



Backgrounds at KLF

Mikhail Bashkanov



- Small x-section
- Neutrons:
 - $-v_n \ll v_{K_L}$
 - Different kinematics

94% of neutrons associated with T < 300 MeV

2 2



Neutrons



- $E_n > 1.6 \ GeV$ (strangeness threshold)~ 1% of neutron flux
- $0.3 < E_n < 1.6 \text{ GeV}$ (above pion threshold)~ 5% of neutron flux
- $E_n < 0.3 \text{ GeV} \sim 94\%$ of neutron flux do not contribute

Neutron Cross-Sections





Neutron Reaction yield





Neutron-induced reaction rate: 233 ev/s

Neutron Background





• Neutron-induced reactions are not an issue for the main program

Useful Neutrons: Calibration





- Neutron-induced reactions are isospin I=1 dominated
- Can be used for calibration $np \rightarrow np\pi^0$
- ~ $24M \gamma' s$ from neutron-induced π^0 production per day
- Neutron-induced reactions have high scientific interest!



Photons

10

Photon background

Photoproduction cross section σ, mb 0 t $\gamma p \rightarrow p\pi$ 10⁸ →Dπ Reaction Yield $(\gamma/(sec cm^2 MeV))$ 10⁶ threshold 10⁴ 10² 10⁰ 10 10⁻² 10 1 10⁰ 10³ E,, GeV 10^{-1} 10⁴

Photon flux at LH2/LD2

Photoinduced reaction rate

Photoinduced reaction rate < 4Hz



Total

Photon background





Photon flux at LH2/LD2

Photoinduced reaction rate

- Photoinduced reaction rate < 4Hz
- Photoproduction reactions are not an issue for the main program



Cosmics

Cosmic muon background







- Cosmic rate induced rate: $\sim 500 \ ev/s$
- Cosmic rays are not an issue for the main program

Total budget



Reactions	Rate [kHz]
<i>K_L</i> –induced	1.0
n —induced	0.4
γ —induced	0.004
cosmics	0.5
Total	~2.0

Conclusion



- Background conditions at KLF are very mild
- No physical background problems at reconstruction level. (from neutron- gamma- induced production in cryogenic target)







