

## What Temperature is optimal for the NPS ?

1. Approved JLab 12 GeV energy experiments E12-13-007 (P<sub>T</sub> in SIDIS); E12-13-010 (DVCS); E12-14-003 (WACS); E12-14-005 (WA exclusive π<sup>0</sup>) and E12-14-006 (Polarization in WACS) require:

- Energy resolution  $\frac{\sigma}{E} \sim \frac{3\%}{\sqrt{E}}$
- Coordinate resolution  $\sigma_{x,y} \sim \frac{2-3}{\sqrt{E}} \text{ mm}$
- Energy range of detected particles (photons) from 300 MeV to 7.4 GeV

2. Let assume NPS will operate at room temperature (at 18°C). Then:

- Light yield is about 15 p.e. per 1 MeV energy deposition (R. Novotny)
- Slow component: ~10% (fraction of LY above 100 ns up to 4 μs)
- Temperature stability: 0.1°C (guarantees a 0.5% resolution stability)
- Spontaneous recovery of radiation damage: from few hours to few days

3. Effects from cooling PbWO<sub>4</sub> from T=+18°C down to T=+14°C or 10°C:

- Light Yield per 1 MeV energy deposition will increase and from 15 p.e. to 16.5 p.e. at T=+14°C and 18.5 p.e. at T=10°C (or will increase by ~11% and ~23%). Note: from PANDA group studies LY(-25)/LY(+18)~3.5-4.0 or increased by 3% per 1°C (assuming exponential growth of LY with temperature:  $LY(T) \approx LY(0^\circ) \times e^{0.03 \times (0^\circ - T)} \approx 25 \times e^{0.03 \times (0^\circ - T)}$ ).
- Slow component also will increase and will reach 12% at T=14°C and 14% at T=10°C. Note: PANDA group studies show that slow component will increase by ~5% per 1°C.
- Temperature stability: 0.1°C (guarantees a 0.5% resolution stability)
- Speed of spontaneous recovery of radiation damage will go down
- Technical difficulties and problems will arise

4. Well known PbWO<sub>4</sub> based calorimeter resolutions:

The energy resolution of a calorimeter can be expressed as

$$\frac{\sigma_E}{E} = \frac{a}{\sqrt{E}} \oplus \frac{b}{E} \oplus c, \text{ where } a/\sqrt{E} \text{ is stochastic term related to intrinsic}$$

shower fluctuations and photo-electron statistics,  $b/E$  is electronic noise term, and  $c$  is constant term that includes instrumental effects and mis-calibration.

- PANDA:  $\frac{\sigma}{E} = \frac{1.74}{\sqrt{E}} + 0.7$  (at  $T=+10^\circ\text{C}$ ), or 2.45% at 1 GeV.

- PANDA:  $\frac{\sigma}{E} = \frac{0.95}{\sqrt{E}} + 0.91$  (at  $T=-25^\circ\text{C}$ ), or 1.86% at 1 GeV.

(Note, from  $T=+10^\circ$  to  $T=-25^\circ\text{C}$  the LY increased by factor of  $\sim 3.5$  while resolution improved only by 30%, from 2.45% to 1.86%).

- HYCAL:  $\frac{\sigma}{\sqrt{E}} = \frac{1.17}{\sqrt{E}} \oplus \frac{1.80}{E} \oplus 1.15$  (at  $T=+14\pm 0.1^\circ\text{C}$ ) or 2.45% at 1 GeV.

- CMS:  $\frac{\sigma}{E} = \frac{2.8}{\sqrt{E}} \oplus \frac{0.12}{E} \oplus 0.30$  (at  $T=+18\pm 0.05^\circ\text{C}$ ) or 2.82% at 1 GeV.

5. NPS resolution at energy 300 MeV and  $T=+18^\circ\text{C}$ , and  $T=+14^\circ\text{C}$  or  $10^\circ\text{C}$ :

- At  $T=18^\circ\text{C}$  Light Yield is  $\sim 15$  p.e. per 1 MeV energy deposition, so photons with energy 300 MeV will generate  $\sim 4500$  p.e.
- At  $T=+14^\circ\text{C}$  photons of same energy (300 MeV) will generate  $\sim 5000$  p.e. ( $\sim 16.5$  p.e. per 1 MeV). At  $T=+10^\circ\text{C}$  same energy (300 MeV) photons will generate  $\sim 5500$  p.e. ( $\sim 18.5$  p.e. per 1 MeV).
- These numbers and CMS/HYCAL/PANDA data show that even at 300 MeV stochastic term “ $a/\sqrt{E} \approx 1/\sqrt{n}$ ” (n is number of photons) is not a limiting factor for the calorimeter’s resolution. The electronic noise (“b”-term), and instrumental effects and calibration (“c”-term) are essential.
- **We think stabilizing the NPS calorimeter at an ambient temperature ( $T=18\pm 0.5^\circ\text{C}$ ) would be quite acceptable for the NPS experiments. This will allow avoid the water condensation problem and related technical complications.**