

Introduction to Bubble Chamber: Ops Training

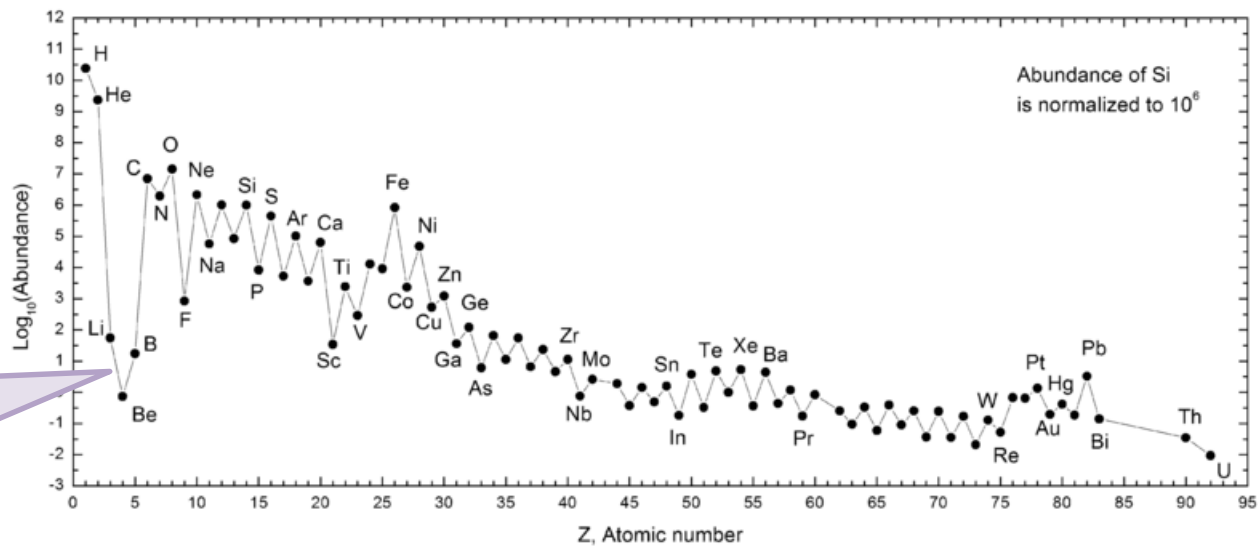
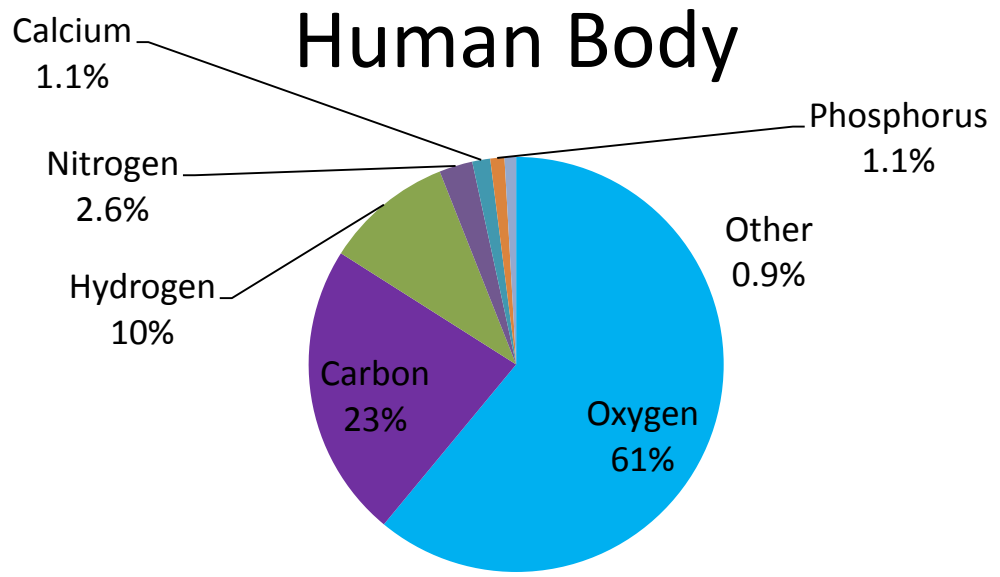
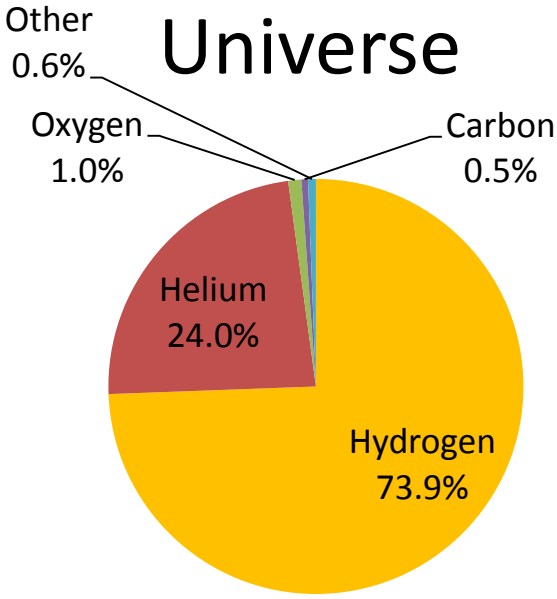
May 9, 2018

[https://wiki.jlab.org/ciswiki/index.php/Bubble Chamber](https://wiki.jlab.org/ciswiki/index.php/Bubble_Chamber)

OUTLINE

- Nucleosynthesis and $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ Reaction
- Time Reversal Reaction: $^{16}\text{O}(\gamma,\alpha)^{12}\text{C}$
- Bubble Chamber
- Engineering Run – May 2018
- Measure $^{19}\text{F}(\gamma,\alpha)^{15}\text{N}$
- Electron Beam Requirements
- Bubble Chamber Test Plans
- Bubble Chamber Safety

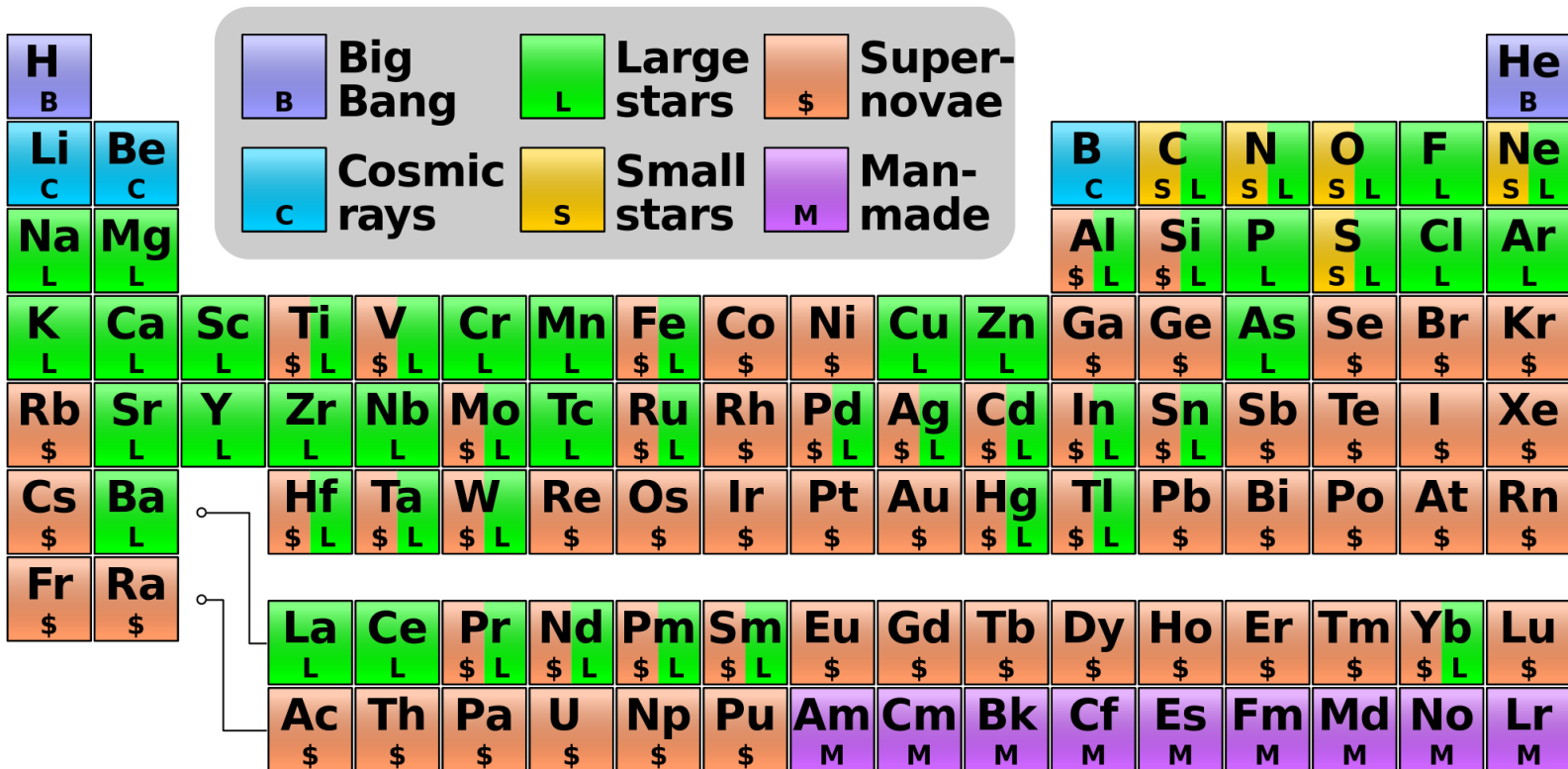
RELATIVE ABUNDANCE OF ELEMENTS BY WEIGHT



This region is bypassed by 3α process

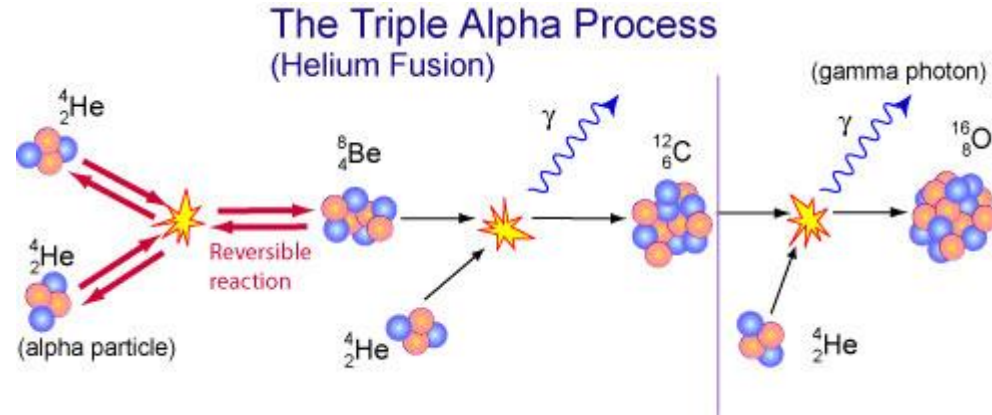
NUCLEOSYNTHESIS

- Big Bang Nucleosynthesis: quark–gluon plasma \rightarrow p, n, He
- Stellar Nucleosynthesis: H burning, He burning, NCO cycle
- Supernovae Nucleosynthesis: Si burning
- Cosmic Ray Spallation



NUCLEOSYNTHESIS AND $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$

Stellar Helium burning

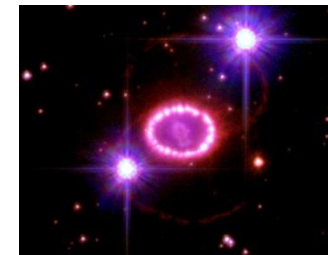
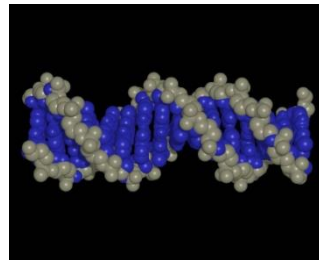


➤ The *holy grail* of nuclear astrophysics:

Periodic Table of the Elements

1	H	He																	Li	Be	B	C	N	O	F	Ne
2	Li	Be																	B	C	N	O	F	Ne		
3	Na	Mg																	Al	Si	P	S	Cl	Ar		
4	K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr								
5	Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe								
6	Cs	Ba	La	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn								
7	Fr	Ra	Ac	Rf	Ha	Sg	Nh	Mt	110	111	112	113														

* Lanthanide Series
+ Actinide Series



Affects synthesis of most of elements in periodic table

Sets $N(^{12}\text{C})/N(^{16}\text{O})$ (≈ 0.4) ratio in universe

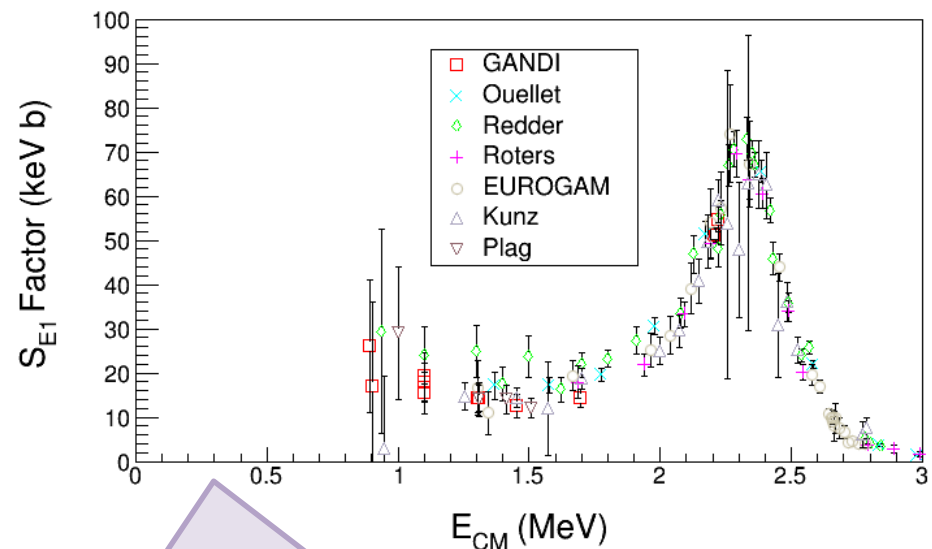
Determines minimum mass star requires to become supernova

HEROIC EFFORTS IN SEARCH OF $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$

- Previous cross section measurements:
 - I. Helium ions on carbon target: $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$
 - II. Carbon ions on helium gas: $^4\text{He}(^{12}\text{C},\gamma)^{16}\text{O}$ or $^4\text{He}(^{12}\text{C},^{16}\text{O})\gamma$
- Define *S-Factor* to remove both $1/E$ dependence of nuclear cross sections and Coulomb barrier transmission probability:

$$S \equiv E_{CM} \sigma(\alpha, \gamma) e^{2\pi\eta}$$

Author	$S_{\text{tot}}(300)$ (keV b)
Hammer (2005)	162 ± 39
Kunz (2001)	165 ± 50



NEW APPROACH: REVERSAL REACTION + BUBBLE CHAMBER

beam

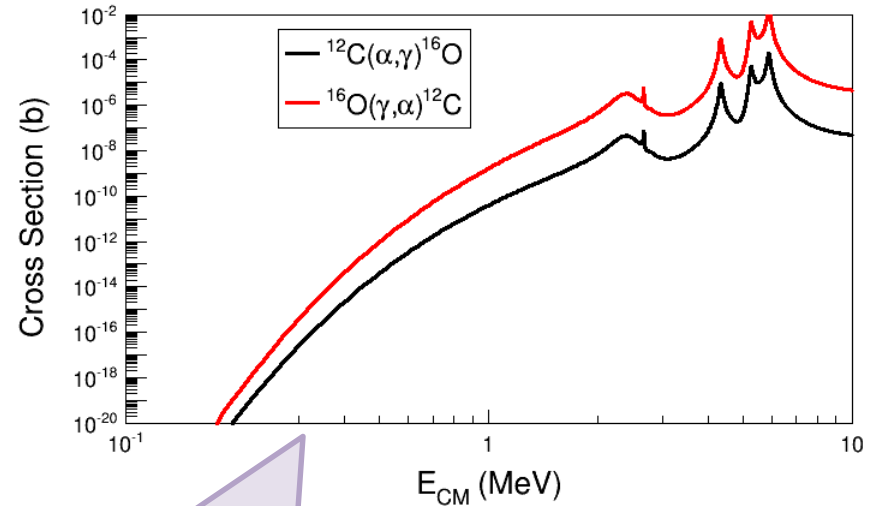
target

signal



$$Q = +7.162 \text{ MeV}$$

$$E_{\gamma} \cong E_{CM} + Q$$

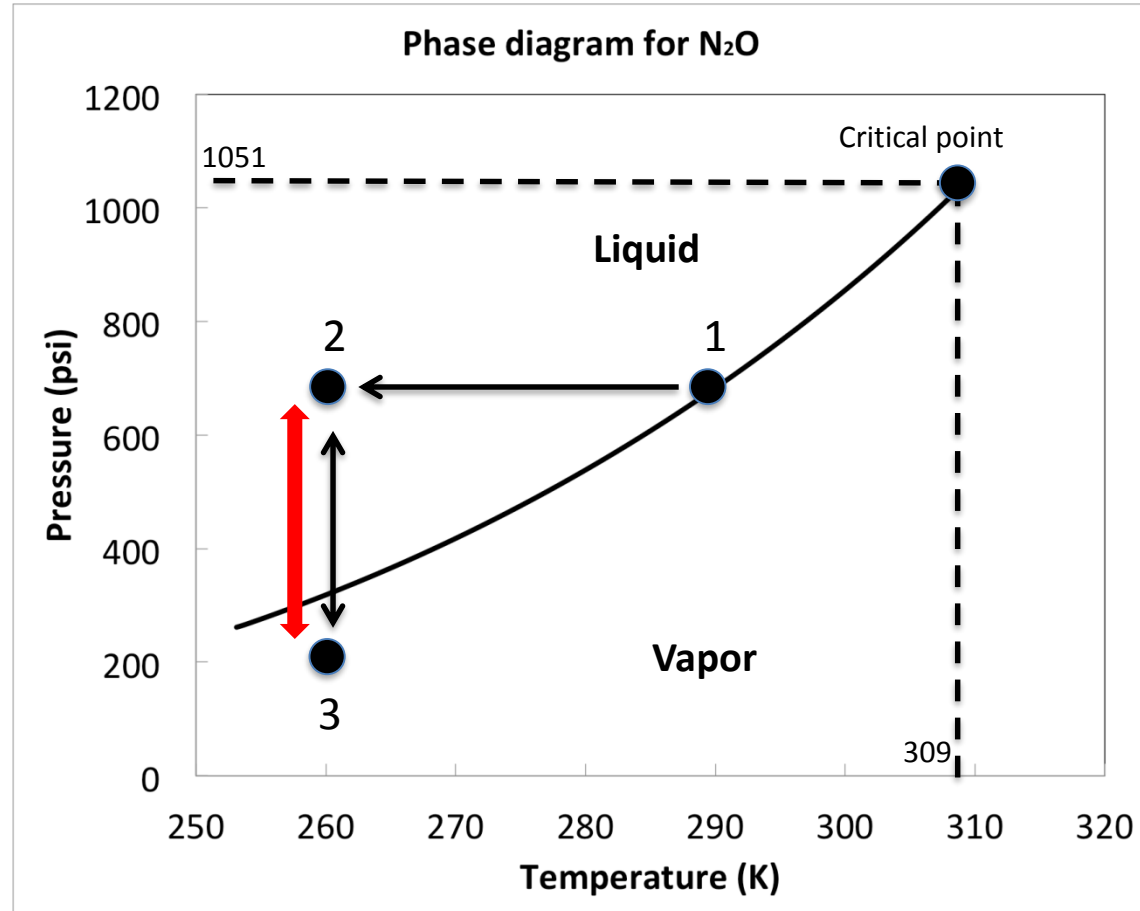


Stellar helium burning at
E = 300 keV, T = 200 × 10⁶ K

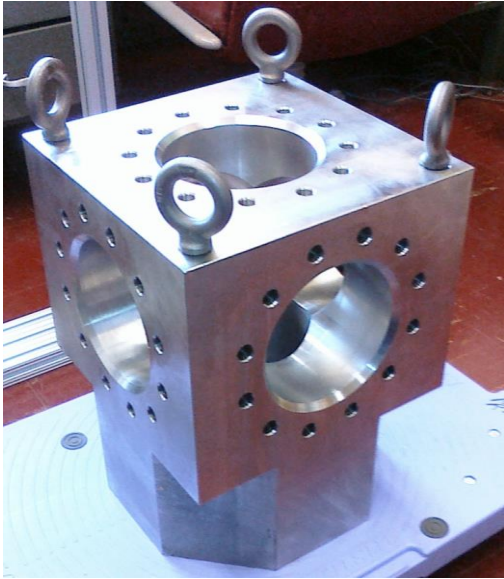
- Extra gain (factor of 100) by measuring time reversal reaction
- Bremsstrahlung at JLab ~ 10⁹ γ/s (top 250 keV)
- Target density up to 10⁴ higher than conventional targets. Number of ¹⁶O nuclei = 3.5 × 10²²/cm² (3.0 cm cell)
- Electromagnetic debris (electrons and gammas, or positrons) do NOT trigger nucleation (detector is insensitive to γ-rays by at least 1 part in 10¹¹)

THE BUBBLE CHAMBER

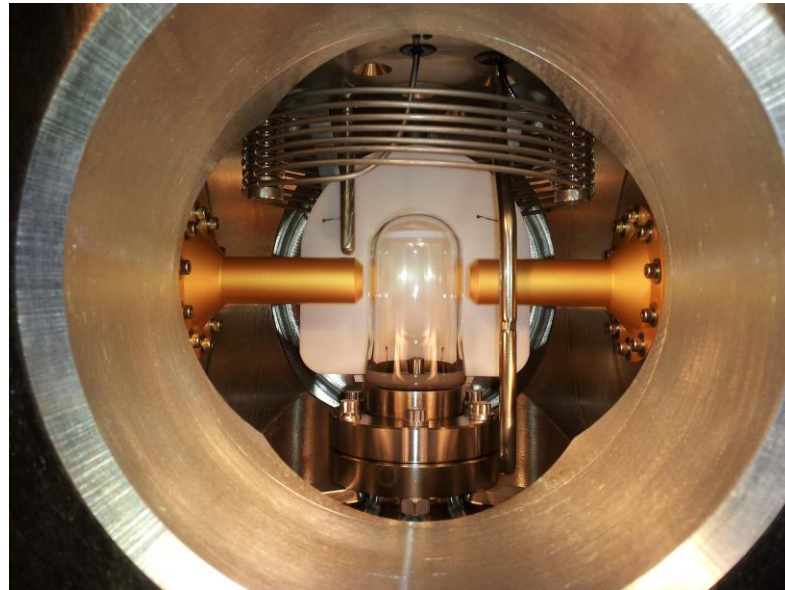
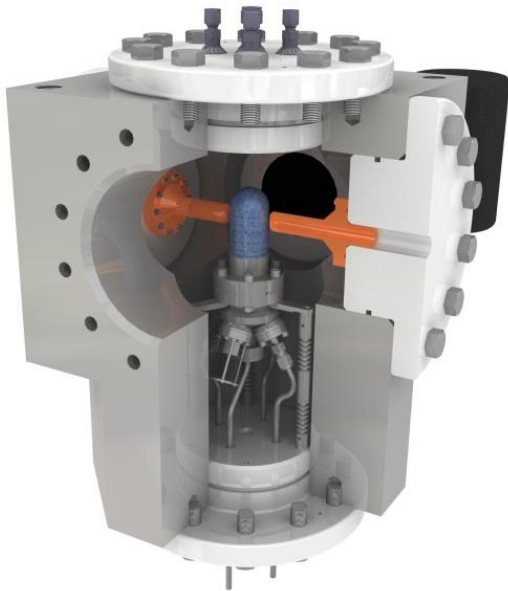
- 1 Cell is cooled then filled with room temperature gas
- 2 Gas is cooled and condenses into liquid
- 3 Once cell is completely filled with liquid, pressure is reduced creating a superheated liquid
- 3 Nuclear reactions induce bubble nucleation
- 2 High speed camera detects bubble and repressurizes
- 3 System depressurizes and ready for another cycle



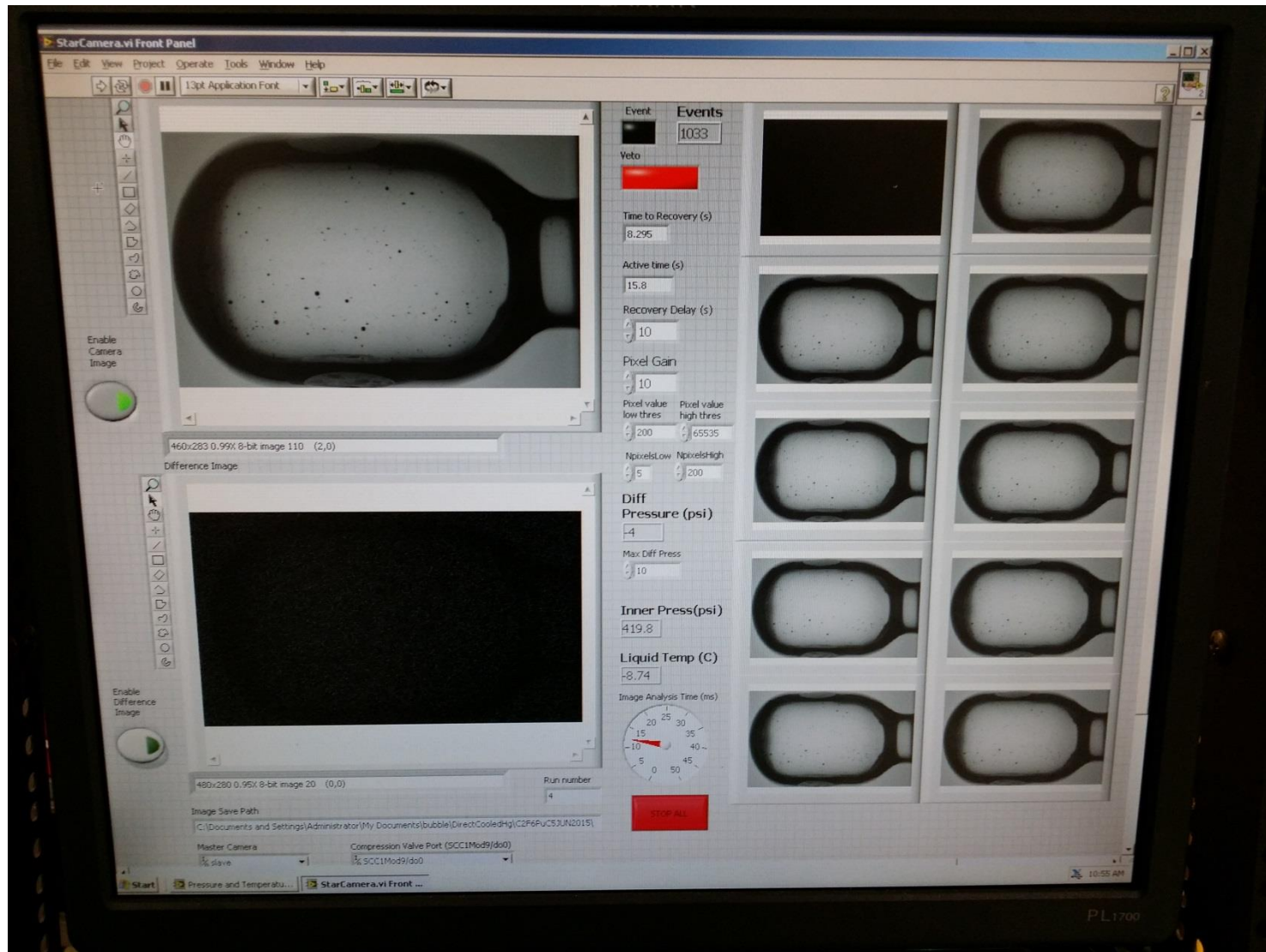
BUBBLE CHAMBER



$T = -10^{\circ}\text{C}$
 $P = 50 \text{ atm}$



USER INTERFACE



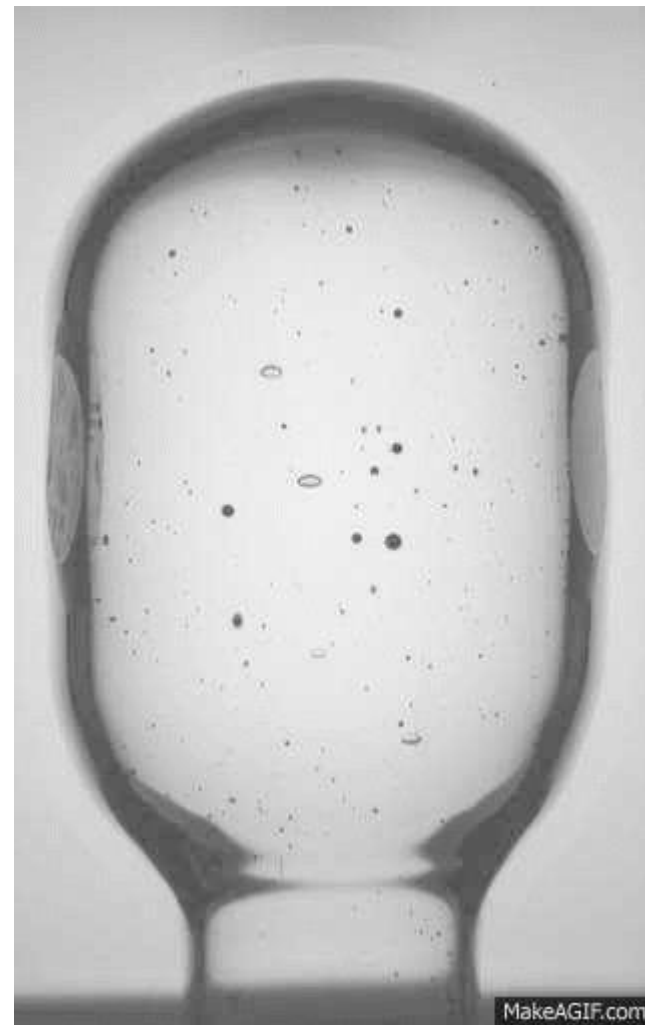
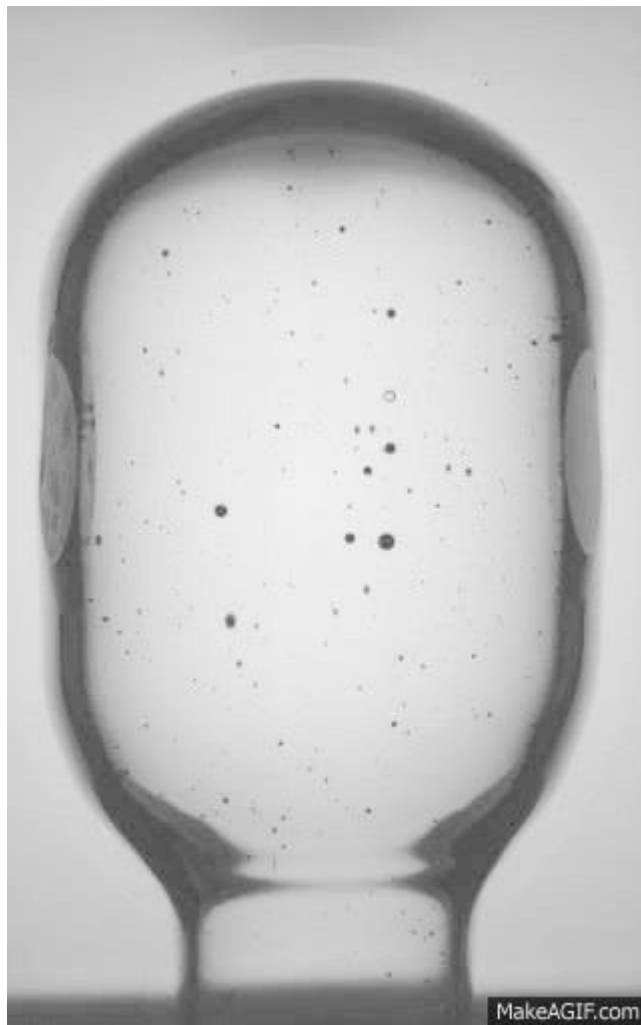
BUBBLE GROWTH AND QUENCHING

3.0 cm



100 Hz Digital
Camera $\Delta t = 10$ ms

N_2O Chamber
with PuC neutron
source



ENGINEERING RUN – MAY 2018

- I. Measure $^{19}\text{F}(\gamma, \alpha)^{15}\text{N}$ ($Q = +4.013$ MeV)
- II. Use C_3F_8 (Octafluoropropane – refrigerant gas)
- III. Only one stable natural isotope (^{19}F)
- IV. Low electron beam kinetic energy (4.6 – 5.2 MeV)
– below threshold of any background reaction

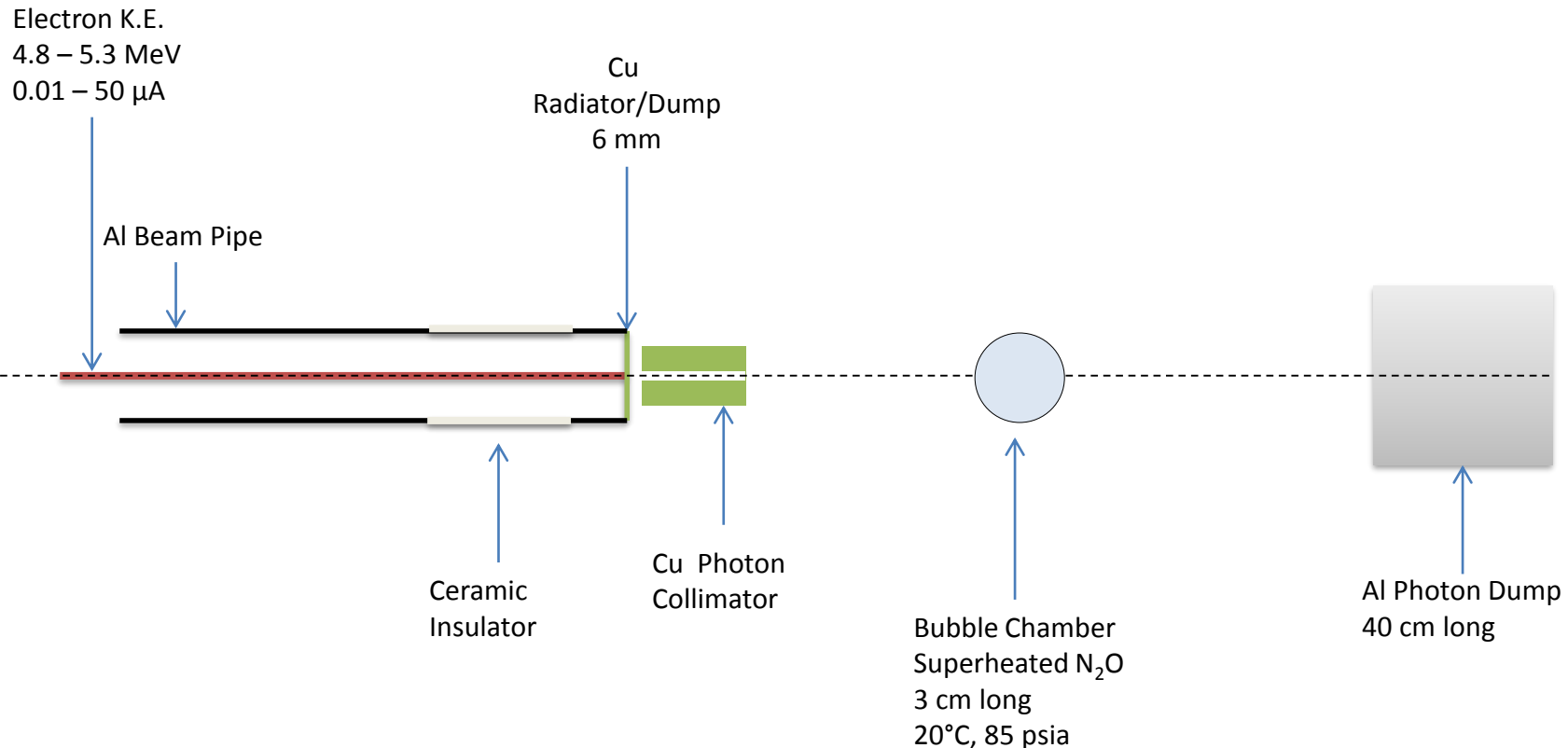
ELECTRON BEAM REQUIREMENTS

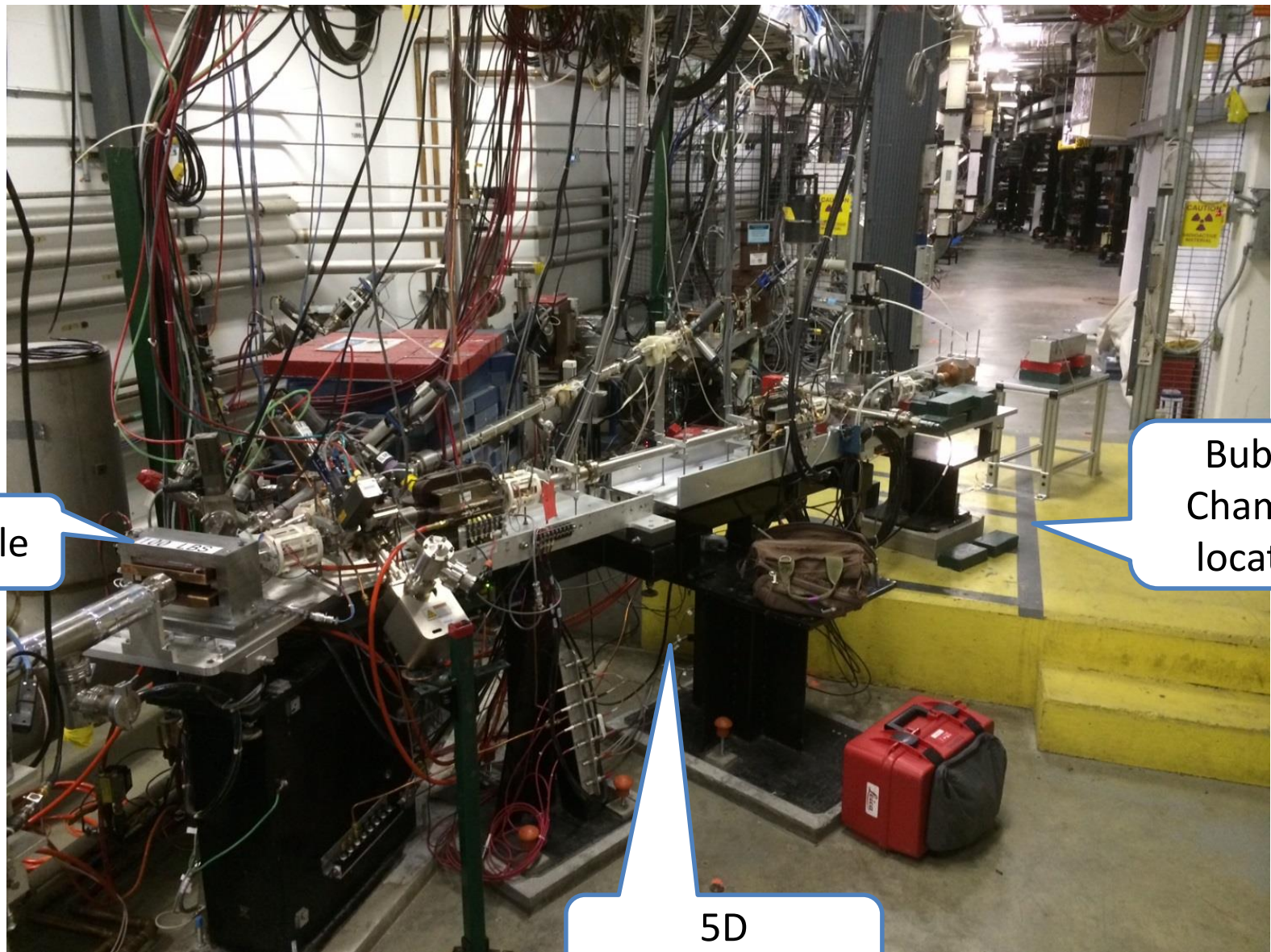
I. Beam Properties at Radiator:

Beam Kinetic Energy, (MeV)	4.8 – 5.3
Beam Current (μA)	0.01 – 50
Absolute Beam Energy Uncertainty	<0.1%
Relative Beam Energy Uncertainty	<0.02%
Energy Resolution (Spread), σ_T/T	<0.06%
Beam Size, $\sigma_{x,y}$ (mm)	1
Polarization	None

SCHEMATICS OF TEST BEAMLINE

- Copper radiator: 6 mm
- Pure Copper and Aluminum (high neutron threshold):
 - I. $^{63}\text{C}(\gamma, n)$ threshold = 10.86 MeV
 - II. $^{27}\text{Al}(\gamma, n)$ threshold = 13.06 MeV



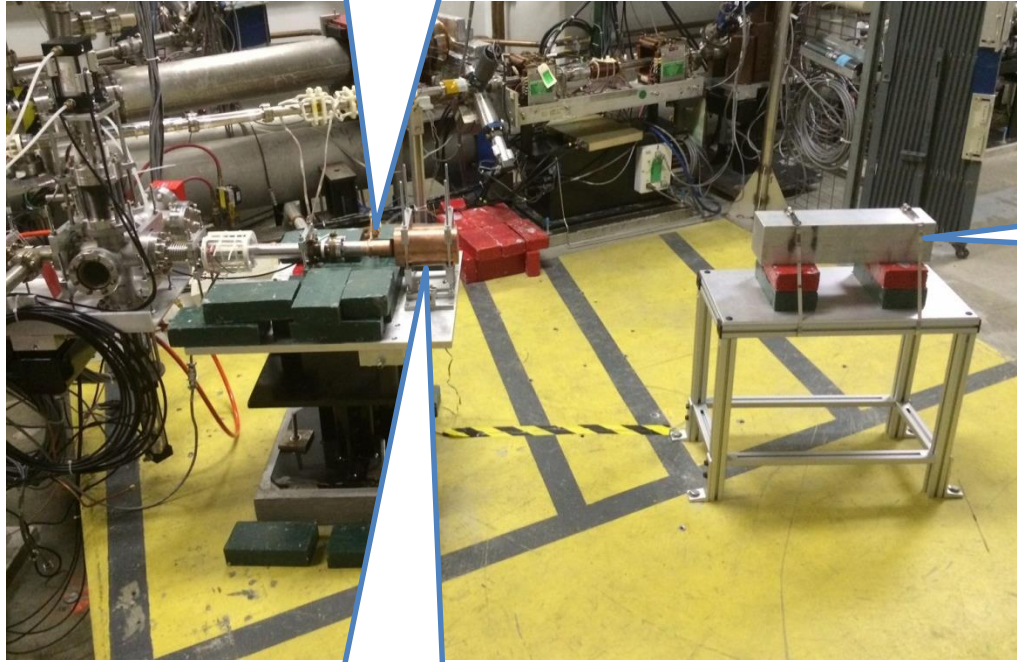


Dipole

Bubble Chamber location

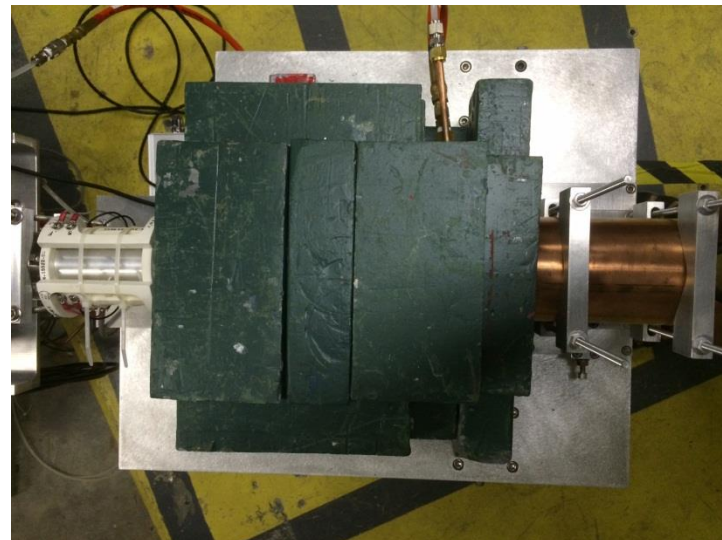
5D Spectrometer

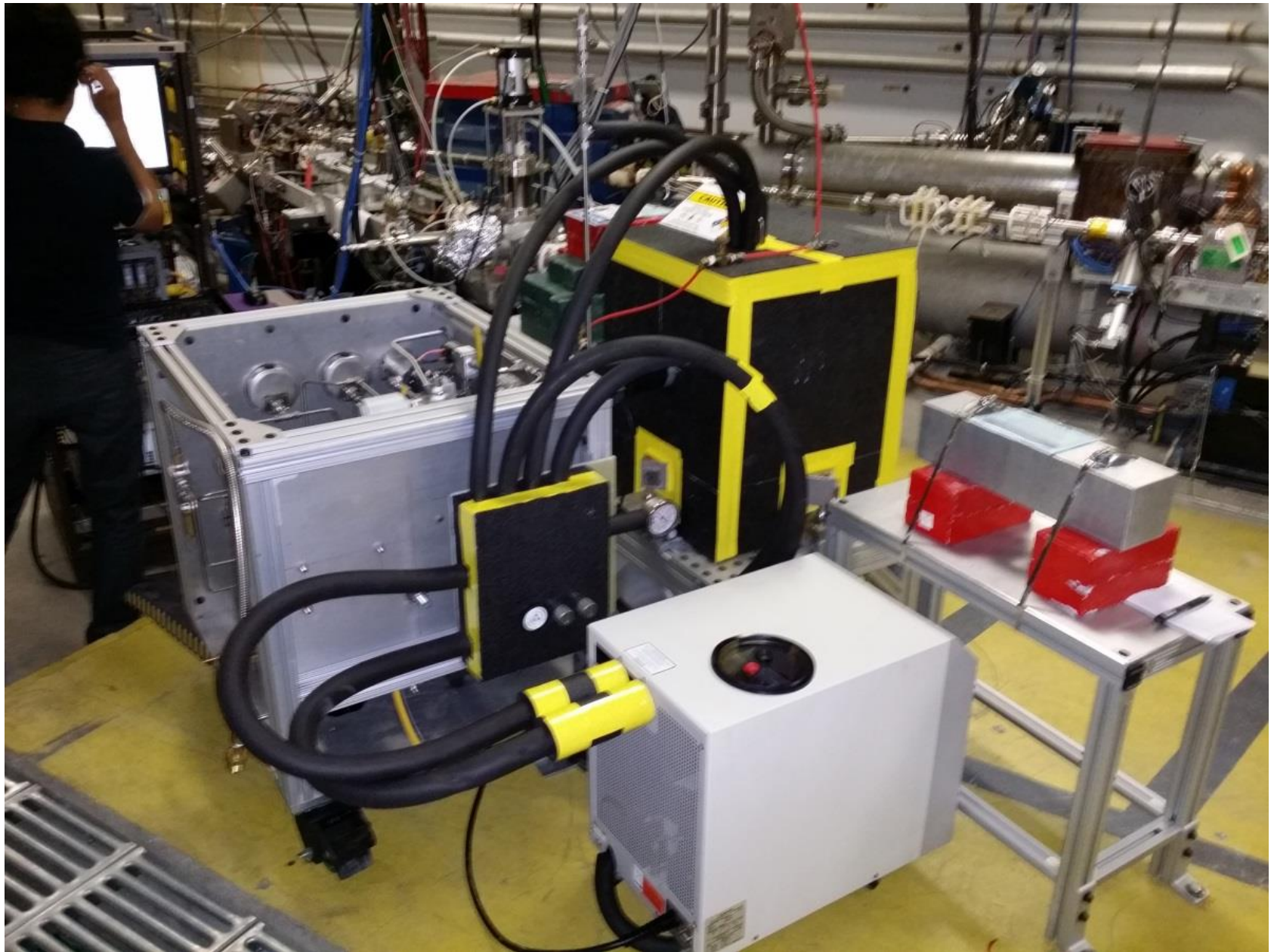
Cu Electron
Radiator/Dump



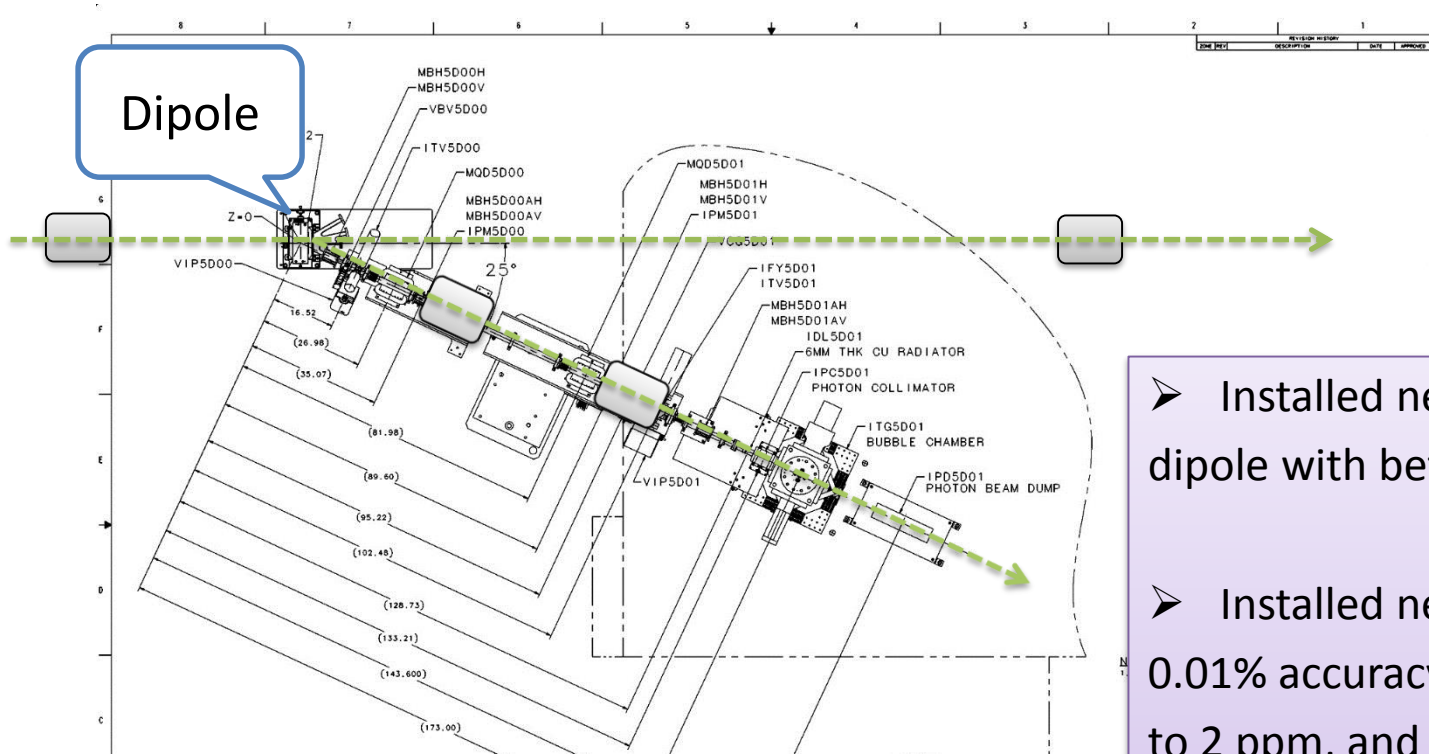
Al Photon
Dump

Cu Photon
Collimator





MEASURING ABSOLUTE BEAM ENERGY



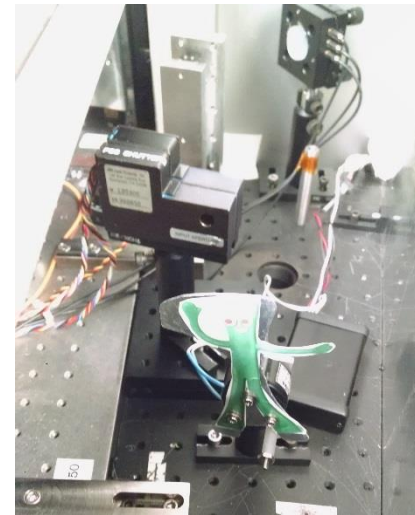
- Installed new higher field dipole with better uniformity
- Installed new Hall probe: 0.01% accuracy, resolution to 2 ppm, and a temperature stability of 10 ppm/°C
- Shielded Earth's and other stray magnetic fields

Electron Beam Momentum

$$p = \frac{\int B dl}{\theta}$$

BUBBLE CHAMBER TEST PLANS

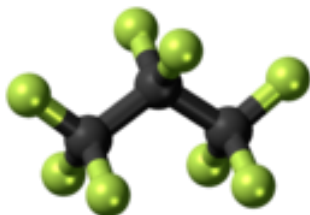
1. Run from MCC fish bowl
2. Use neutron source – test bubble chamber systems operation
3. Measure cosmic background
4. With beam (24/7, May 10 – 18)
 - I. Check CCD camera with beam-on
 - II. Test bubble chamber laser shutter (*gumby shutter*) – Stop beam once a bubble is detected for 5 seconds to allow for time to quench bubble
 - III. Collect bubbles at few beam energies – expect a bubble every 10 sec at 5.3 MeV and every 10 minutes at 4.8 MeV



BUBBLE CHAMBER SAFETY REVIEWS

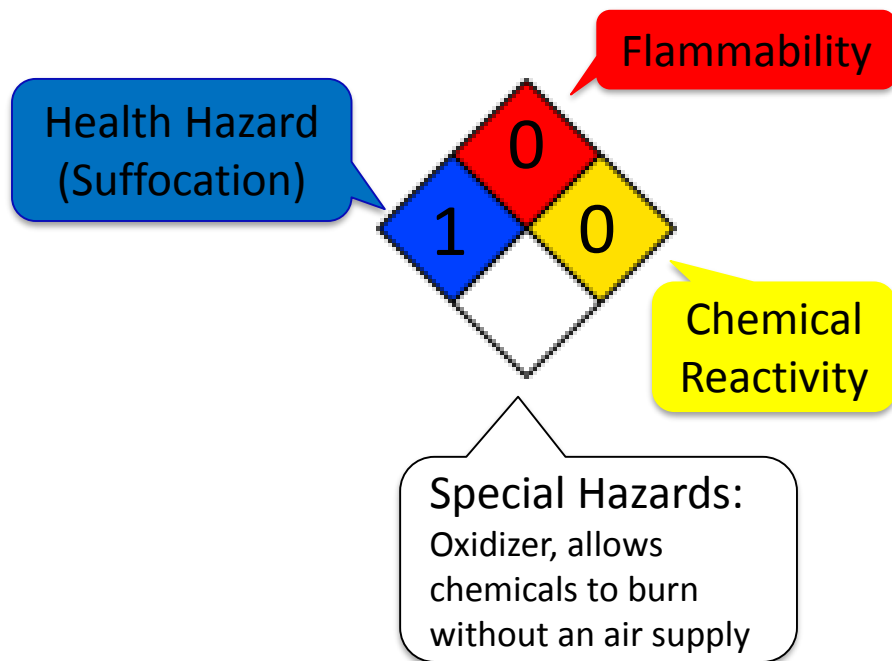
➤ Superheated liquid: C₃F₈ (Octafluoropropane – refrigerant gas)

I. Fluorocarbon non-flammable greenhouse gas



➤ High pressure system:

- I. Design Authority: Dave Meekins
- II. Temperature = 20°C
- III. Operational Pressure = 85 psia
- IV. Quenching Pressure = 350 psia



➤ Temporary Operational Safety Procedures (**TOSP**) is approved