

Radiation Tolerance Issues with Implementation of Silicon Photomultipliers

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*Tech Transfer Workshop
Opportunities with Detector Technologies in Nuclear Physics
Catholic University of America
Washington, DC
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OUTLINE

Introduction to SiPM tech (user view)

Go over some of the key aspects of SiPM operational characteristics

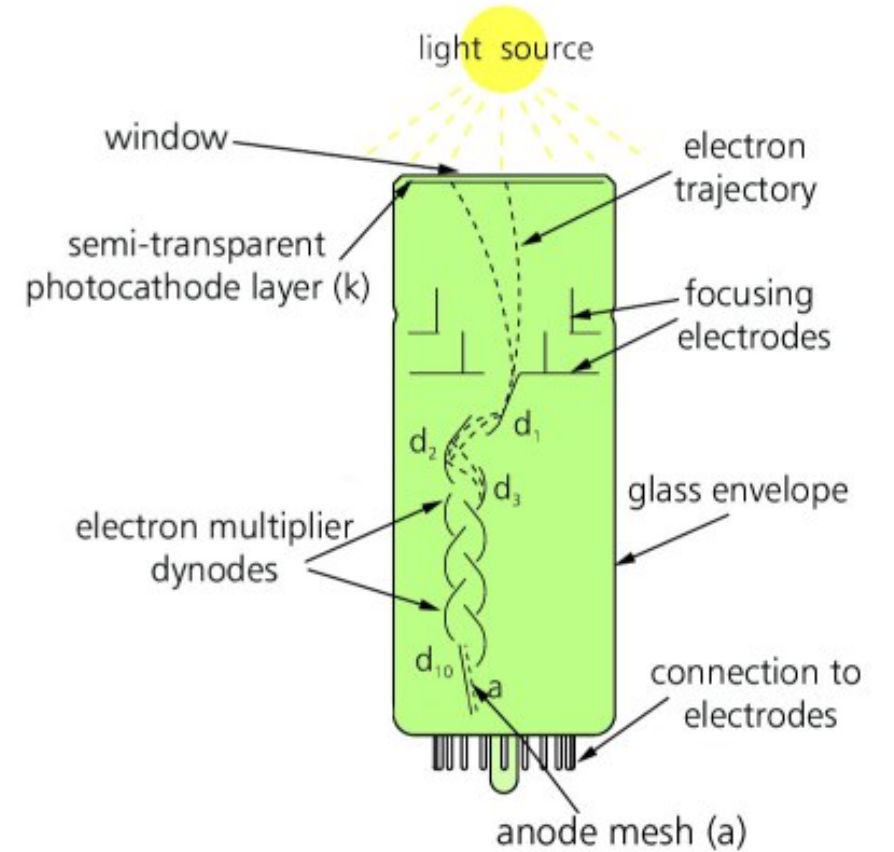
Radiation damage to SiPM (user view)

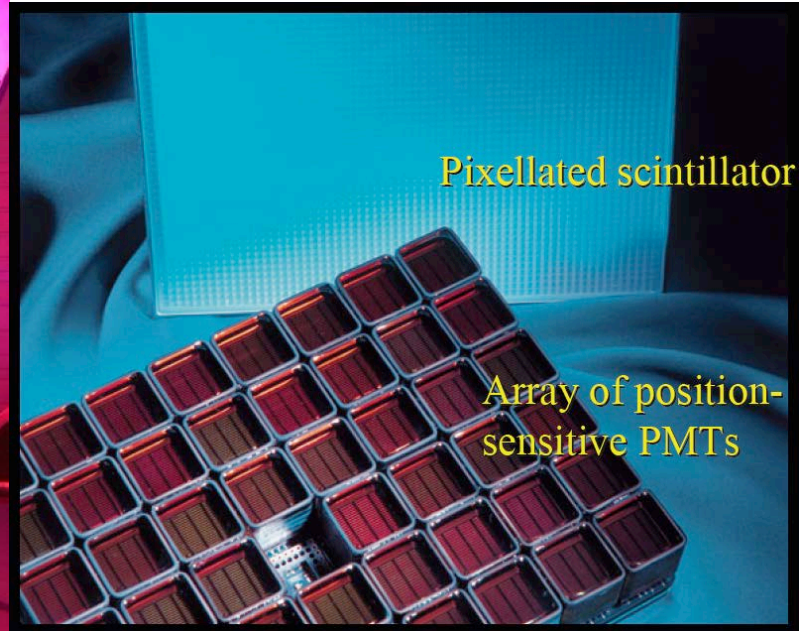
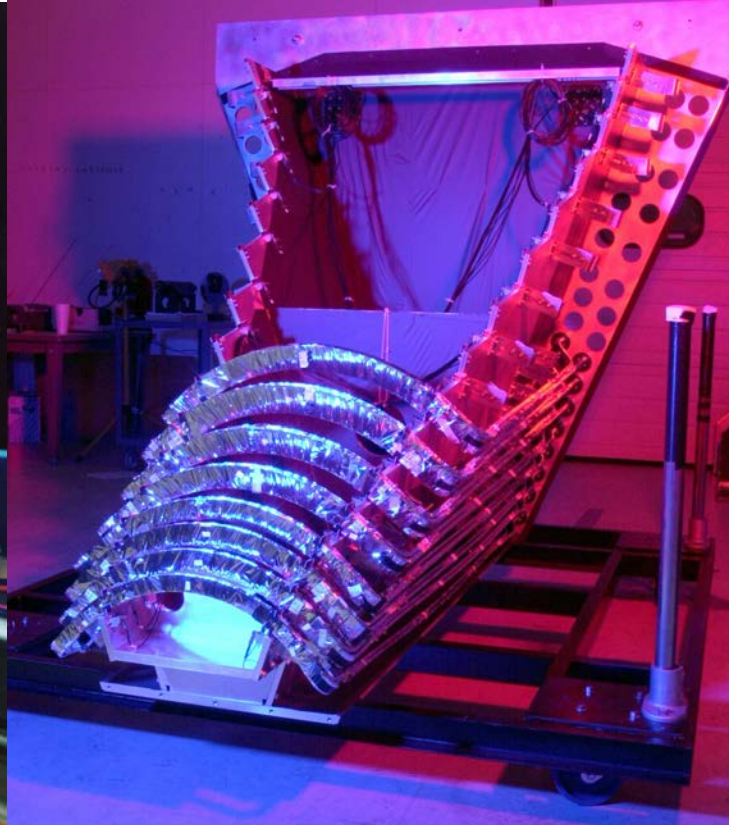
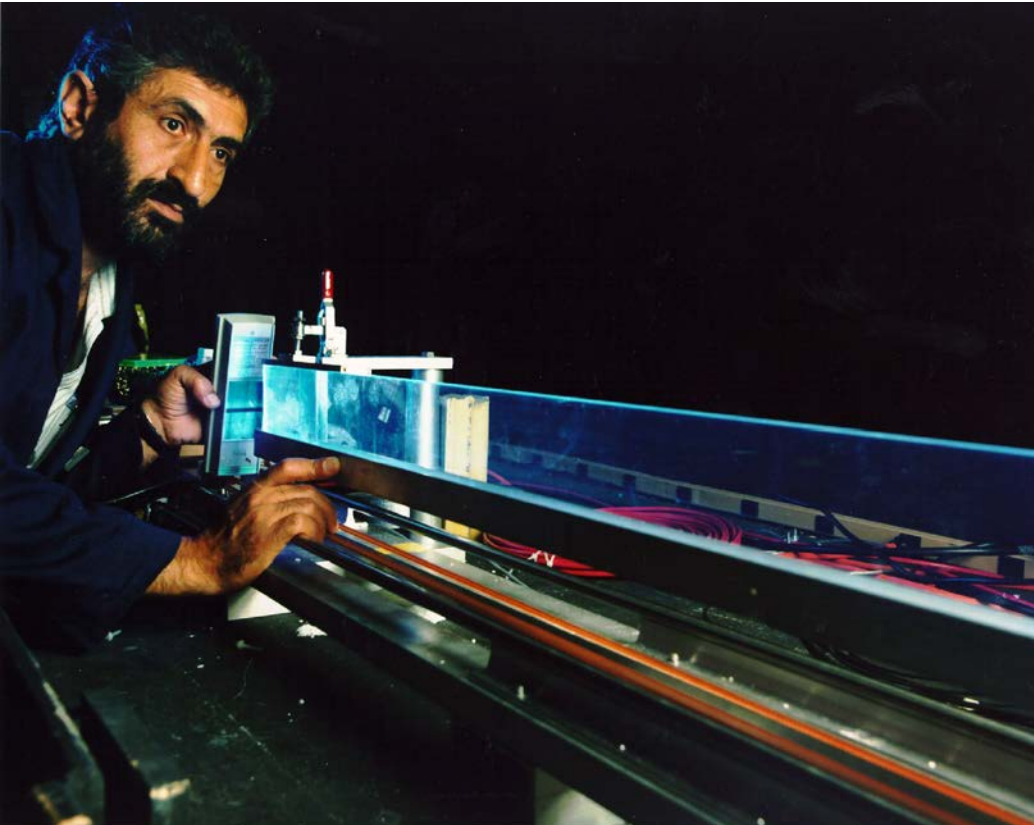
Go over the key phenomena

More importantly, go over some ways to abate the damage

Suggestions for future work

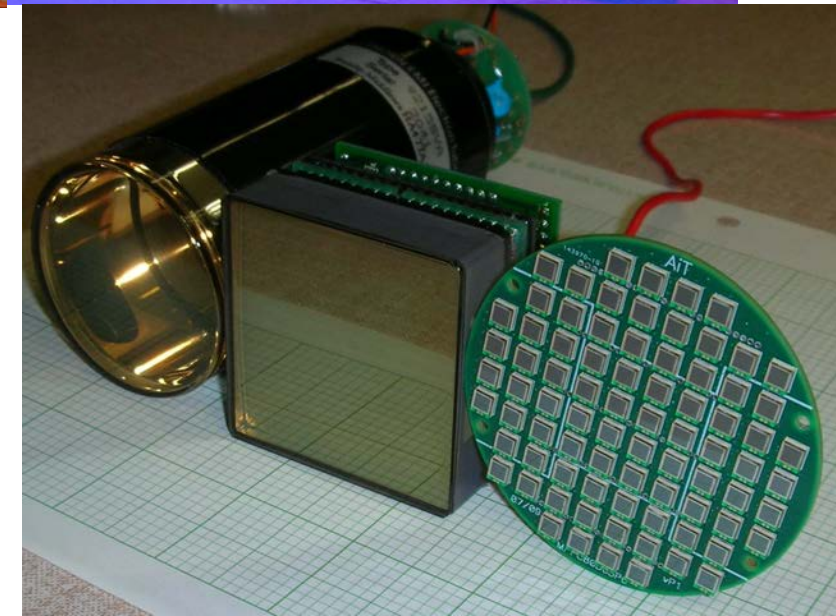
Traditional Photodetector "Workhorse": *Vacuum Photomultipliers*





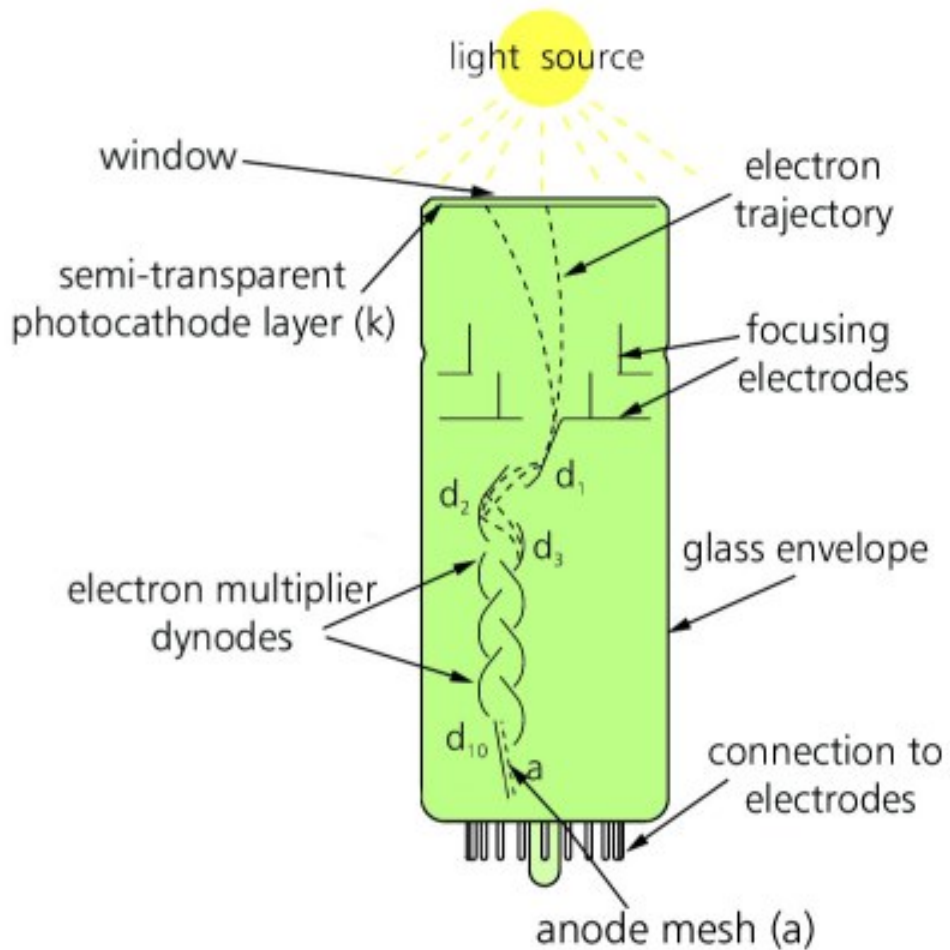
Pixellated scintillator

Array of position-sensitive PMTs

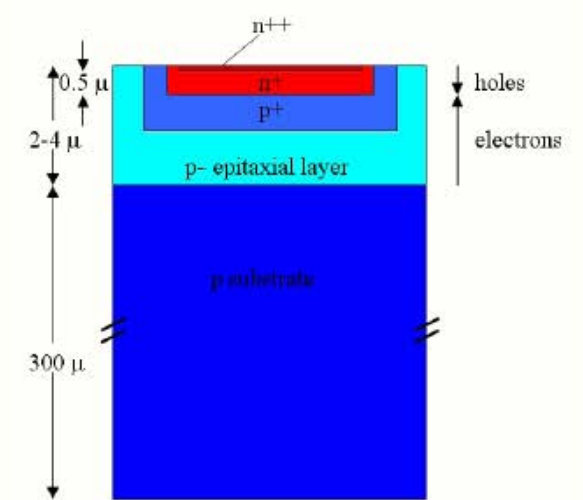
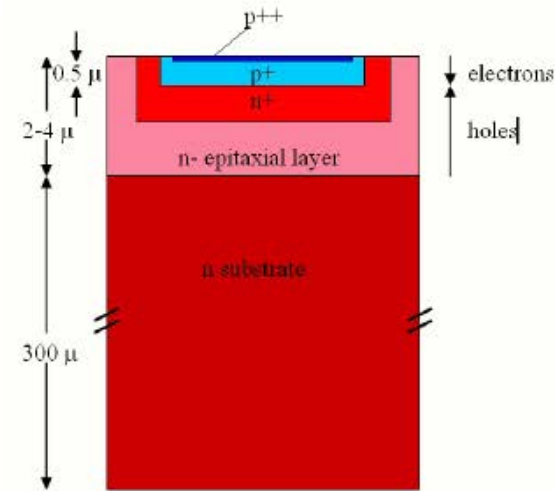


Photomultipliers

PMT

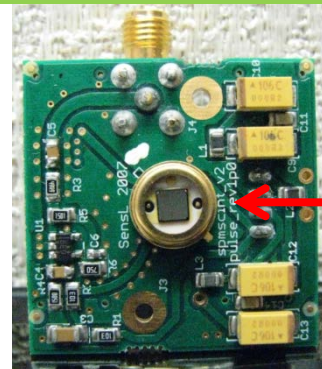


SiPM



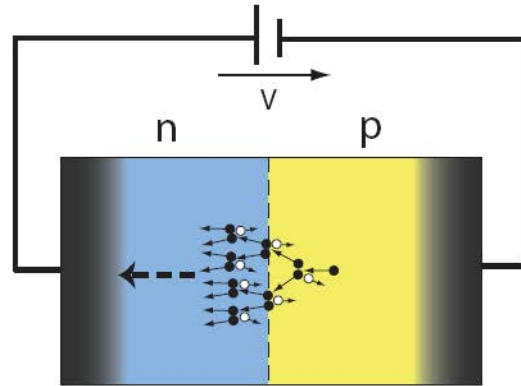
"p on n"
Higher breakdown voltage (70V)
Blue-peaked sensitivity
Less dark noise

"n on p"
Lower breakdown voltage (30V)
Green-red sensitivity
More dark noise

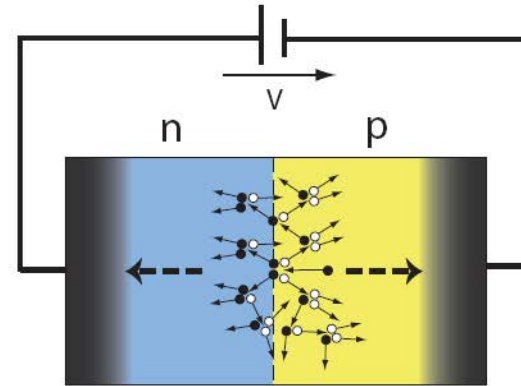


3x3 mm²

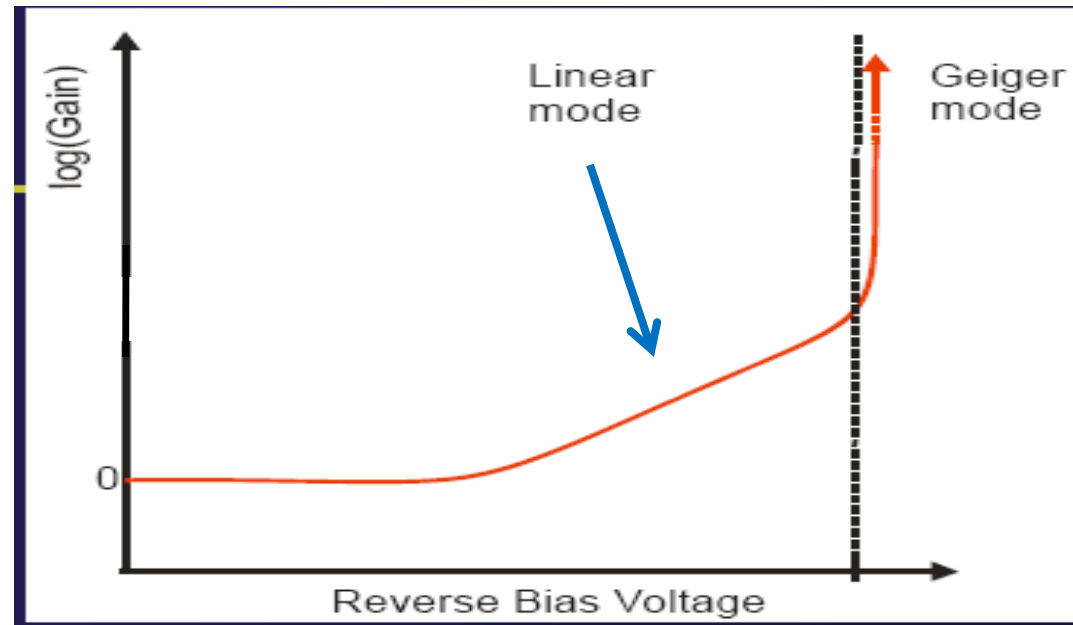
Avalanche Photodiode (APD)



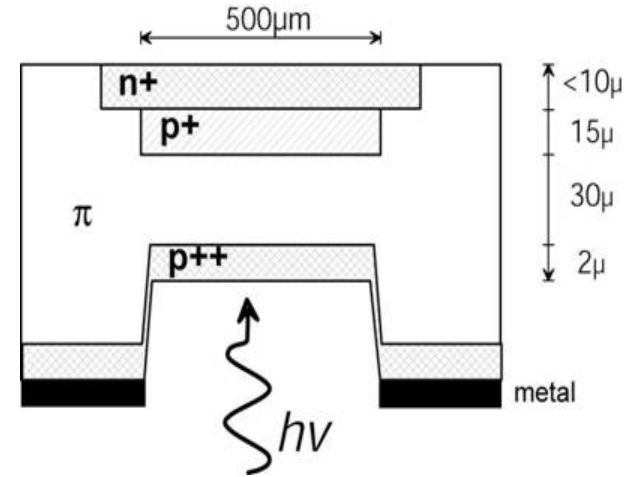
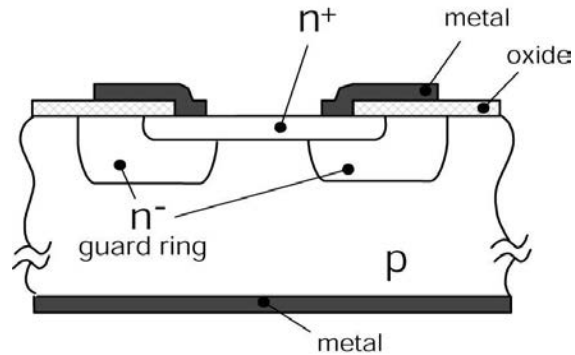
Linear Mode



Geiger Mode

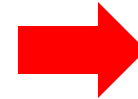


A Bit of History



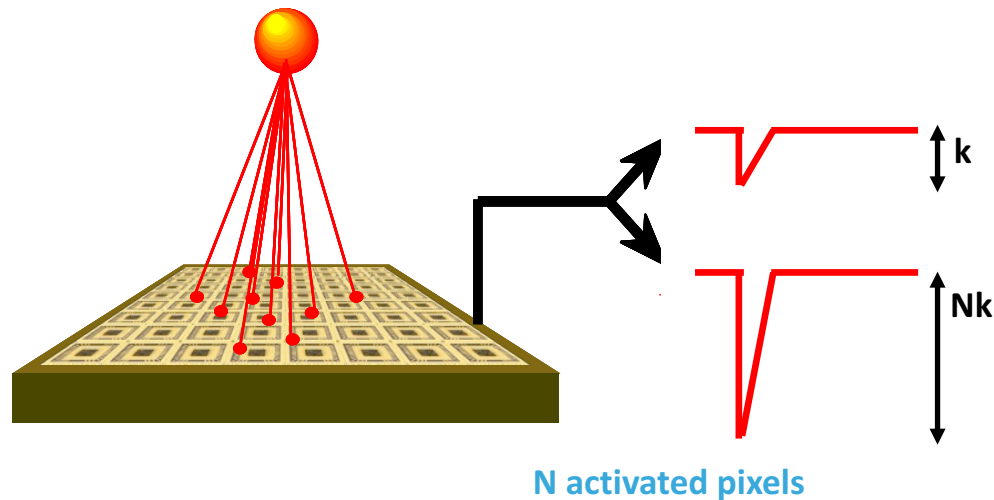
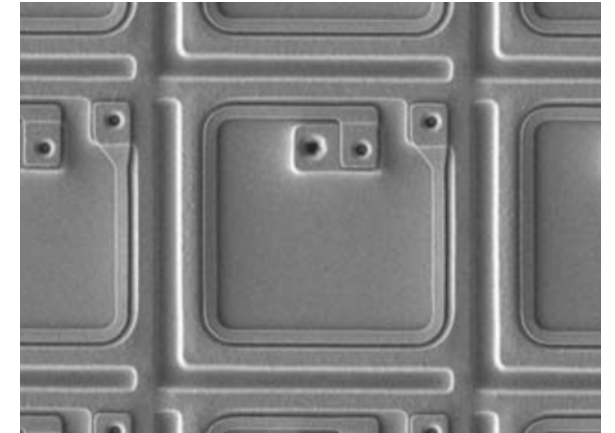
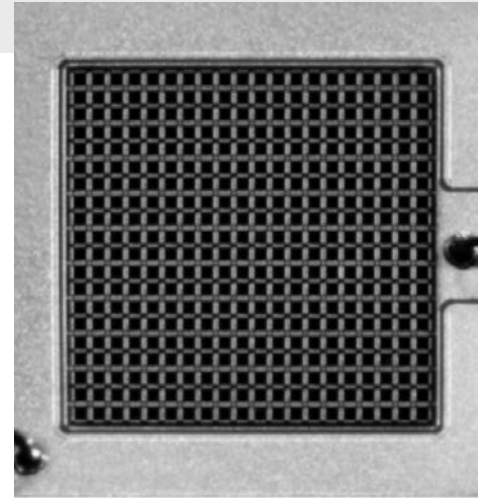
First geiger-mode APDs made in 60/70s could only be used as single photon counters.

Still available today - commonly used in biomolecular work

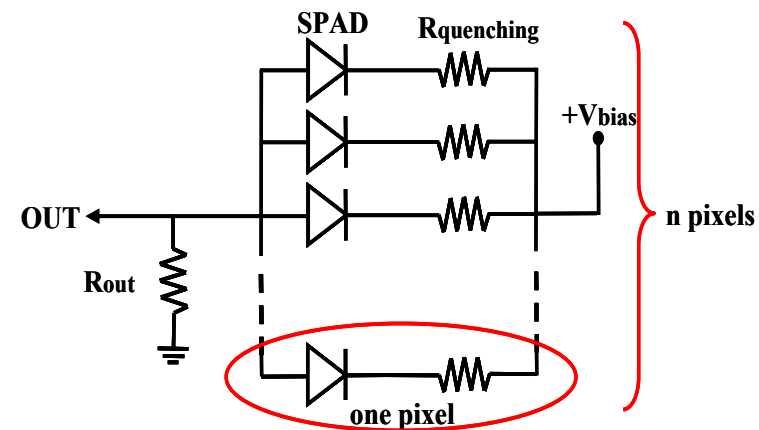


SiPM \rightarrow Geiger mode SPAD array

- Array of several micro-cells, connected in parallel
- Each cell is a Single Photon Avalanche Diode (SPAD)
- The device is biased above its breakdown voltage
- Each cell is sensitive to one photon (digital response)
- The whole array is an analogue device

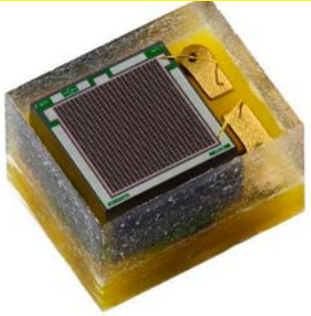


SiPM equivalent circuit

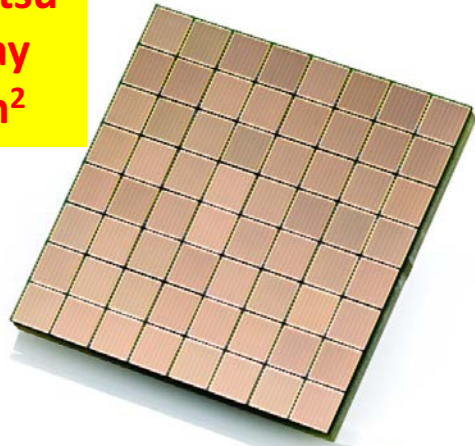


Pics of SiPMs

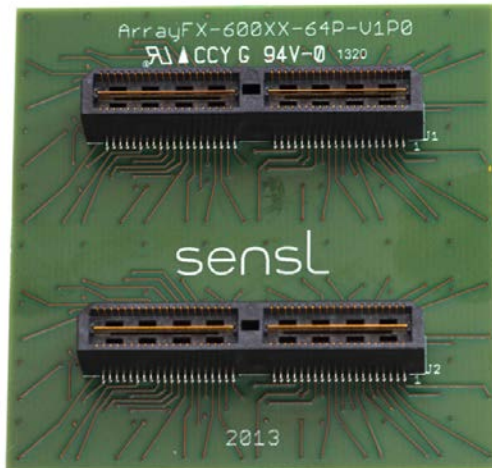
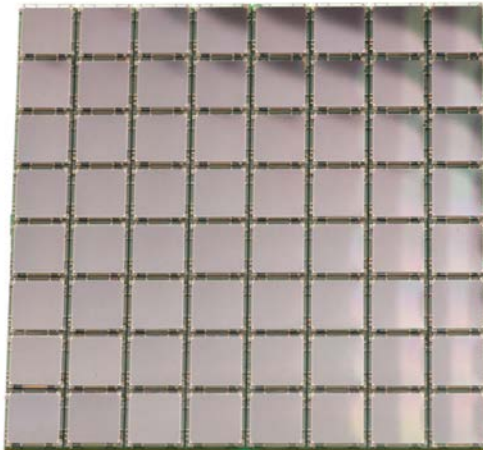
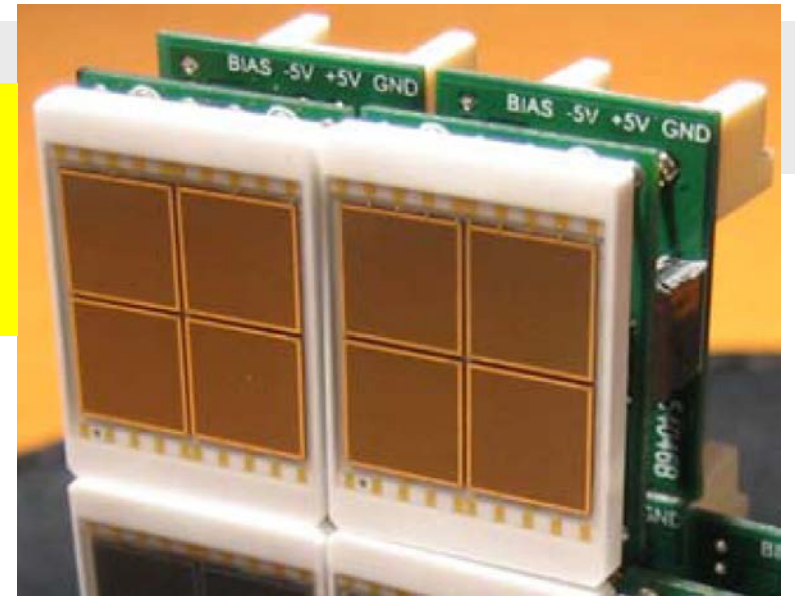
1x1 mm² Ketek



Hamamatsu
8x8 array
3x3 mm²



Hamamatsu
Hall D array
4x4 array
3x3 mm²

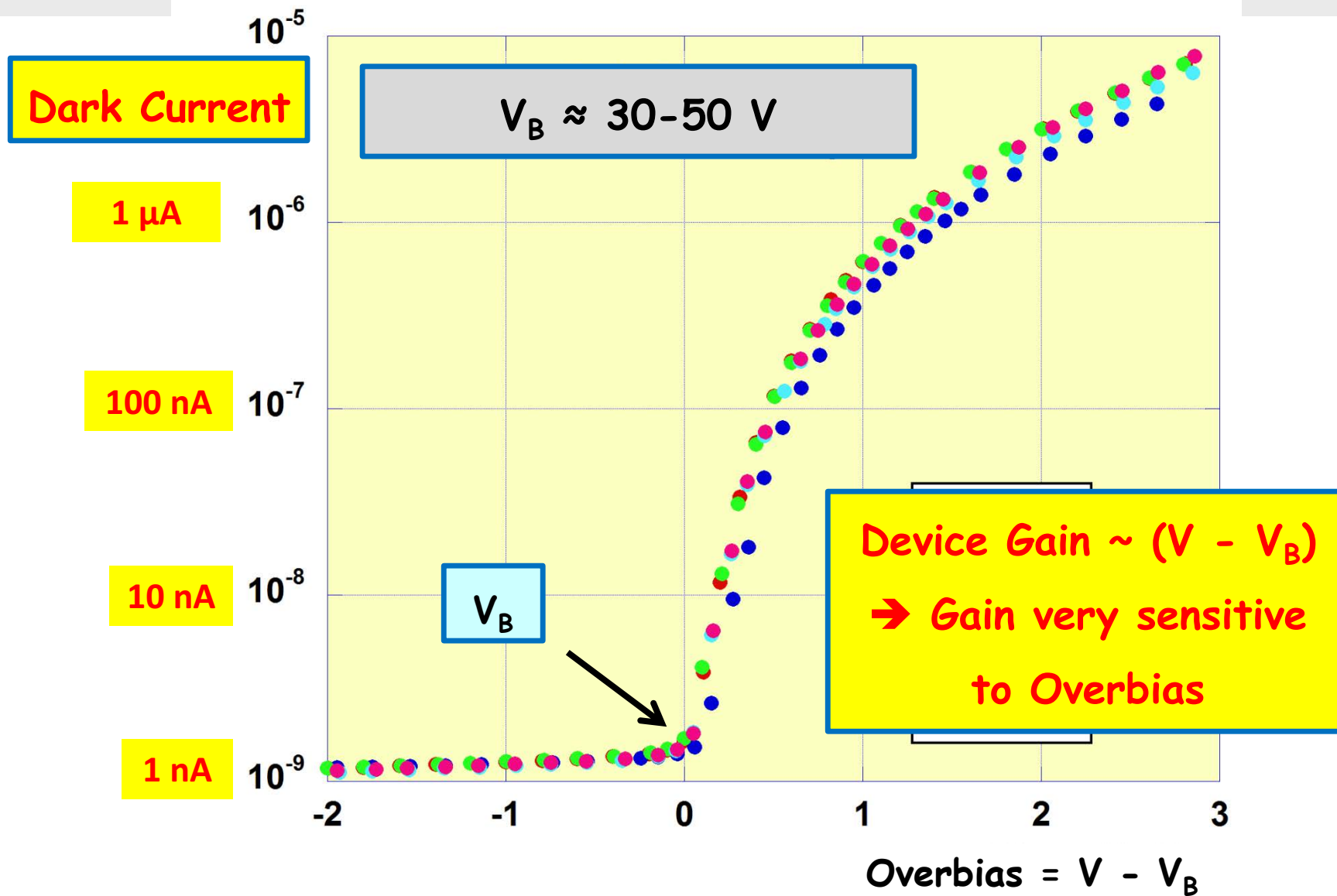


SensL
8x8 array
3x3 mm²



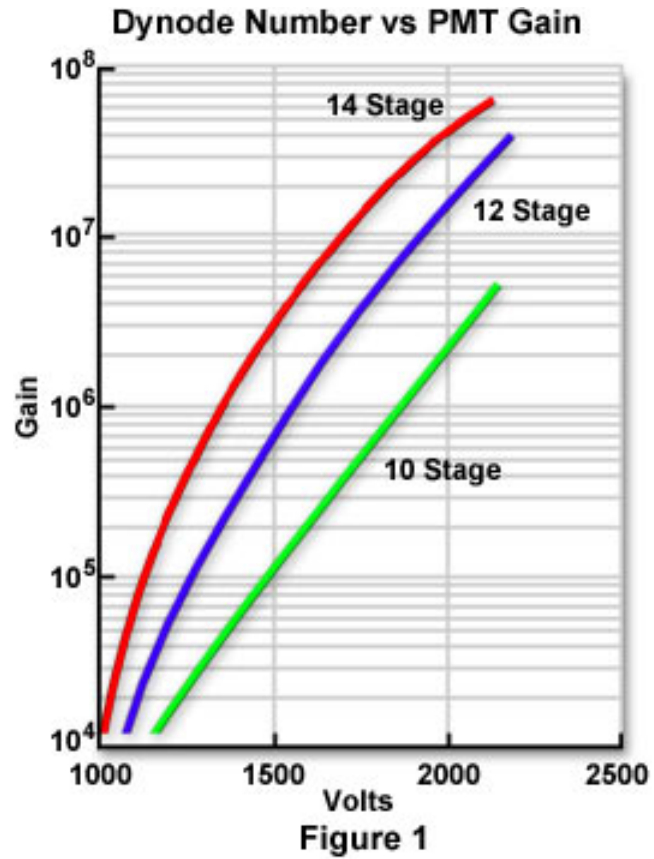
16x16 array
3x3 mm²
SiPM replacement
for 5x5 cm² PSPMT

Turning on a SiPM (in the *Dark*)

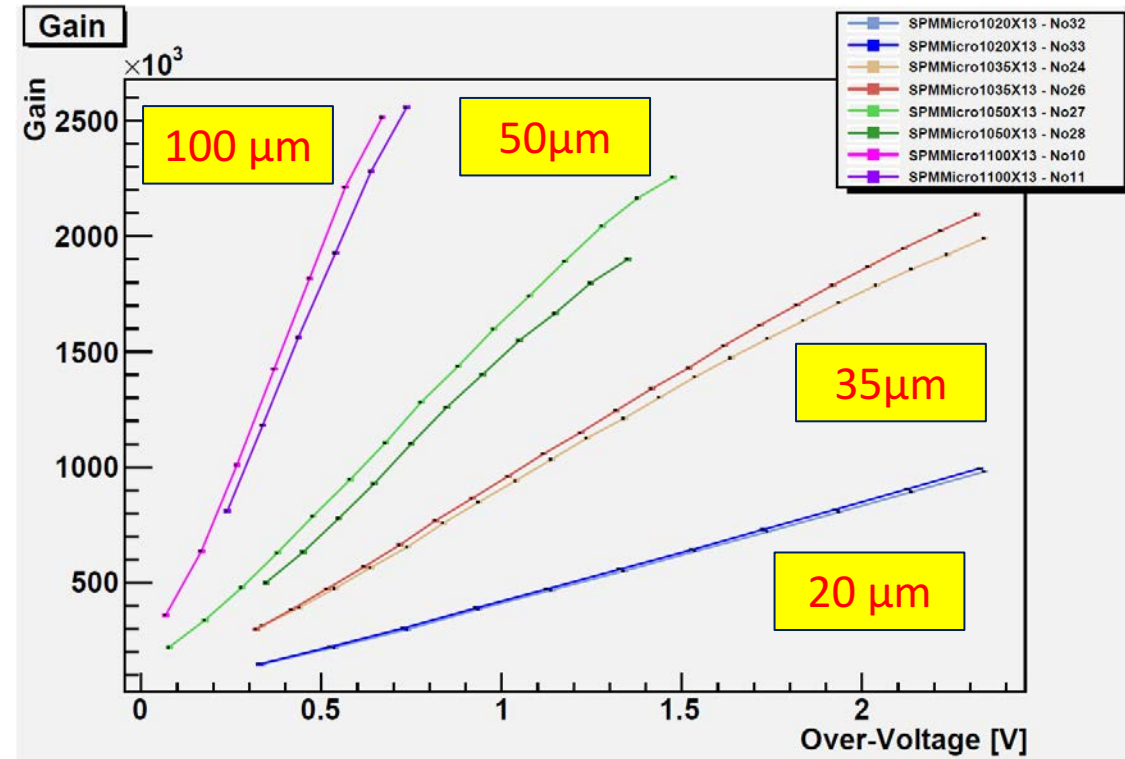


Signal Gain

PMT



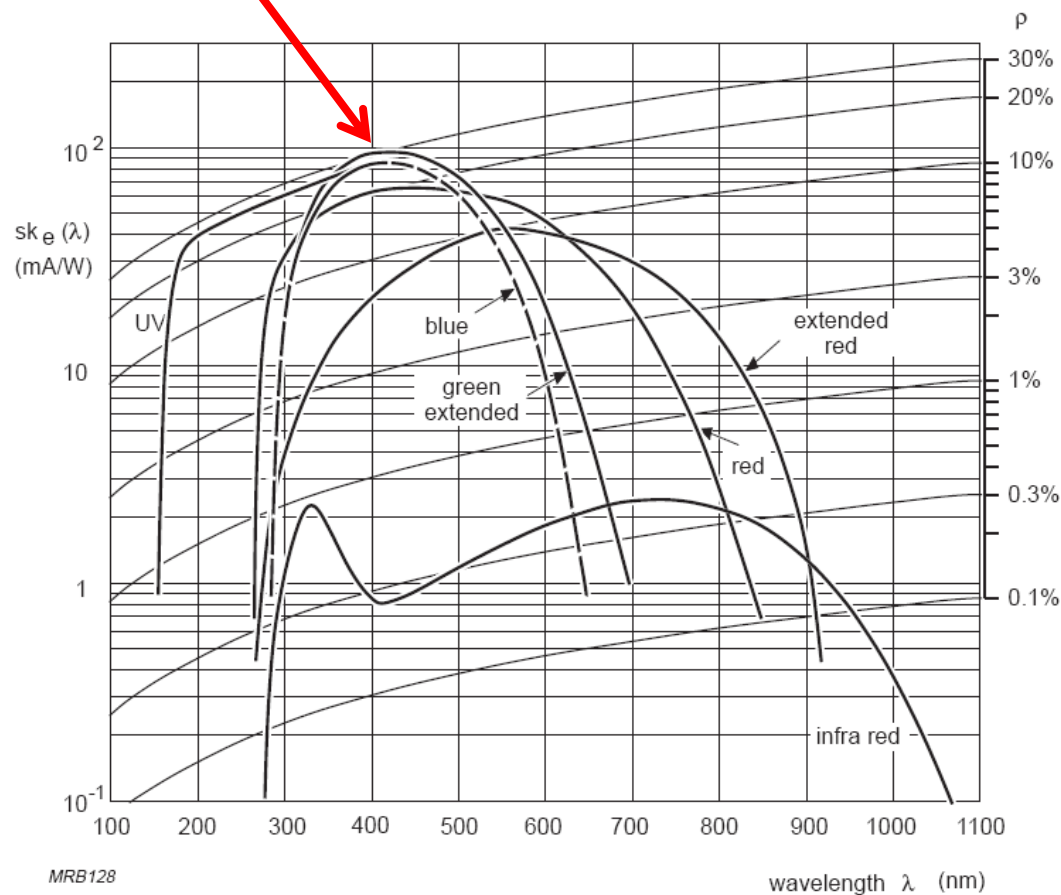
SiPM*



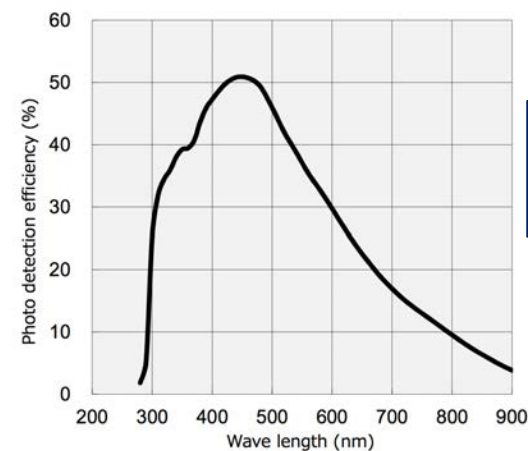
Quantum Efficiency vs Wavelength

PMT

20-30% peak is typical

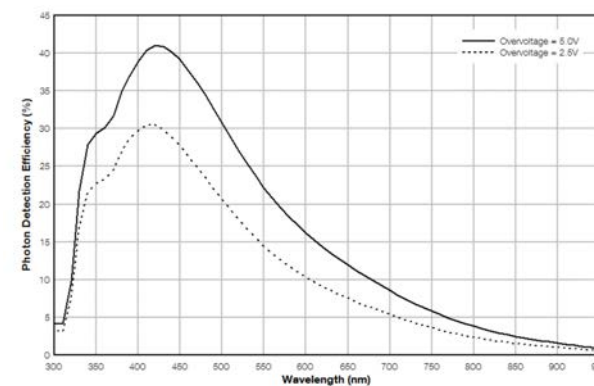


SiPM

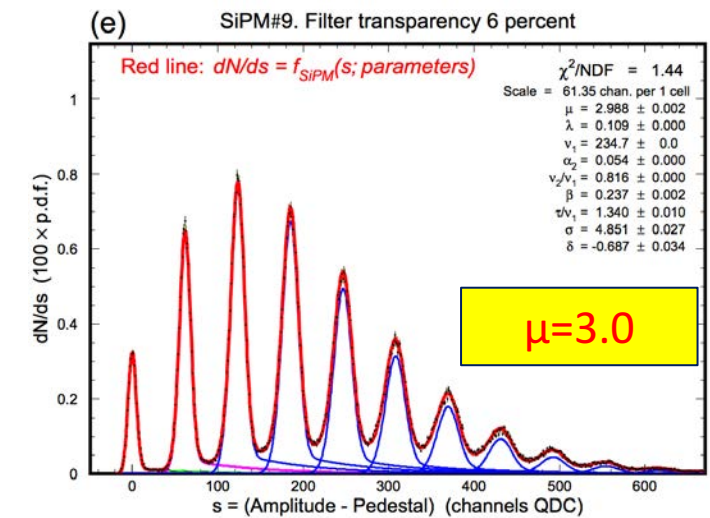
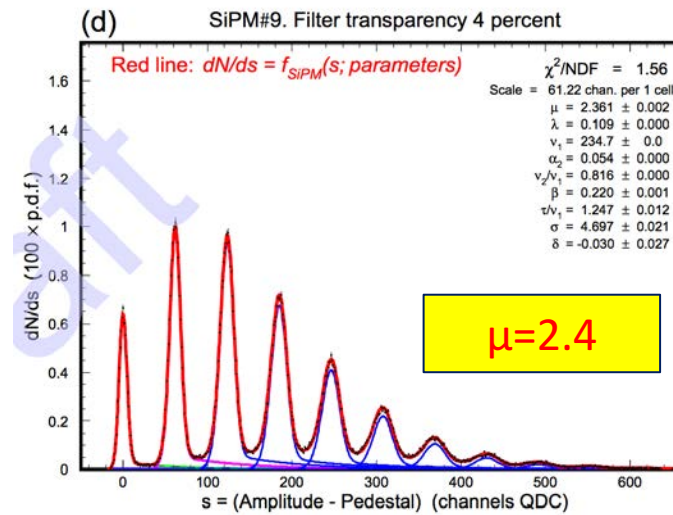
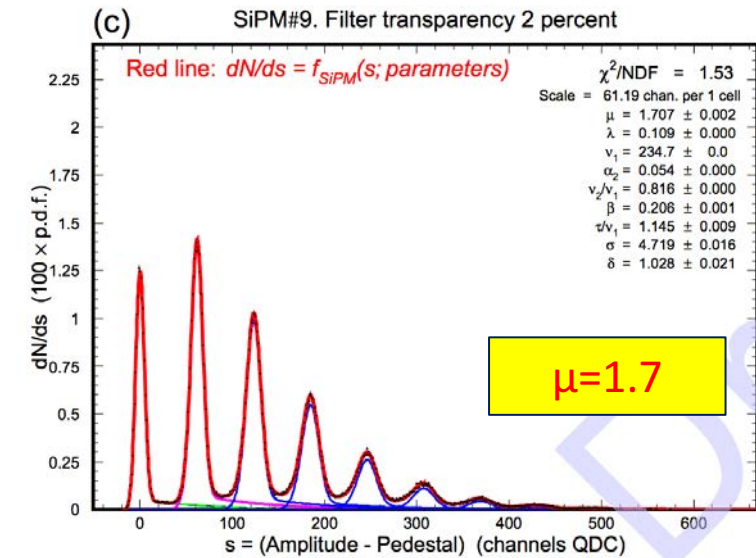
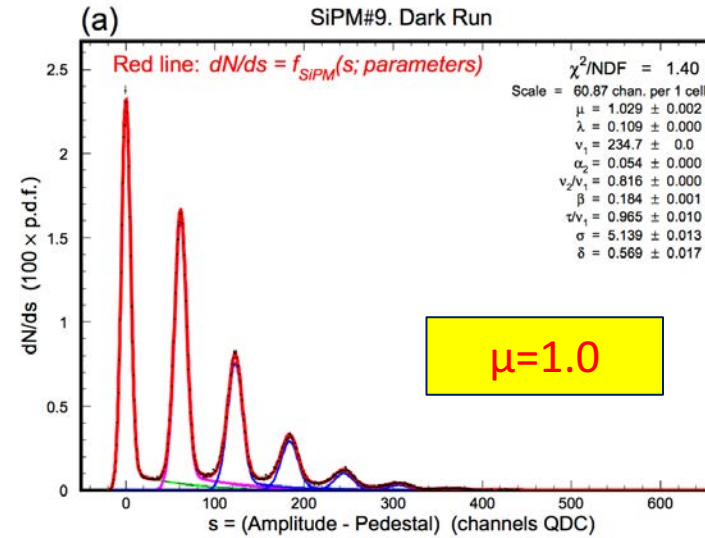
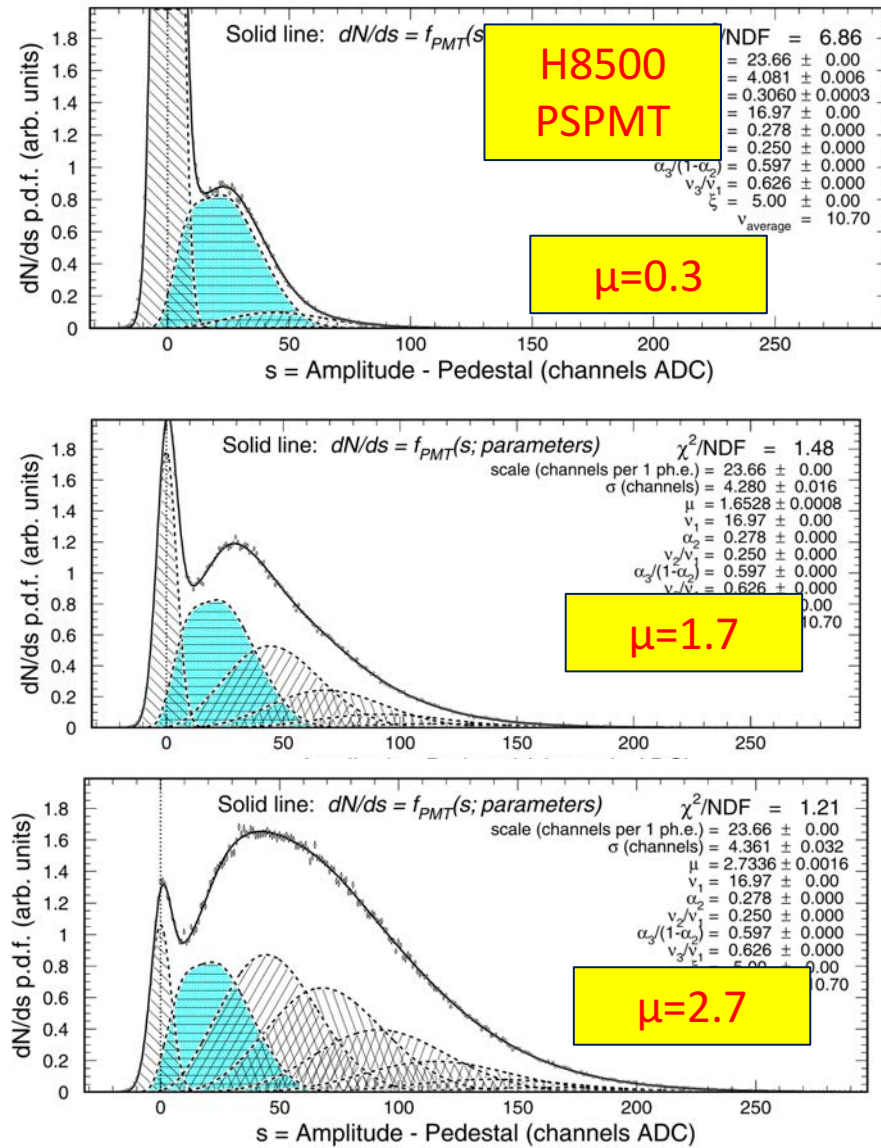


Hamamatsu
14160 series

~40-50% peak is typical now

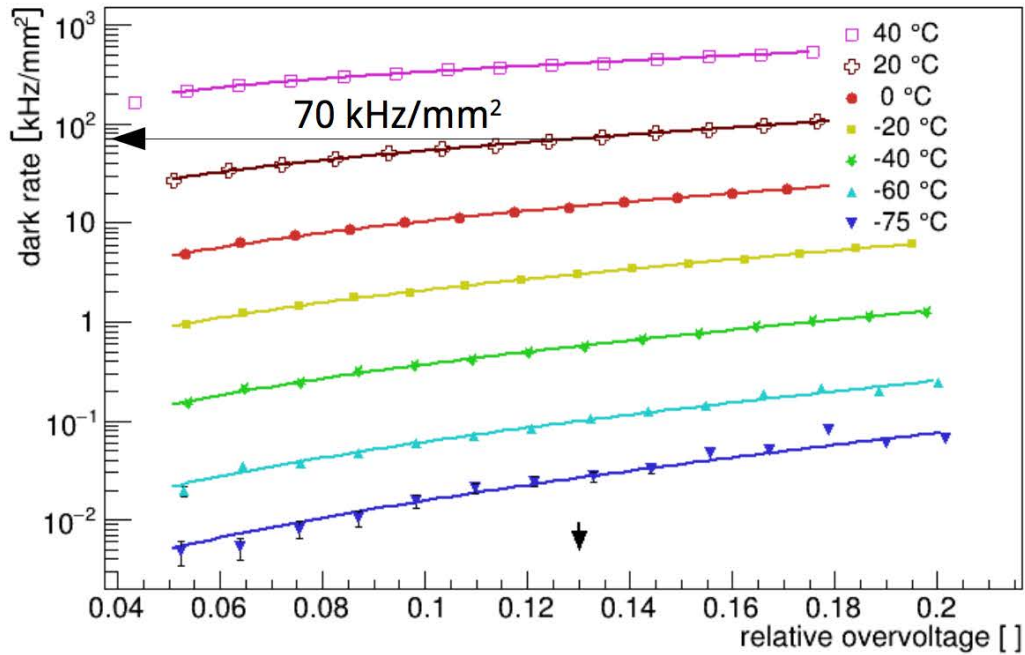


SensL
C series

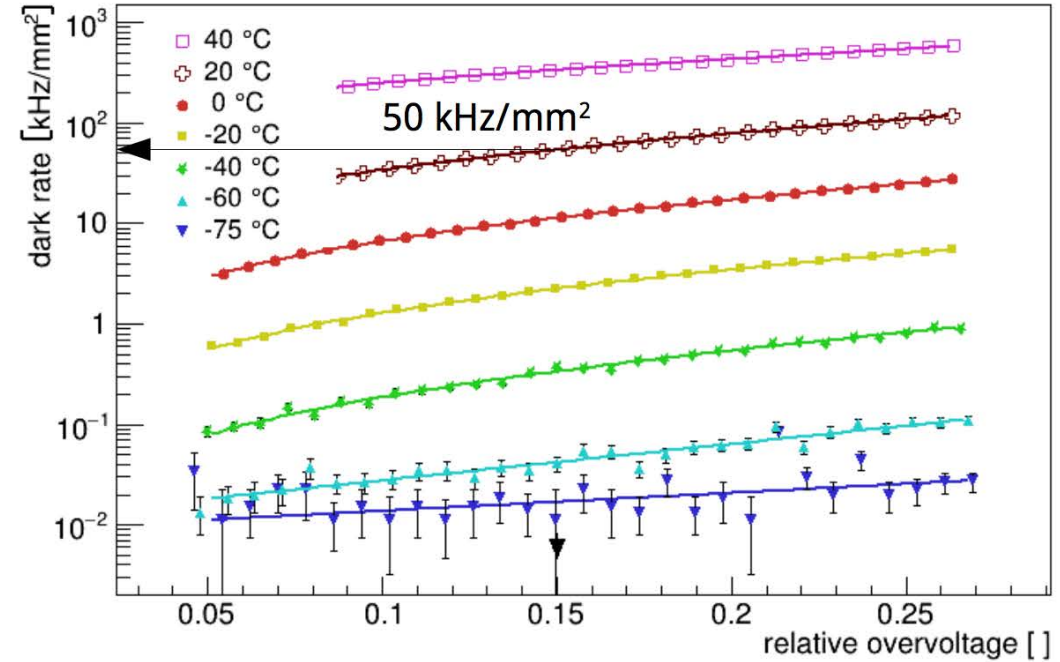


Dark Rate vs Temperature

Hamamatsu LVR2 6050 CN

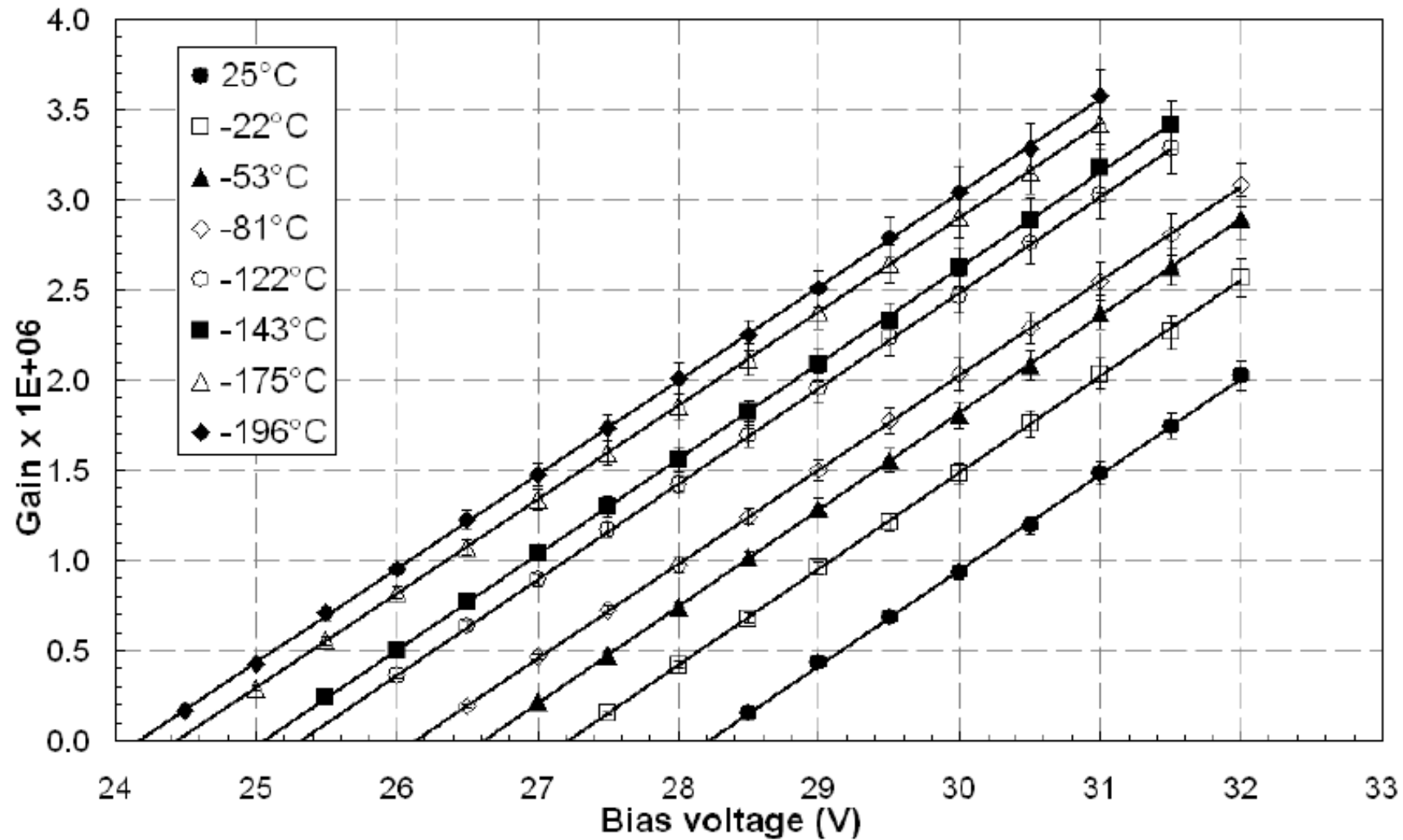


KETEK PM 3325 WB



Otte et al. NIM A 846 (2017) 106-125

Gain vs Temperature



← V_B as temp. decreases

Effect of transverse field on Burle PMT

43 gauss

75 gauss

B field



Top View

Burle
Planacon

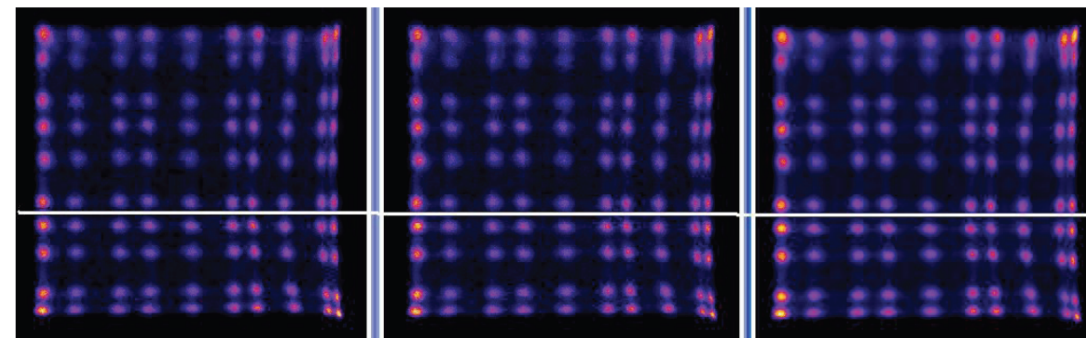


100 gauss

133 gauss

$\mathbf{v} \times \mathbf{B}$ gives downward force

Effect of MRI fields on SiPM

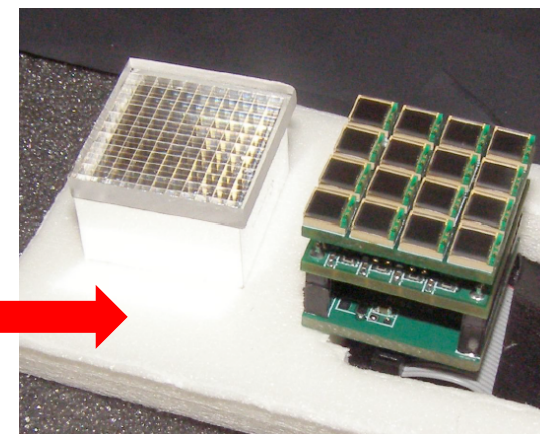


0 Tesla

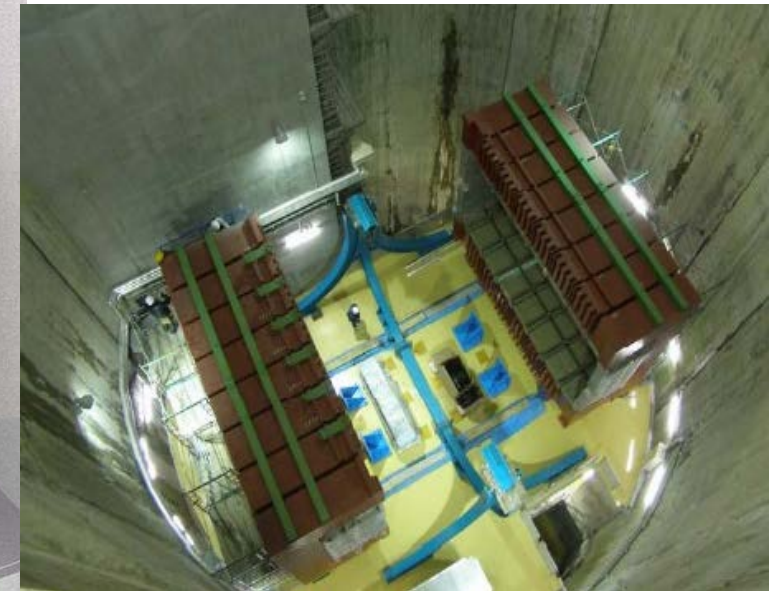
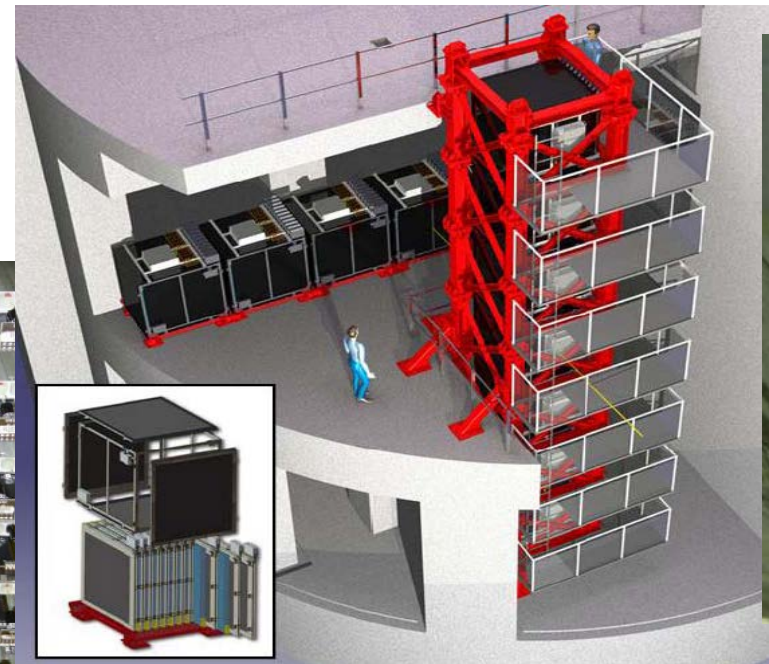
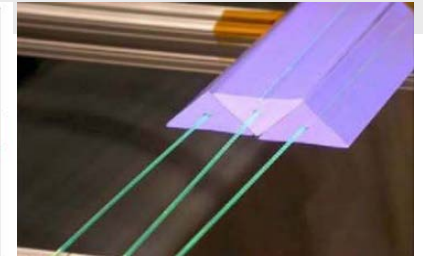
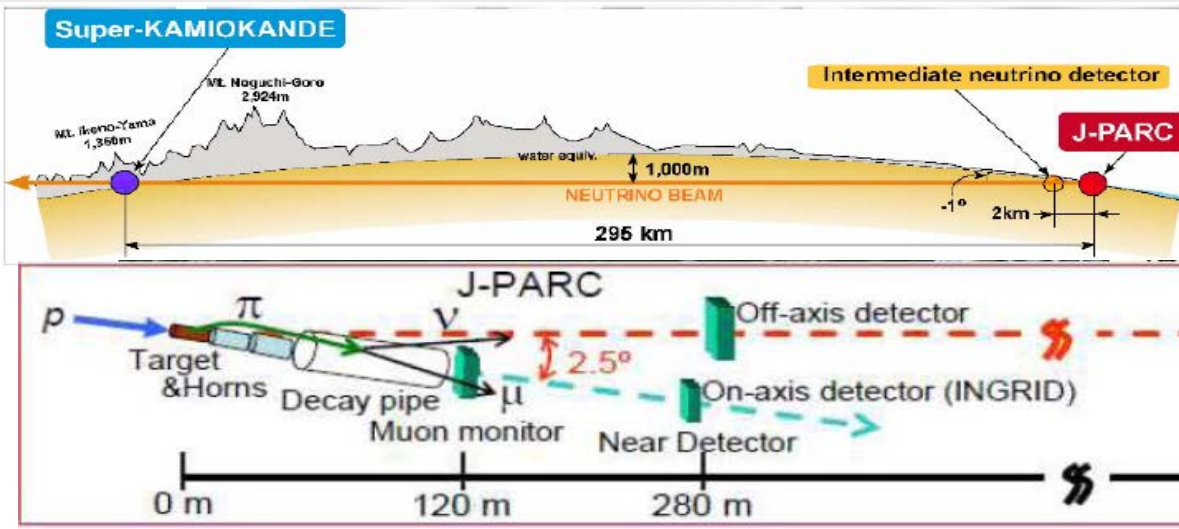
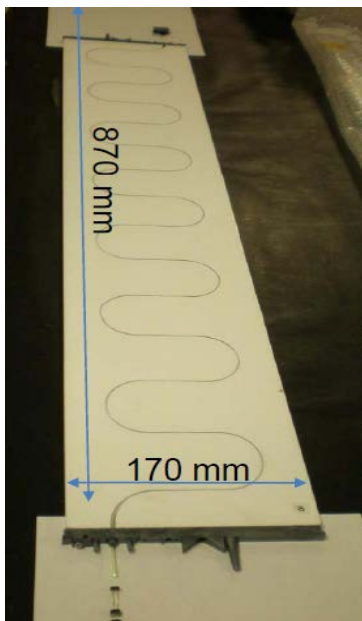
3 Tesla ||

3 Tesla ⊥

Array of SiPMs +
scintillating crystal
array



First High Volume Use of SiPMs - T2K neutrino experiment - Japan



T2K

Super-KAMIOKANDE

ML. Teno-Yama
1,360m

ML. Naguchi-Goro
2,524m

Intermediate neutrino detector

J-PARC

water equiv.

1,000m

NEUTRINO BEAM

296 km

-1°

2km

J-PARC

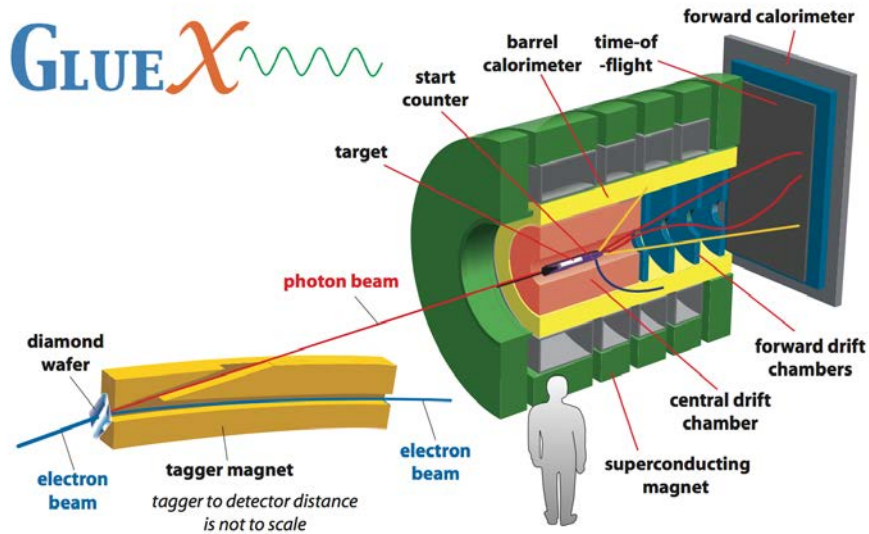
Off-axis detector

2.5°

Far Detector
(SK)

- > 60,000 delivered by Hamamatsu ($\sim 0.1 \text{ m}^2$)
- $1.3 \times 1.3 \text{ mm}^2$ specifically designed for T2K
- Well suited for 1 mm diameter fiber
- 667 pixels
- 26×26 array of $50 \text{ }\mu\text{m}$ pixels
- Dark noise < 1.2 MHz at nominal voltage ($\sim 700 \text{ kHz/mm}^2$)
(7.5×10^5 gain at 25°C)

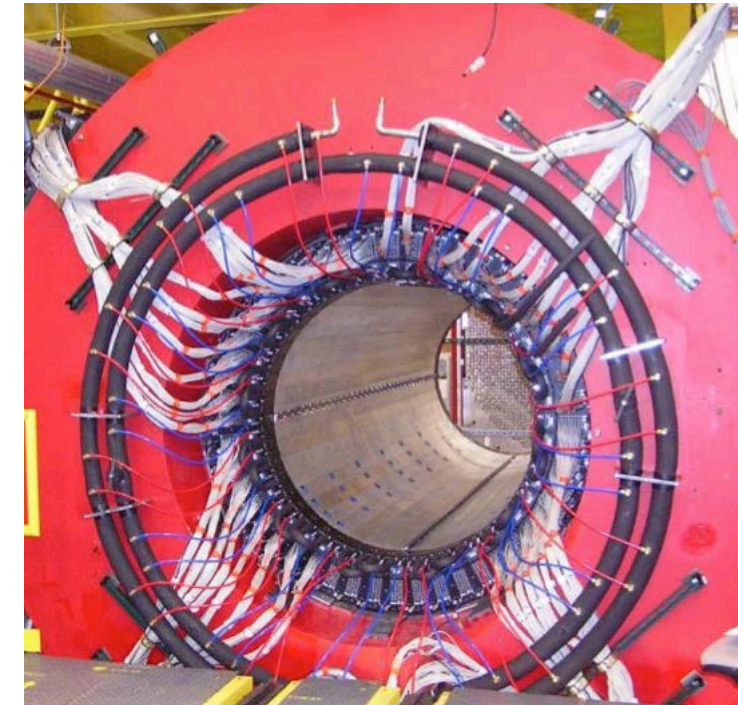
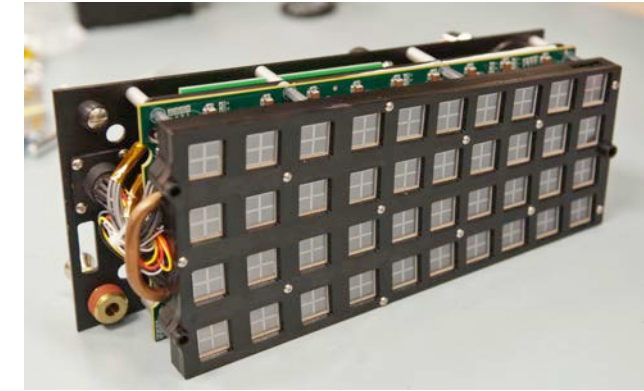
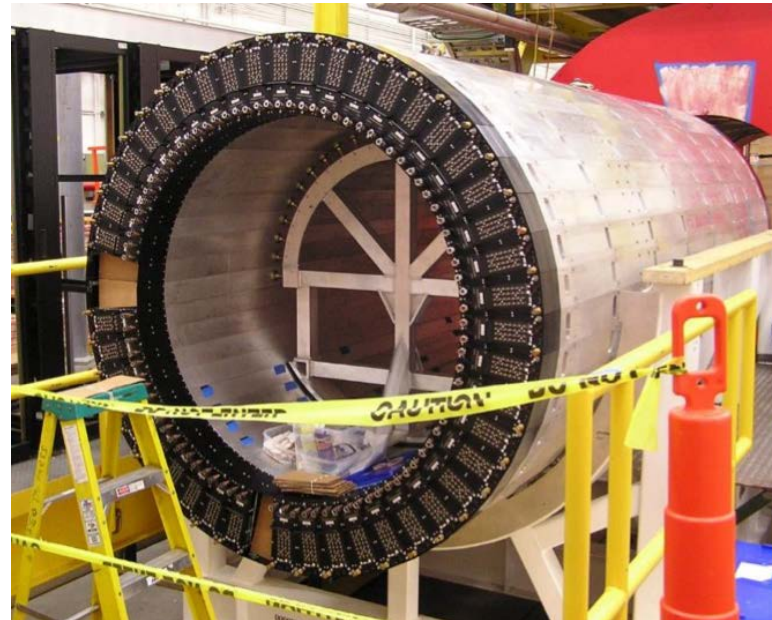
Hall D – First Large Scale Use of SiPMs at JLAB



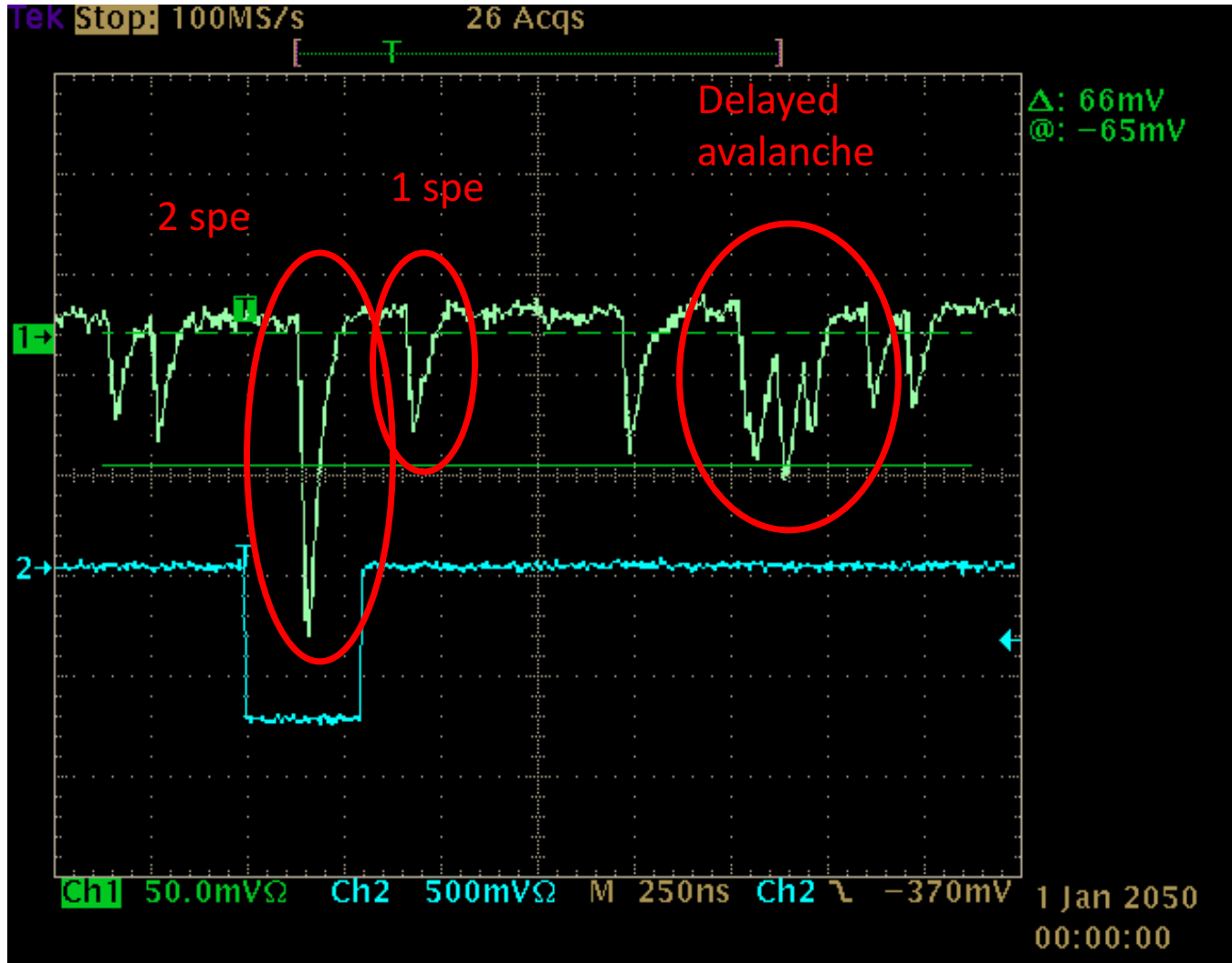
Total of 0.55 m²
of SiPMs



Barrel Calorimeter
3840 SiPM arrays
4x4 array of 3x3 mm² 50 μm
microcell (Hamamatsu)
PDE~20%, DCR=110 kHz/mm²



Noise in SiPMs: Dark current, crosstalk, afterpulses



- Noise derived from defects in lattice that create intermediate energy levels between valence and conduction band
- Radiation can increase number of such defects leading to increased noise - **dark current**
- Internal avalanche emits near-IR light - this can trigger avalanches in nearby pixels - **crosstalk**
- Also possible to have delayed avalanches - **afterpulses**
- These are also affected by radiation
- The suppression of this radiation induced noise is the key to future SiPM implementation

Particle interactions with SiPMs

X-rays (~ 10s keV)

Limited penetration depth leads to increased noise from surface interactions

High energy electrons and gammas (~ MeV)

Volumetric penetration - leads to defect buildup throughout SiPM via ionization processes

High energy neutrons (~MeV)

Principal source of defect buildup - neutron can displace atom from lattice site leading to interstitial vacancy - primary knock-on atom can continue to create other defects in material - essentially a non-ionizing source of noise buildup

Thermal neutrons (meV)

Nuclear transmutation ($^{30}\text{Si} + n \rightarrow ^{31}\text{Si} \rightarrow ^{31}\text{P} + \beta^-$)

High energy charged hadrons (i.e., protons)

Essentially combination of ionization and non-ionizing effects

Overall effects have to be judged by experimental conditions of background radiation sources

Gamma Irradiation - Co-60 - 1 MeV

R. Pagano et al. / Nuclear Instruments and Methods in Physics Research A 767 (2014) 347–352

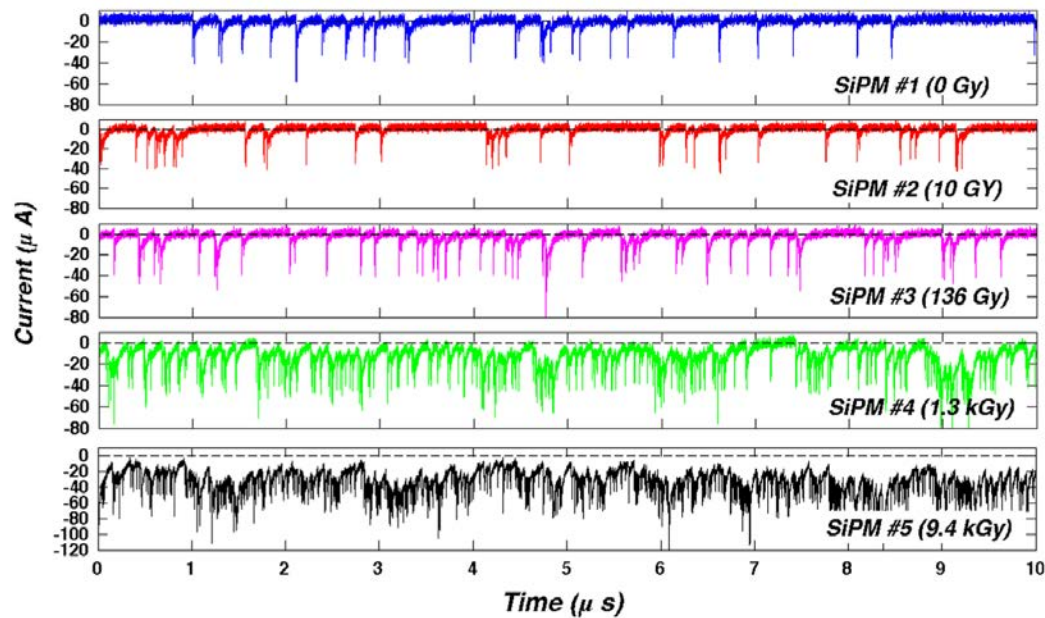


Fig. 2. Dark current traces at RT ($21^\circ \pm 3^\circ\text{C}$) for a $\Delta V=3\text{ V}$ for the SiPM before and after irradiation at different doses.

Pagano et al. NIM A 767 (2014) 347-352

R. Pagano et al. / Nuclear Instruments and Methods in Physics Research A 767 (2014) 347–352

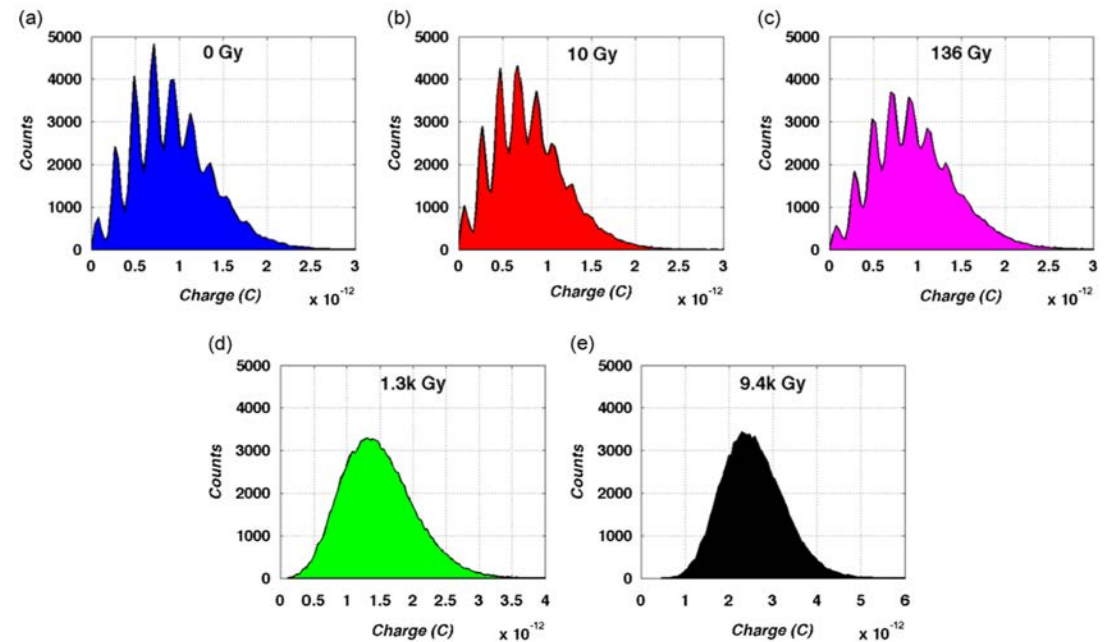
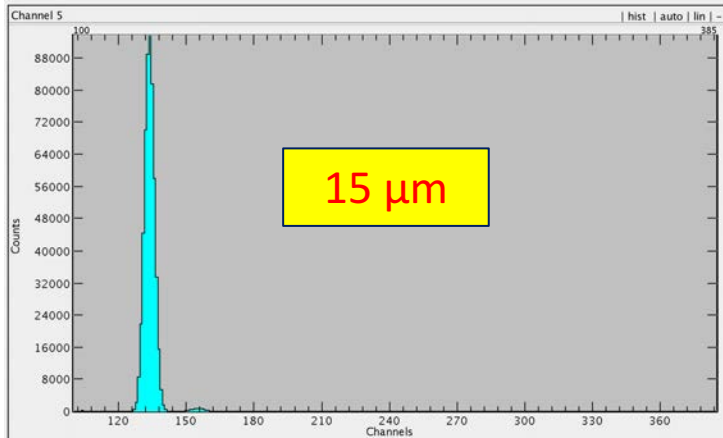


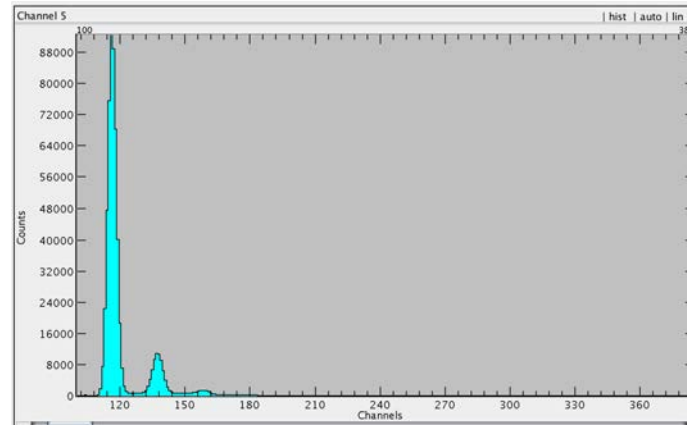
Fig. 5. SPS at $\Delta V=3\text{ V}$ and at RT ($21^\circ \pm 3^\circ\text{C}$) before and after γ -rays irradiation.

Annealed at room temperature

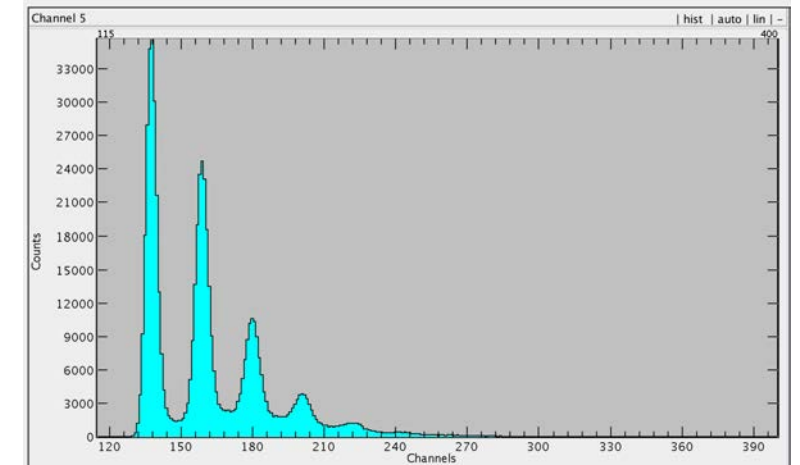
Dark Rate increase with irradiation by high energy neutrons (Am/Be source) (post-irradiation annealed at 60°C for 24 hrs)



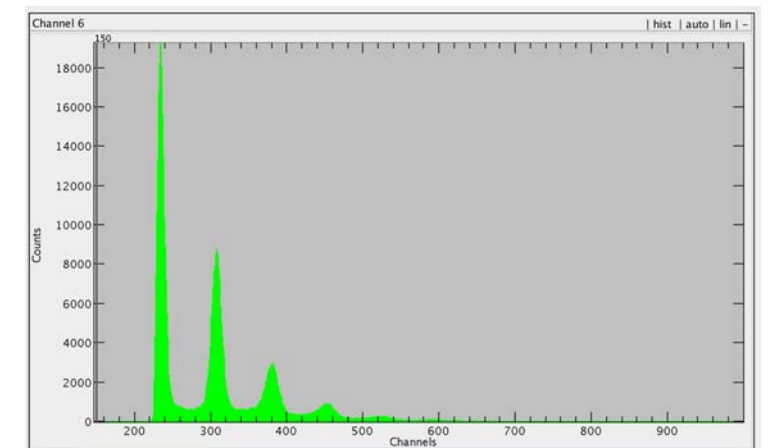
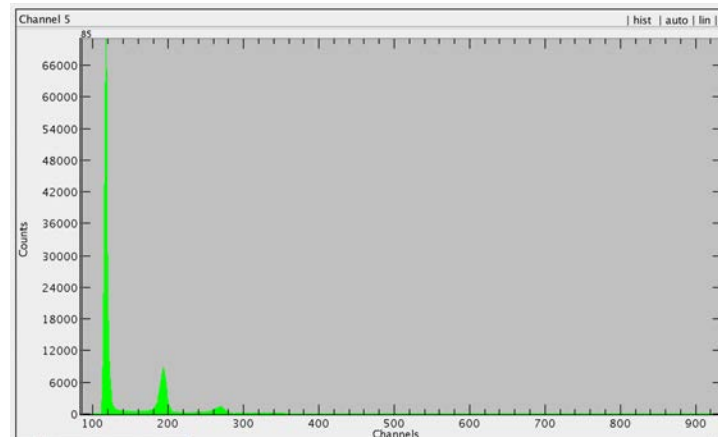
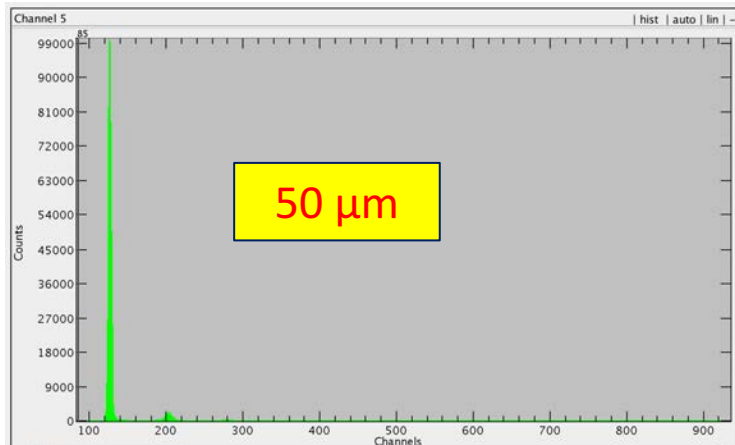
Pre-irradiation



10^9 n/cm^2

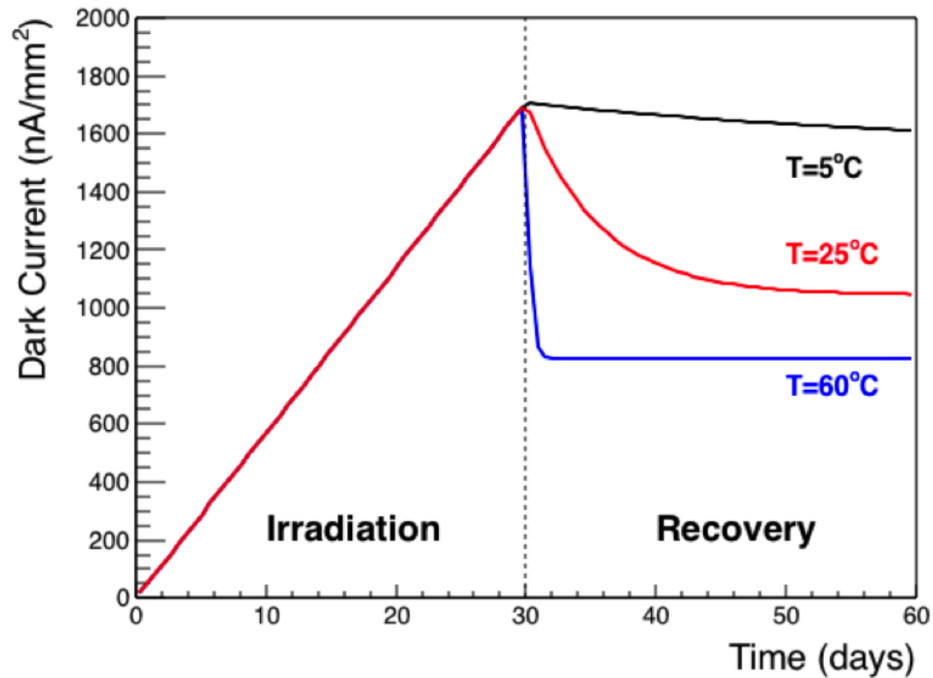


$5 \times 10^9 \text{ n/cm}^2$

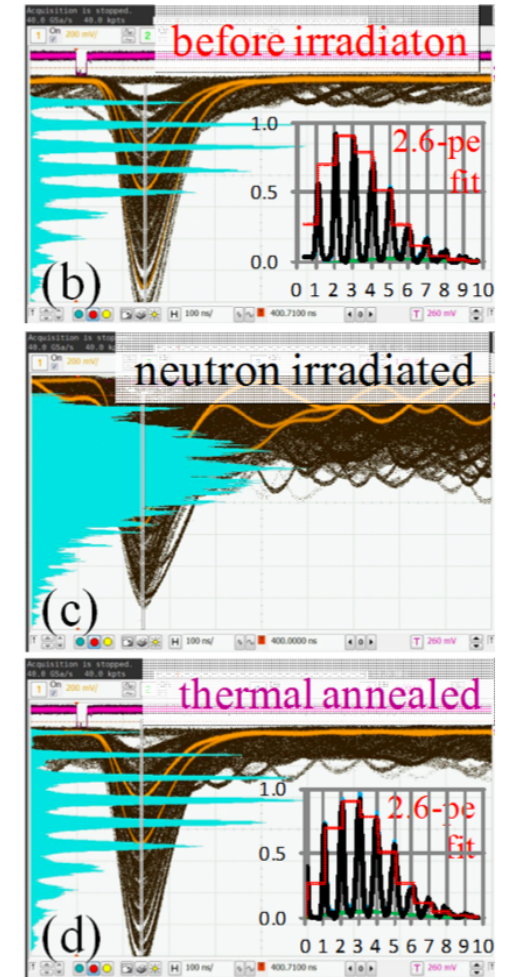
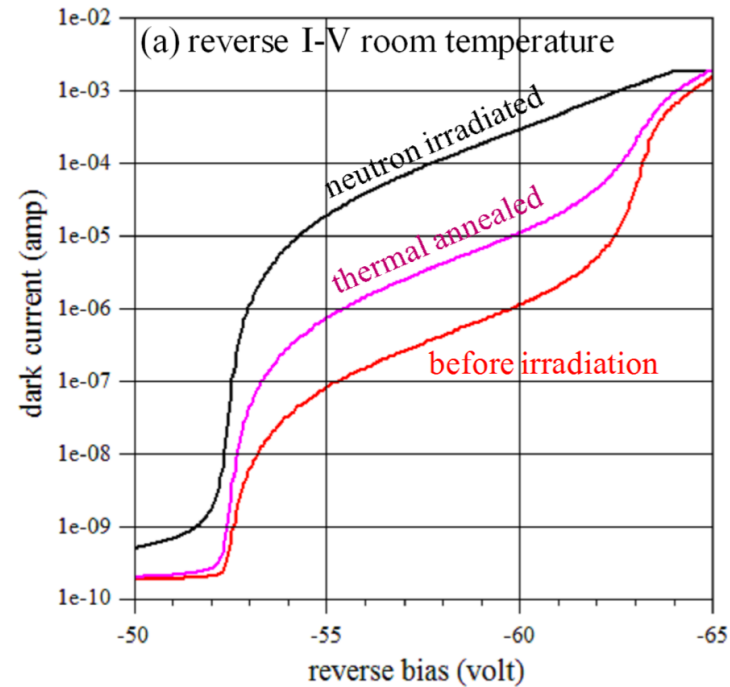


Radiation Damage - Annealing of Prompt effects

JLAB SiPM raddam study
Qiang et al., NIM A 698 (2013) 234-241

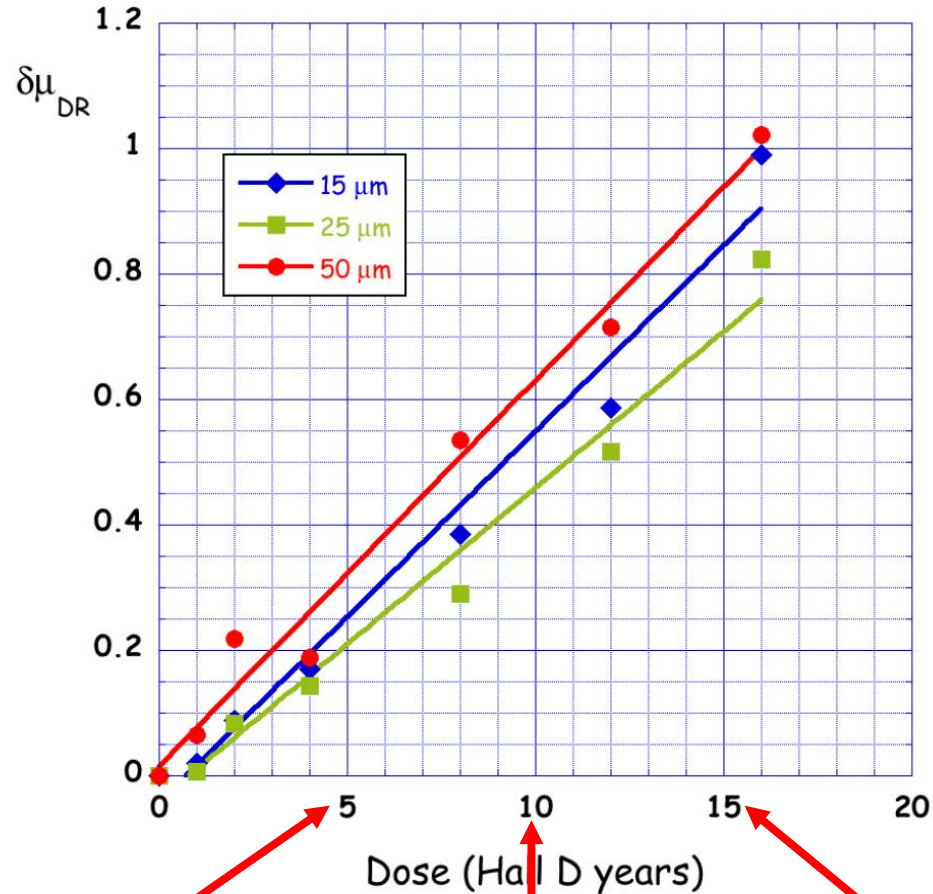


BNL study - 10^9 n/cm²
Thermal anneal - 250°C
+ 8 mA forward bias



Tsang et al., 2016 JINST 11 12002

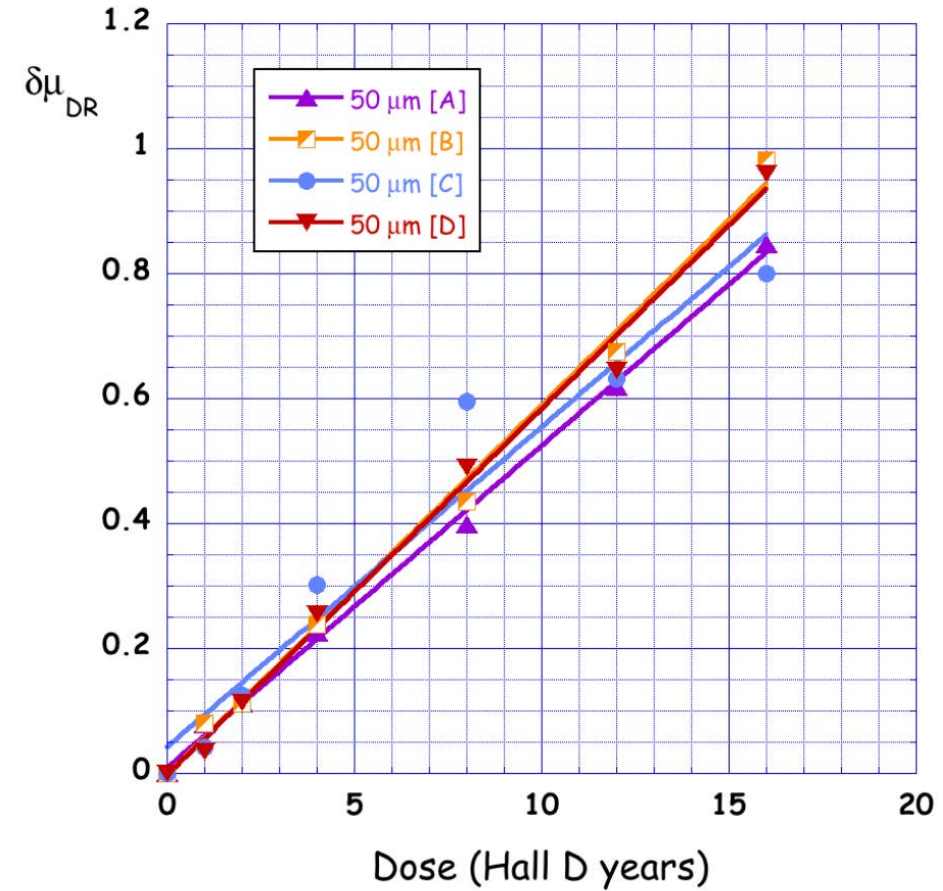
Dark Noise increase with dose - Hamamatsu SiPMs (MPPCs)



$1.5 \times 10^9 \text{ n/cm}^2$

$3 \times 10^9 \text{ n/cm}^2$

$4.5 \times 10^9 \text{ n/cm}^2$



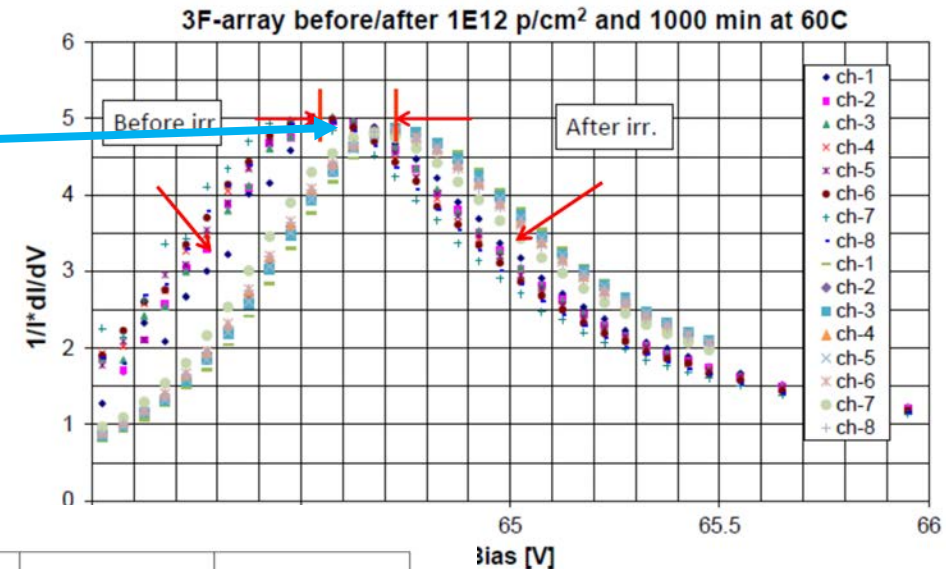
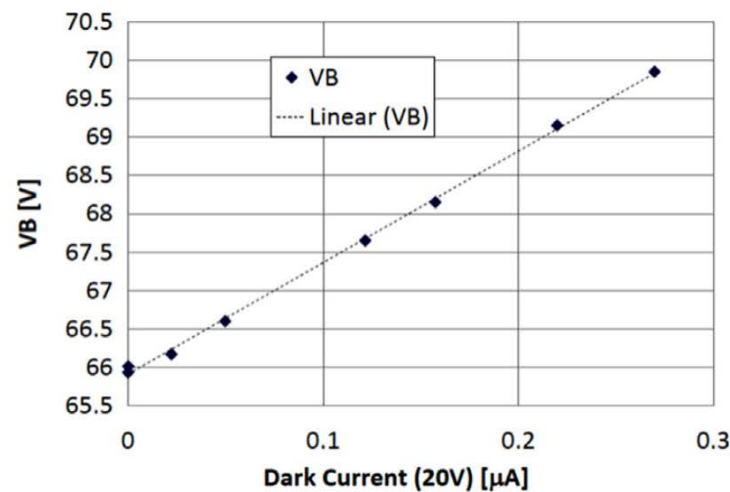
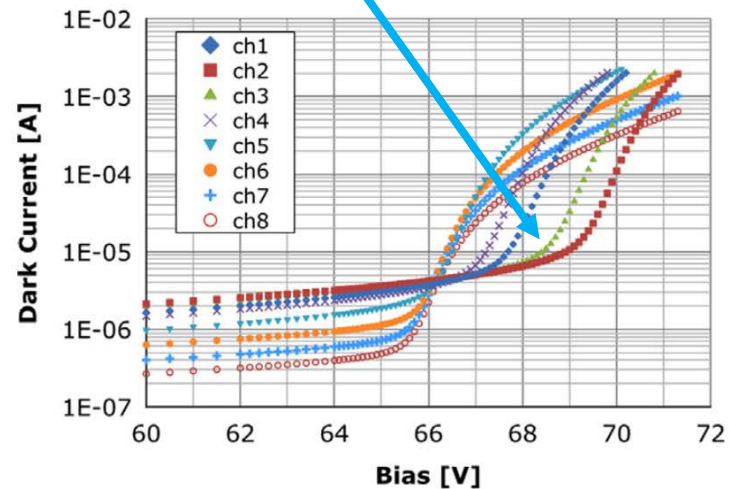
Irradiation of $3 \times 3 \text{ mm}^2$
SiPMs with Am/Be source

Barbosa et al., 2012 IEEE proceedings

Effects of Very High Doses - Shift in V_B - affects Gain and PDE

Maximum dose
 $2.2 \times 10^{14} \text{ n/cm}^2$

175 mV increase in
breakdown voltage
 $2 \times 10^{12} \text{ n/cm}^2$



Heering et al. NIM A 824 (2016) 111-114

Rundown of Effects of Radiation on SiPMs

Increased dark current

- leads to smearing of photoelectron resolution
- threshold 10^8 n/cm² or 1 kGy
- above 10^{11} n/cm², dark noise can be so high as to saturate device response
- longer recovery time of larger pixels is relevant parameter - smaller pixels recover faster

Change in Breakdown Voltage

- VB increases slightly - example: +175 mV for $> 10^{12}$ n/cm²
- threshold $> 10^{10}$ n/cm²
- creates a gain and PDE decrease - will need to shift the bias voltage

Thermal Neutron Contribution to Dark Noise

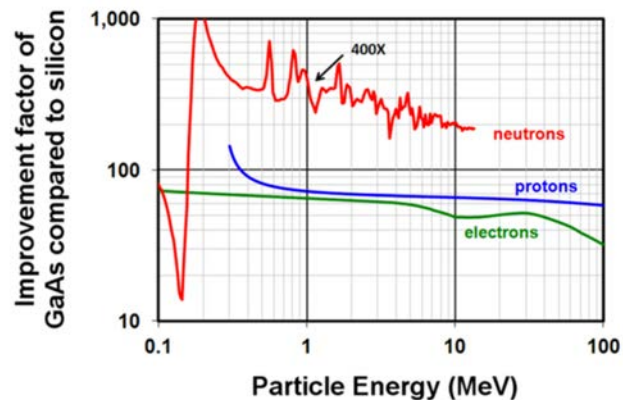
- seems to be significant only $> 10^{11}$ n/cm²

Transmission window of SiPM may yellow with irradiation

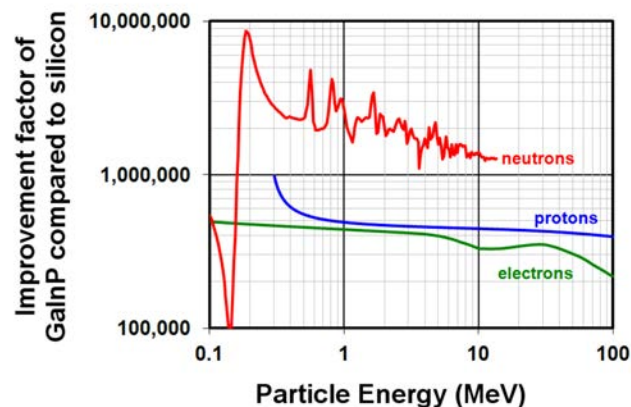
- neutrons have almost no effect
- > 10 kGy to see any effect from gammas

Farther down the Line - new tech - *LightSpin Technologies (NY)* - GaAs/GaInP "SiPM"

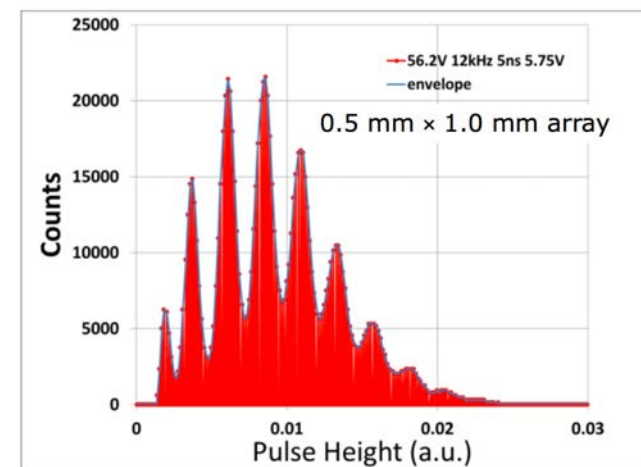
Radiation Hardness GaAs vs. silicon



Radiation Hardness GaInP vs. silicon

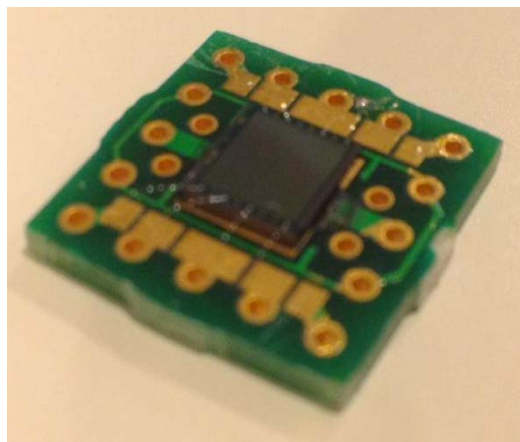
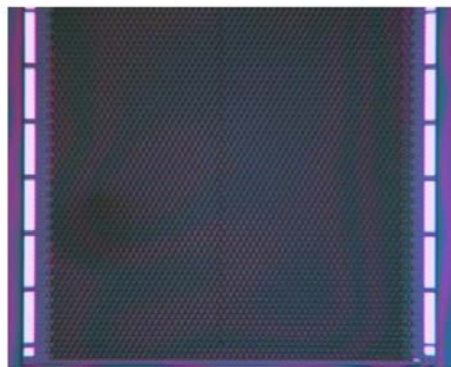


Photon Number resolving



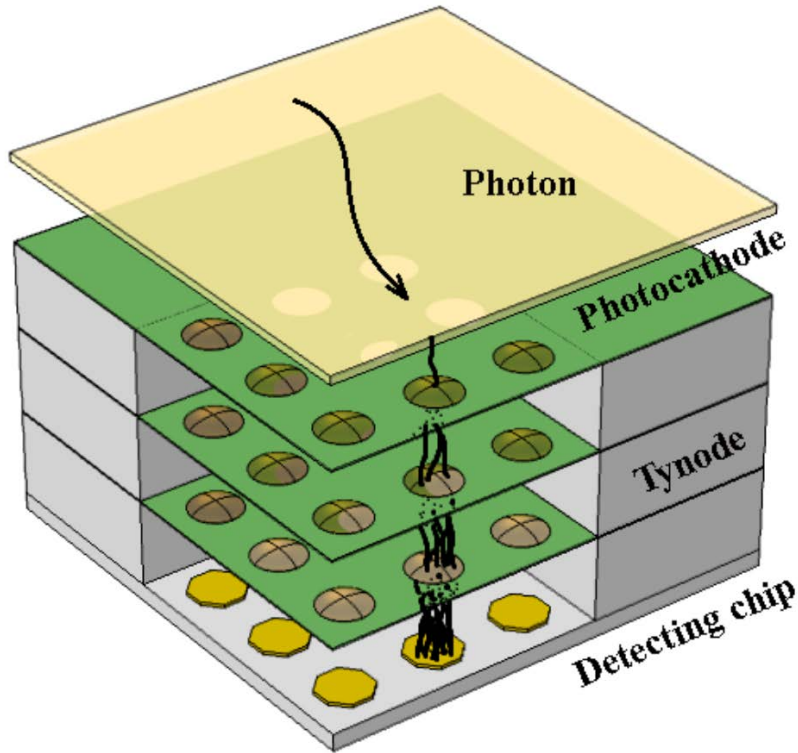
LightSpin Technologies, Inc.

3 mm × 3 mm die
← 3 mm →

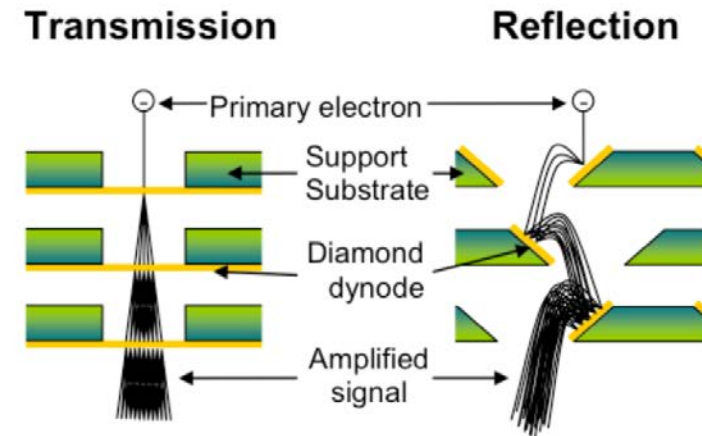


Another example of possible Future Tech - **TiPC** - Harry van der Graaf

Timed Photon Counter



Trynode – transmission dynode



PHOTONIS

TU Delft

BROOKHAVEN
NATIONAL LABORATORY

Argonne
NATIONAL LABORATORY

NIKHEF

- ➔ Expect very low noise compared to SiPM
- ➔ And very high timing resolutions (< 10 ps)

In Conclusion...

Present knowledge base is effort of many people dealing with their own experimental priorities

Time to pool knowledge and craft a R&D path with the goal of creating a certified community accepted knowledge base

Industry is interested in our feedback to them - raddam especially

Needed to successfully take advantage of this remarkable device