (3.0 cm cell)
- Electromagnetic debris (electrons and gammas, or positrons) do NOT trigger nucleation (detector is insensitive to  $\gamma$ -rays by at least 1 part in  $10^{11}$ )

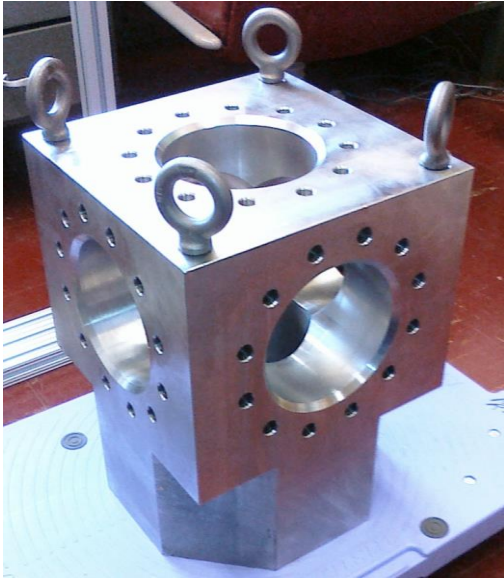
# THE BUBBLE CHAMBER

- 1 Cell is cooled then filled with room temperature gas
- 2 Gas is cooled and condenses into liquid
- 3 Once cell is completely filled with liquid, pressure is reduced creating a superheated liquid
- 3 Nuclear reactions induce bubble nucleation
- 2 High speed camera detects bubble and repressurizes
- 3 System depressurizes and ready for another cycle



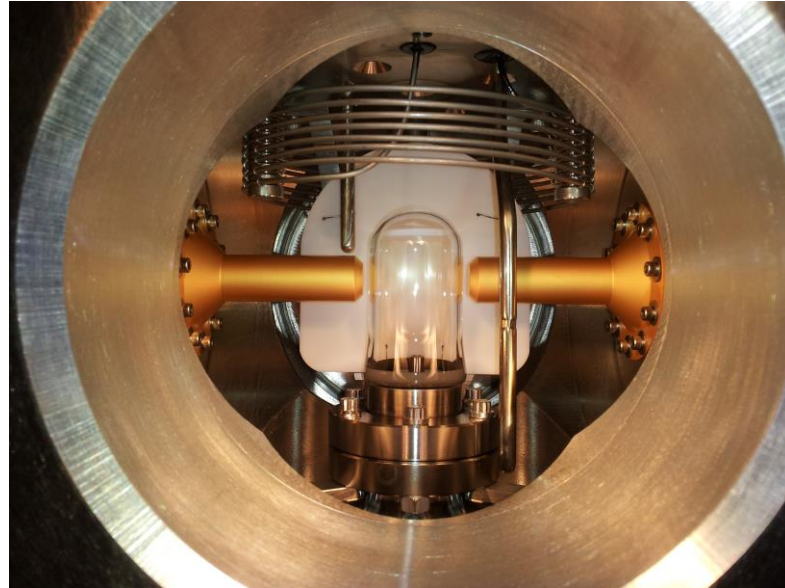
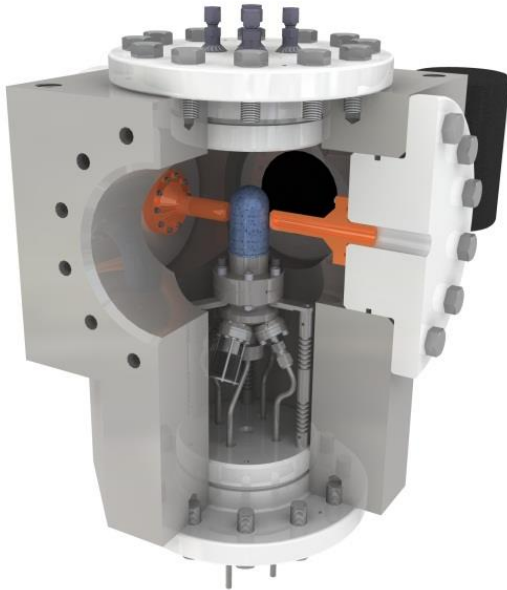


# $\text{N}_2\text{O}$ (LAUGHING GAS) BUBBLE CHAMBER



$T = -10^{\circ}\text{C}$

$P = 50 \text{ atm}$



# USER INTERFACE

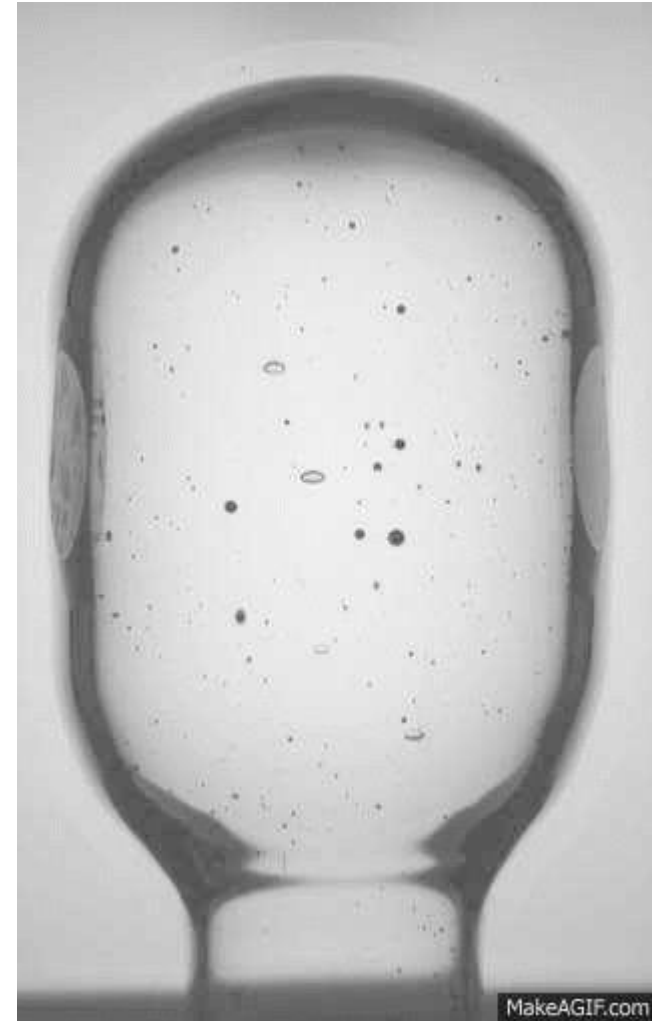
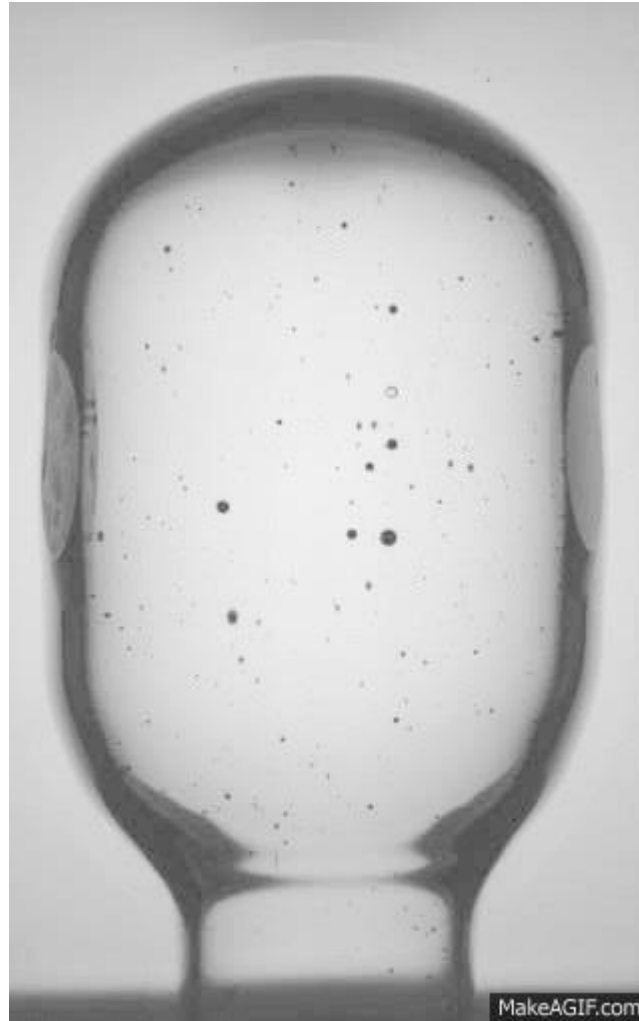


# BUBBLE GROWTH AND QUENCHING

3.0 cm

100 Hz Digital  
Camera  $\Delta t = 10$  ms

$\text{N}_2\text{O}$  Chamber  
with PuC neutron  
source



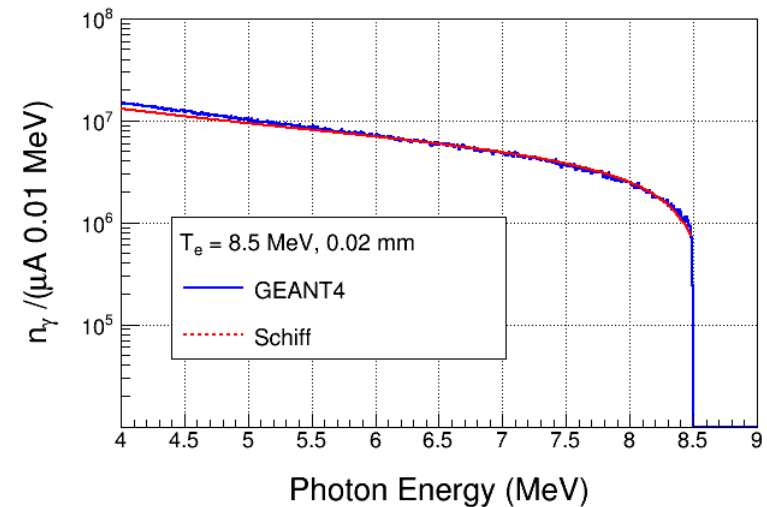
# ELECTRON BEAM REQUIREMENTS

## I. Beam Properties at Radiator:

Beam Kinetic Energy, (MeV)	7.9 – 8.5
Beam Current ( $\mu\text{A}$ )	0.01 – 100
Absolute Beam Energy Uncertainty	<0.1%
Relative Beam Energy Uncertainty	<0.02%
Energy Resolution (Spread), $\sigma_T/T$	<0.06%
Beam Size, $\sigma_{x,y}$ (mm)	1
Polarization	None

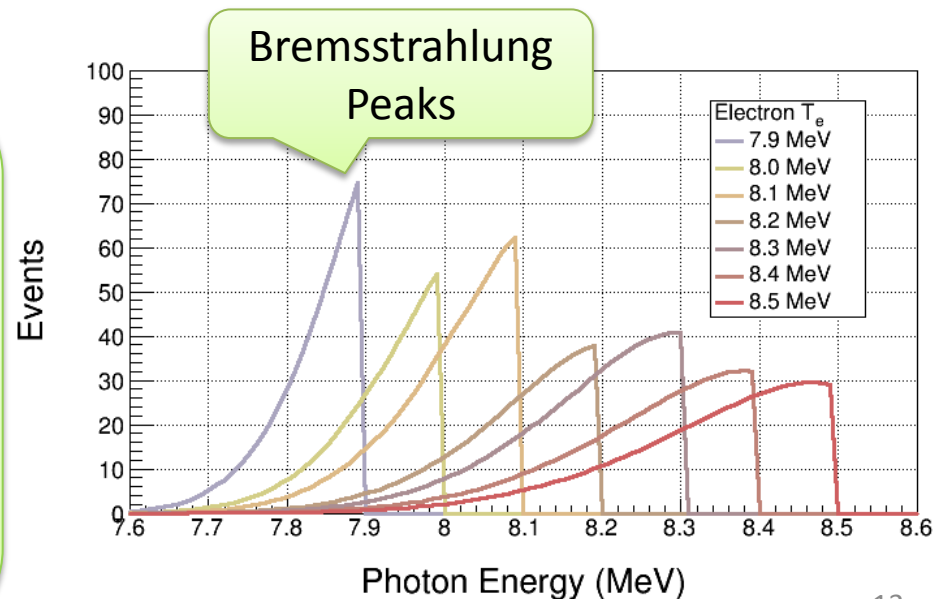
# BREMSSTRAHLUNG BEAM

- Use both GEANT4 and FLUKA to calculate Bremsstrahlung spectra (we will not measure Bremsstrahlung spectra)
- Monte Carlo simulation of Bremsstrahlung at radiotherapy energies is well studied, accuracy:  $\pm 5\%$



$^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$  is ideal case for Bremsstrahlung beam and Penfold–Leiss Unfolding:

- Very steep cross section; only photons near endpoint contribute to yield
- No-structure (resonances)



# PENFOLD-LEISS CROSS SECTION UNFOLDING

- Measure yields at:  $E = E_1, E_2, \dots, E_n$  where,  $E_i - E_{i-1} = \Delta$ ,  $i = 2, n$

$$Y(E_i) = \int_{th}^{E_i} n_\gamma(E_i, k) \sigma(k) dk \approx \sum_{j=1}^i N_\gamma(E_i, \Delta, E_j) \sigma(E_j)$$

Volterra Integral Equation of First Kind

- Solution can be written in two forms:

$$\sigma_i = \frac{1}{N_{ii}} \left[ y_i - \sum_{j=1}^{i-1} (N_{ij} \sigma_j) \right]$$

Method of Quadratures:  
numerical solution of integral  
equation based on replacement  
of integral by finite sum

- Or, Matrix form:

$$\begin{bmatrix} y_1 \\ y_2 \\ \vdots \\ y_n \end{bmatrix} = \begin{bmatrix} N_{\gamma,11} & 0 & \cdots & 0 \\ N_{\gamma,21} & N_{\gamma,22} & \cdots & 0 \\ \vdots & \ddots & \ddots & 0 \\ N_{\gamma,n1} & N_{\gamma,n2} & \cdots & N_{\gamma,nn} \end{bmatrix} \begin{bmatrix} \sigma_1 \\ \sigma_2 \\ \vdots \\ \sigma_n \end{bmatrix}$$

$$[Y] = [N] \bullet [\sigma]$$

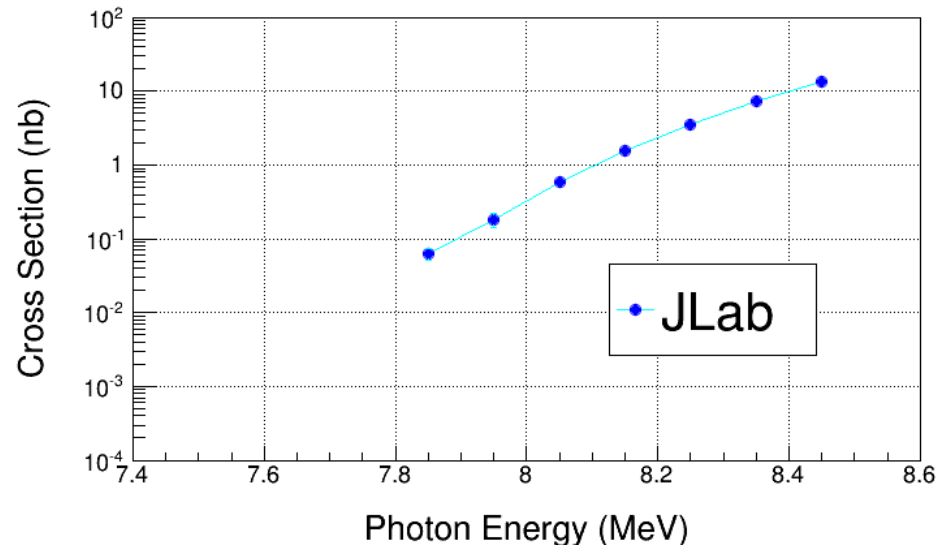
$$[\sigma] = [N]^{-1} \bullet [Y]$$

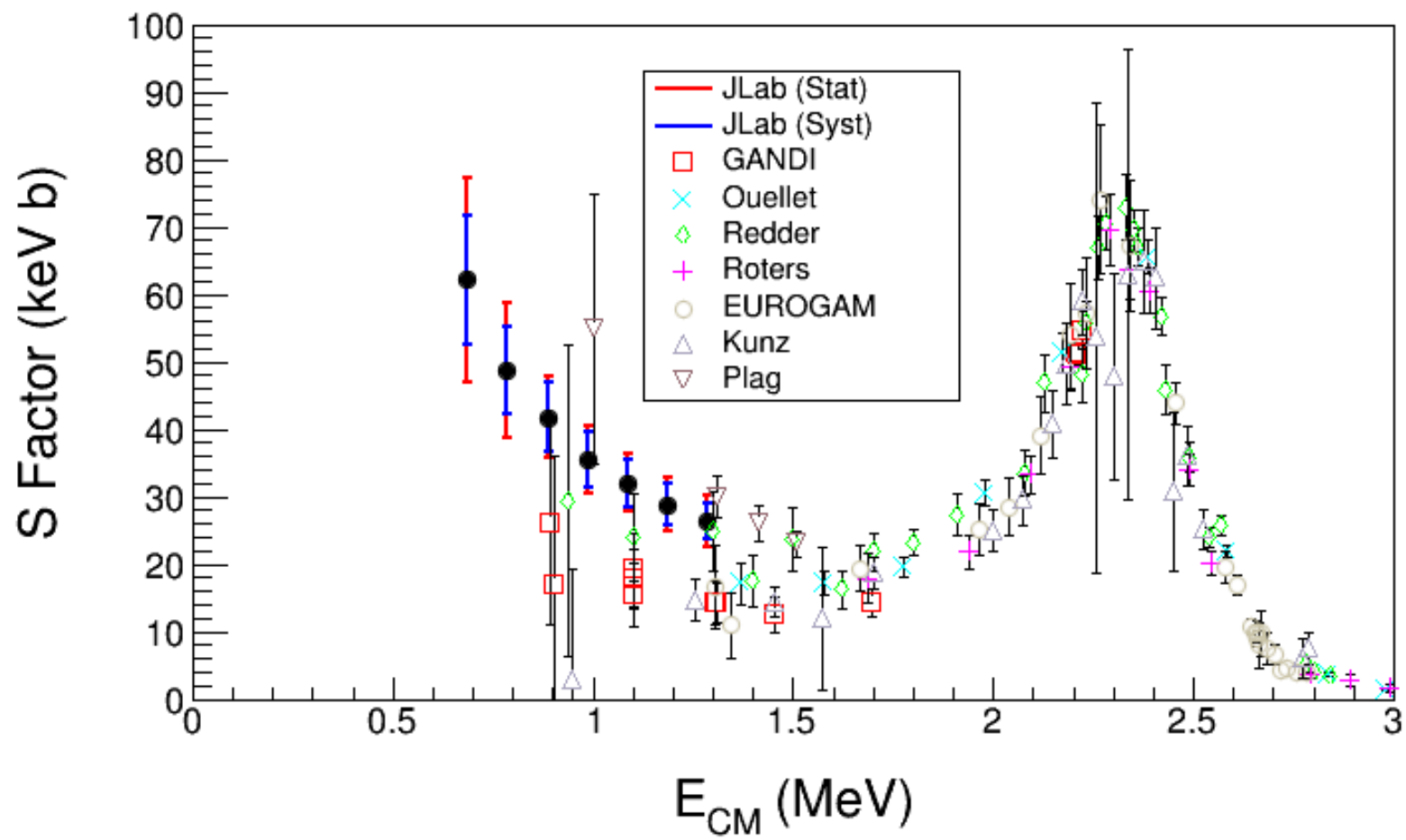


# JLAB PROJECTED $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ S-Factor

- Statistical Error: dominated by background subtraction from  $^{18}\text{O}(\gamma,\alpha)^{14}\text{C}$  (depletion = 5,000)

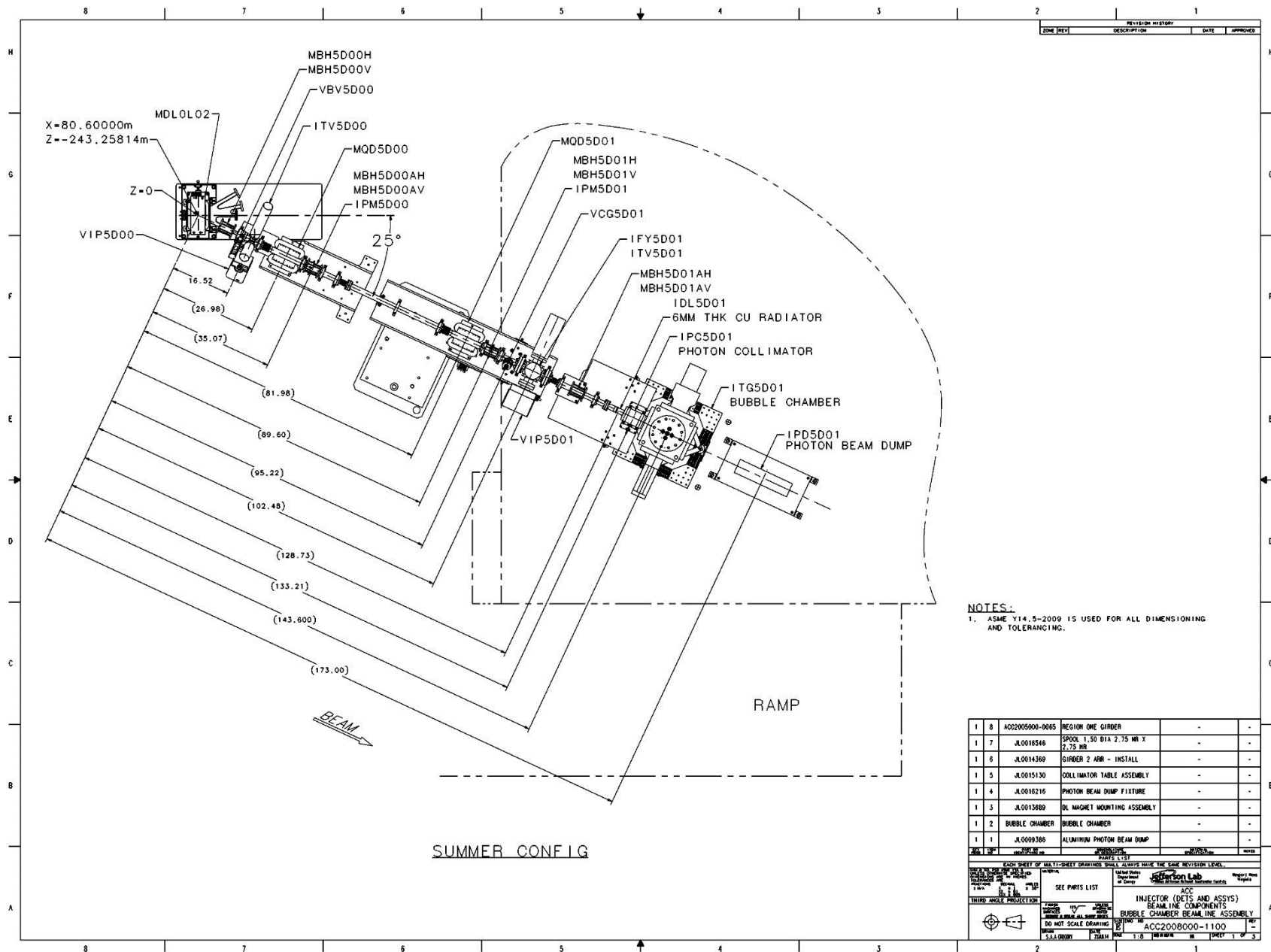
Electron Beam K. E.	Gamma Energy (MeV)	$E_{CM}$ (MeV)	Cross Section (nb)	$S_{tot}$ Factor (keV b)	Stat Error (%)	Sys Error (Total, %)
7.9	7.85	0.69	0.046	62.2	24.5	15.3
8.0	7.95	0.79	0.185	48.7	20.7	13.5
8.1	8.05	0.89	0.58	41.8	14.7	12.2
8.2	8.15	0.99	1.53	35.5	13.8	11.4
8.3	8.25	1.09	3.49	32.0	13.3	10.7
8.4	8.35	1.19	7.2	28.8	13.8	10.5
8.5	8.45	1.29	13.6	26.3	14.8	10.1







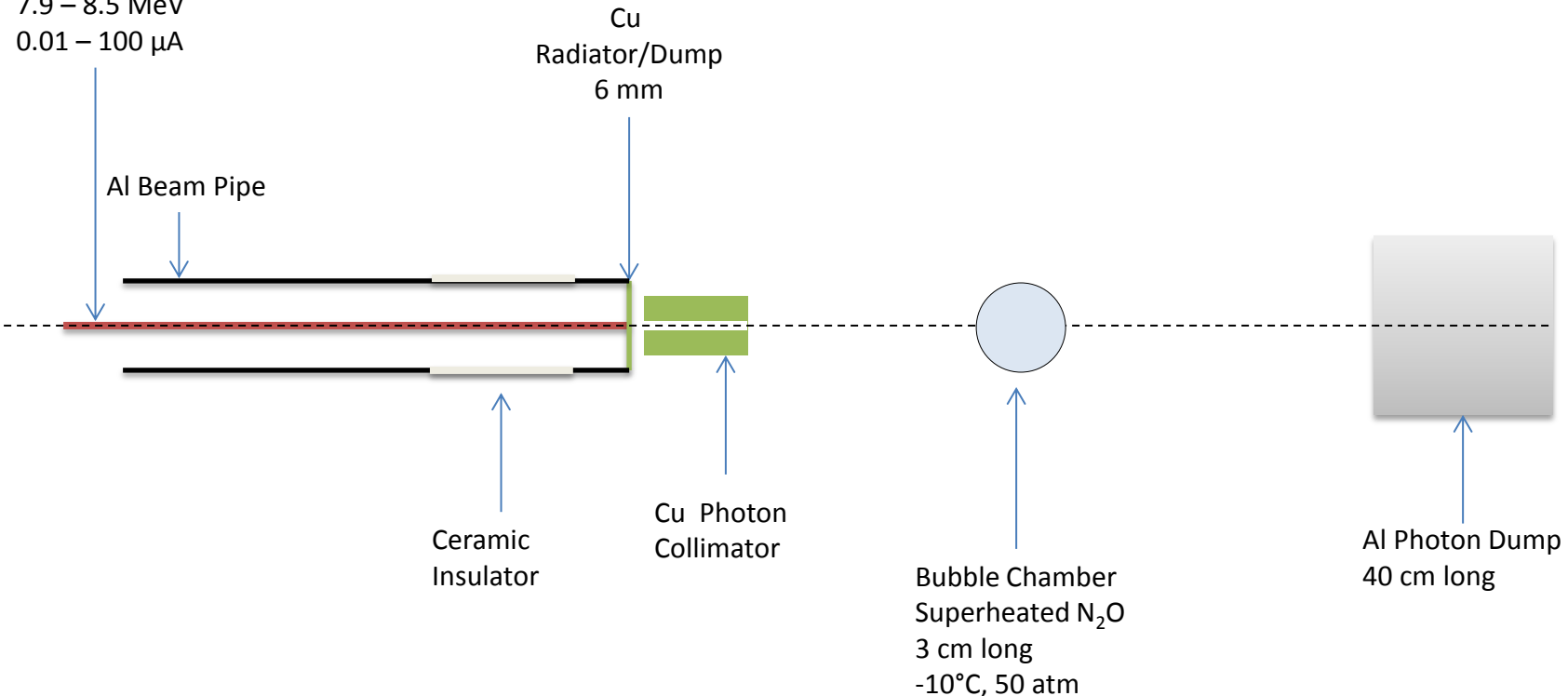
# TEST BEAMLINE

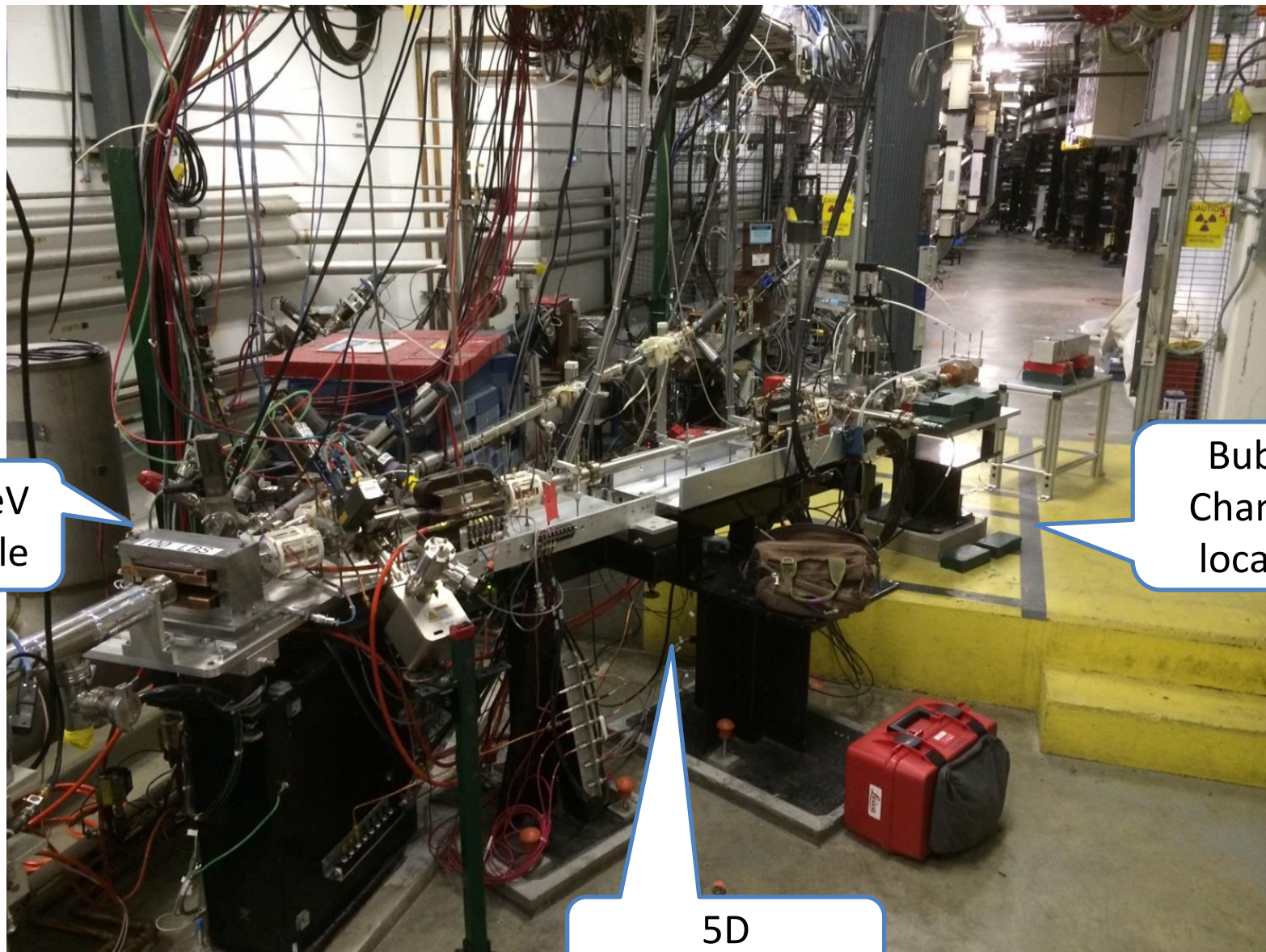


# SCHEMATICS OF TEST BEAMLINE

- Power deposited in radiator (100  $\mu$ A and 8.5 MeV) :
  - 6 mm: Energy loss = 8.5 MeV, P = 850 W
- Pure Copper and Aluminum (high neutron threshold):
  - $^{63}\text{C}(\gamma, n)$  threshold = 10.86 MeV
  - $^{27}\text{Al}(\gamma, n)$  threshold = 13.06 MeV

Electron K.E.  
7.9 – 8.5 MeV  
0.01 – 100  $\mu$ A





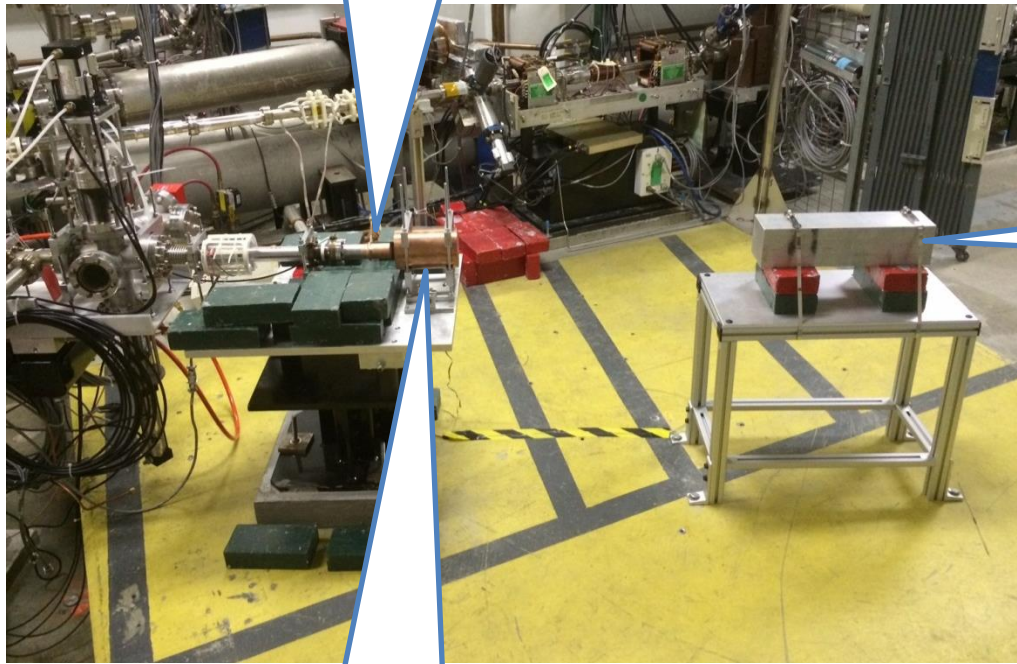
5 MeV  
Dipole

Bubble  
Chamber  
location

5D  
Spectrometer

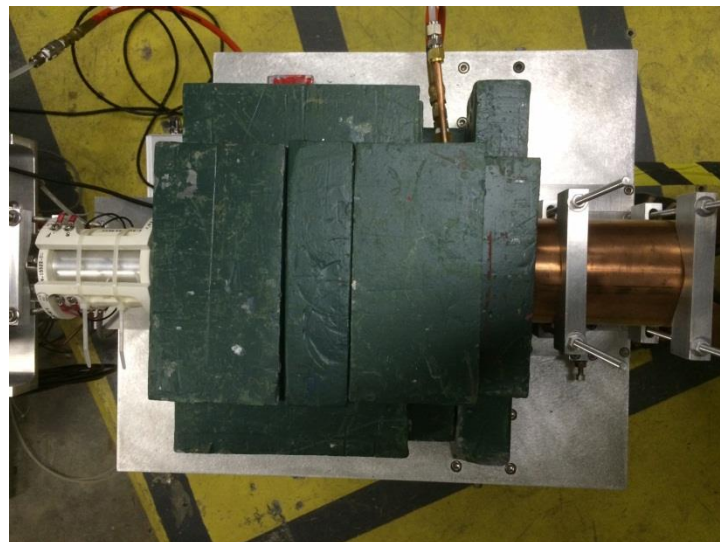


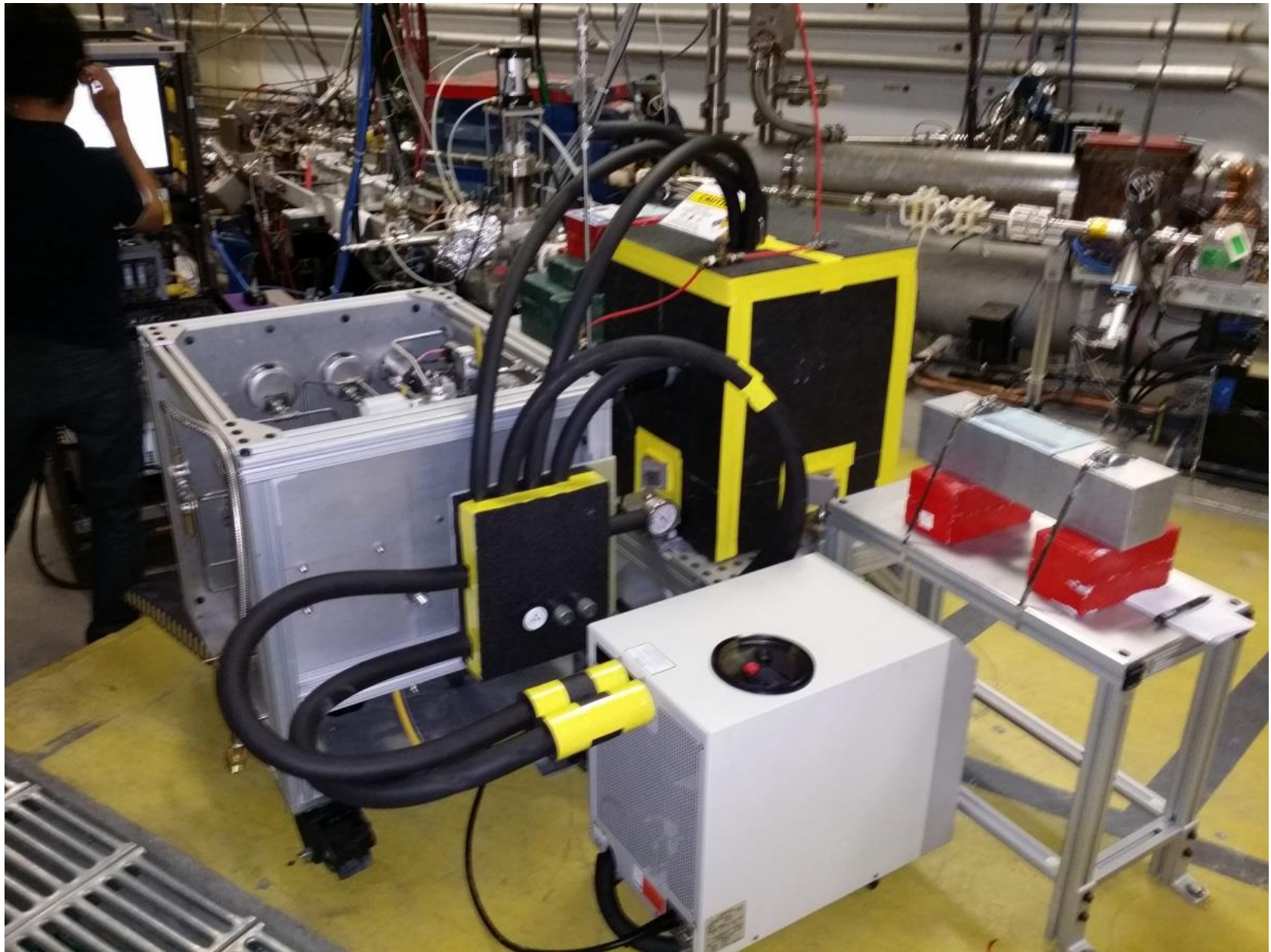
Cu Electron  
Radiator/Dump



Al Photon  
Dump

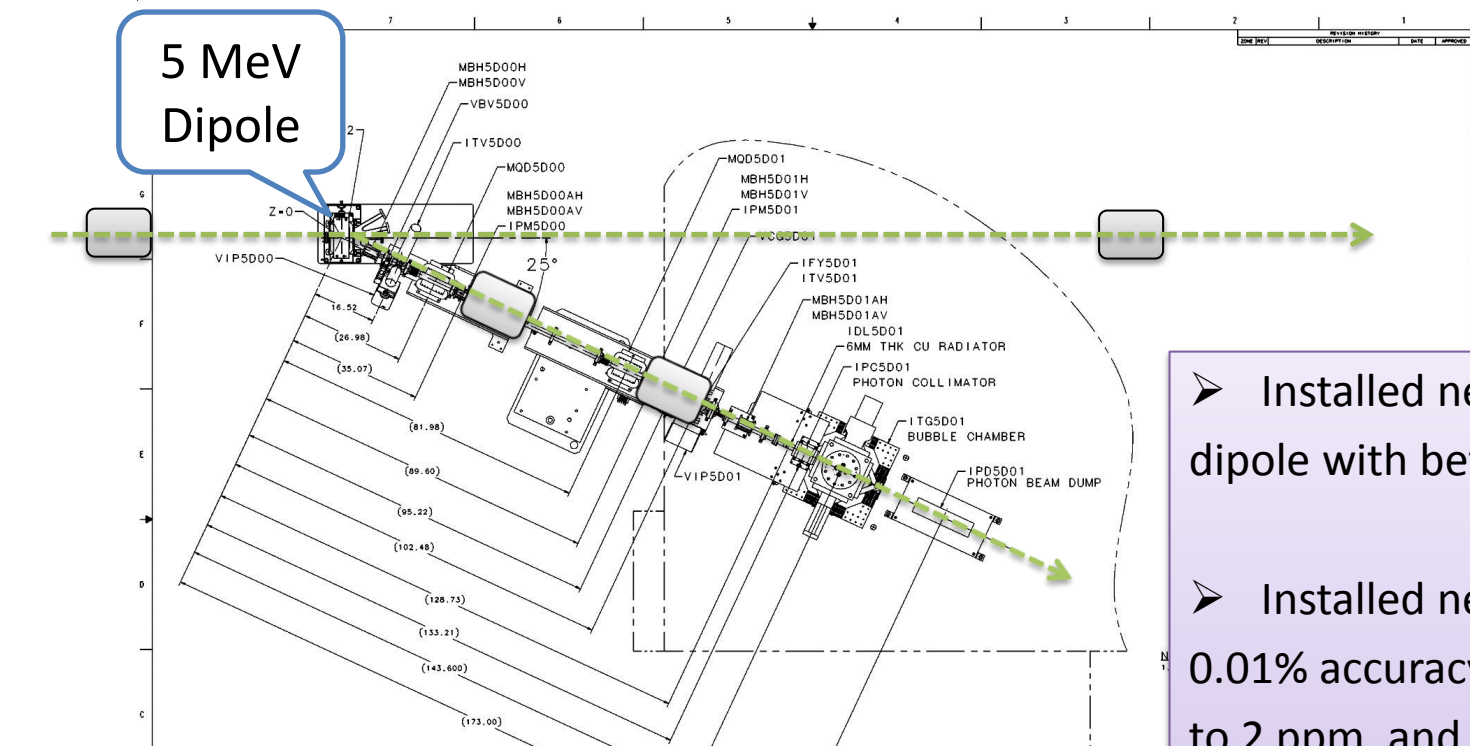
Cu Photon  
Collimator







# MEASURING ABSOLUTE BEAM ENERGY



 Beam Position Monitor (BPM)

Electron Beam Momentum

$$p = \frac{\int B dl}{\theta}$$

- Installed new higher field dipole with better uniformity
- Installed new Hall probe: 0.01% accuracy, resolution to 2 ppm, and a temperature stability of 10 ppm/°C
- Still need to shield Earth's and other stray magnetic fields

# TEST BEAMLINE COMMISSIONING

- Beamline was ready since Fall 2014
- Approved to run 10  $\mu\text{A}$  CW and total energy of 10 MeV
- Completed hot checkout and beam checkout
- Beam Studies completed so far:
  - I. Delivered 10.0  $\mu\text{A}$  and 9.65 MeV (kinetic) for 5 hours in August 2015
  - II. Measured beam momentum at different  $\frac{1}{4}$  cryo-unit settings
  - III. Measured beam charge at different beam currents
- Re-doing realistic thermal analysis to run at 100  $\mu\text{A}$

# BUBBLE CHAMBER TEST PLANS

1. Fill with natural  $\text{N}_2\text{O}$  – test bubble chamber systems operation
  2. Study Chamber with beam (1:00 – 11:00 pm, Sept 10 – 18)
  3. Background measurements
  4. Fill with  $\text{C}_2\text{F}_6$  – test bubble chamber systems operation
  5. With beam (planned in Oct 16 – 22, 2015)
    - I. Measure  $^{19}\text{F}(\gamma, \alpha)^{15}\text{N}$  ( $Q = +4.013$  MeV)
    - II. Compare measured cross section to our HIGS data
- Fluorine is suitable for a first Penfold-Leiss unfolding:
- Only one stable natural isotope ( $^{19}\text{F}$ )
  - Low electron beam kinetic energy (4.6 – 5.2 MeV) – below threshold of any background reaction



# BUBBLE CHAMBER SAFETY REVIEWS

## ➤ Superheated liquid: N<sub>2</sub>O, Nitrous oxide (laughing gas)

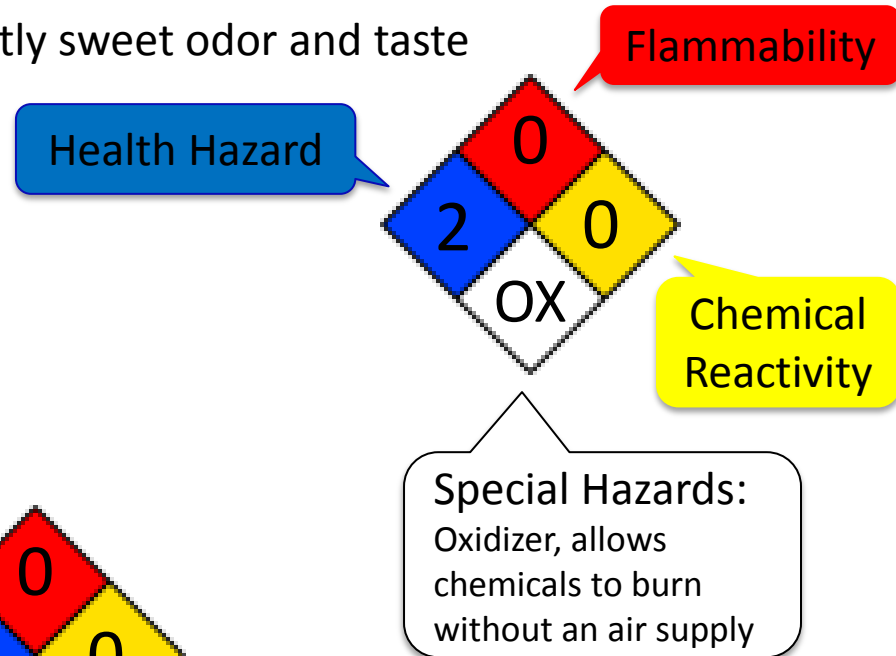
- I. Colorless, non-flammable gas, with slightly sweet odor and taste

## ➤ High pressure system:

- I. Design Authority: Dave Meekins
- II. T = -10°C
- III. P = 50 atm

## ➤ Buffer liquid: Mercury

- I. Closed system
- II. Volume: 150 mL



## ➤ Bubble Chamber Safety Review was on Aug 18, 2015

## ➤ Temporary Operational Safety Procedures (**TOSP**) is approved